

https://cds.cern.ch/record/1742528



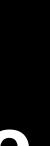
Introduction to CERN and High Energy Particle **Physics Programs**

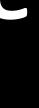
N. SRIMANOBHAS, Chulalongkorn U. (TH), CMS Collaboration 24 Dec 2022, OBEC

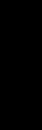




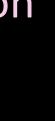






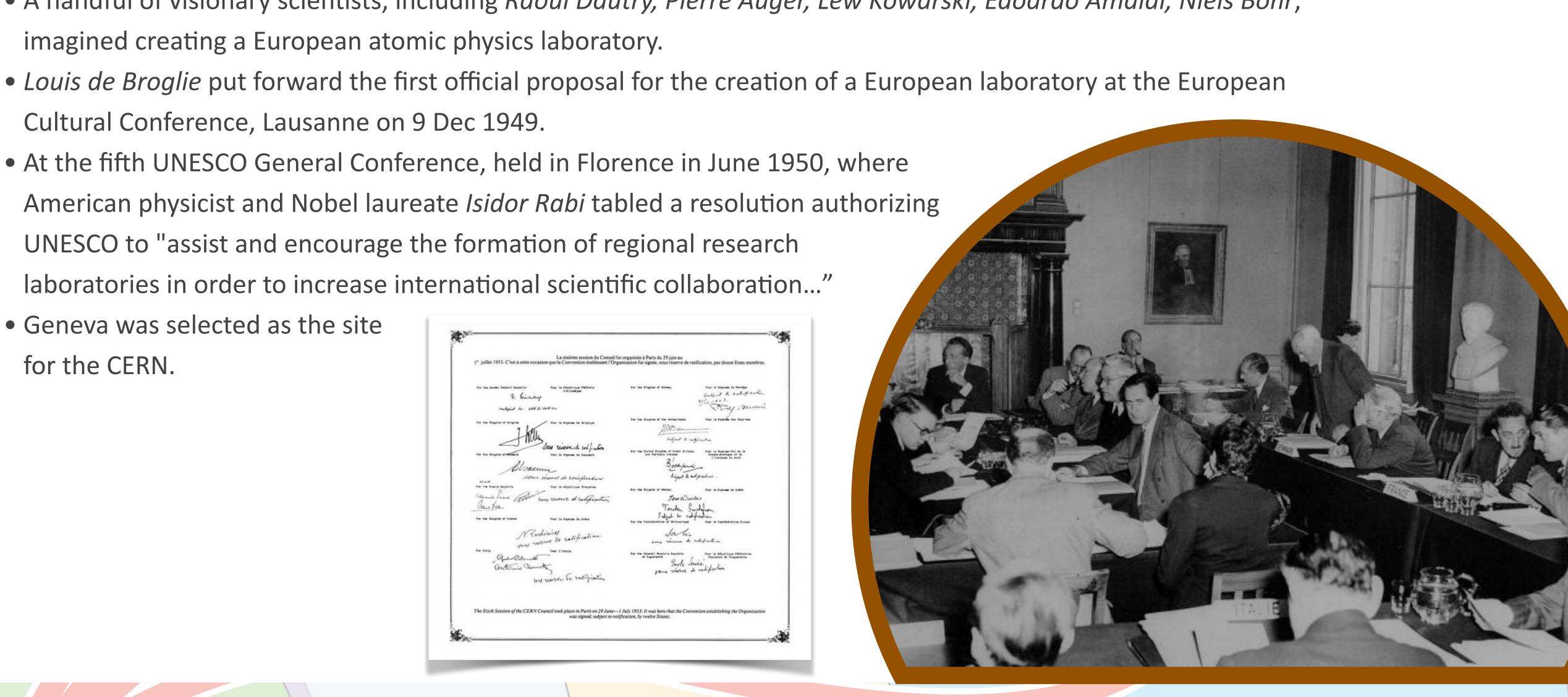






CERN: Conseil Européen pour la Recherche Nucléaire

- At the end of the Second World War, European science was no longer world-class.
- A handful of visionary scientists, including Raoul Dautry, Pierre Auger, Lew Kowarski, Edoardo Amaldi, Niels Bohr, imagined creating a European atomic physics laboratory.
- Cultural Conference, Lausanne on 9 Dec 1949.
- At the fifth UNESCO General Conference, held in Florence in June 1950, where American physicist and Nobel laureate *Isidor Rabi* tabled a resolution authorizing UNESCO to "assist and encourage the formation of regional research laboratories in order to increase international scientific collaboration..."
- Geneva was selected as the site for the CERN.



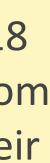
CERN: Conseil Européen pour la Recherche Nucléaire

MEMBER STATES ASSOCIATE MEMBER STATES **CIATE MEMBERS IN** THE PRE-STAGE TO MEMBERSHIP OBSERVERS **OTHER STATES**

- Today 23 member states
- 10 associated member states (3 in pre-stage to member states)
- 1 to become associated member state
- Thailand has an international Cooperation Agreement with CERN since 2018 CERN employs just over 2500 people and some 12,000 visiting scientists from over 70 countries and with 120 different nationalities come to CERN for their
- research

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• Found in 1954 with 12 European member states





CERN: Conseil Européen pour la Recherche Nucléaire

Distribution of All CERN Users by Location of Institute on 27 January 2020

MEMBER STATESAustria86Belgium145Bulgaria40Czech Republic246Denmark43Finland85France832Germany1 258Greece147Hungary74Israel71Italy1 498Netherlands180Norway86Poland298Portugal88Romania115Serbia38Slovakia75Spain350Sweden100Switzerland364United Kingdom944									
United Kingdom 944 7 163	THE PRE-STAGE								
ASSOCIATE MEMBERS Croatia 41 India 186 Lithuania 21 Pakistan 39 Turkey 128 Ukraine 35 450	34 OBSERVERS Japan 245 Russia 1 071 USA 1 960 3 276	OTHERS Algeria Argentina Armenia Australia Azerbaijan Bahrain Belarus Brazil	Canada Chile 3 China 16 Colombia 13 Cuba 23 Ecuador 2 Egypt 3 Estonia 27 Georgia 114 Hong Kong	206Iceland22Indones362Iran21Ireland3Jordan4Korea16Kuwait24Latvia37Lebanor21Malaysi	$ \begin{array}{r} 11\\ 7\\ 1\\ 143\\ 2\\ 2\\ n\\ 15\\ \end{array} $	Malta Mexico Mongolia Montenegro Morocco New Zealand Oman Peru Puerto Rico Singapore	53 S 2 T 5 T	South Africa Sri Lanka Taiwan Thailand J.A.E. 13	80 8 55 18 2







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CERN Missions

Research

Seeking and finding answers to questions about the Universe

Technology Advancing the frontiers of technology

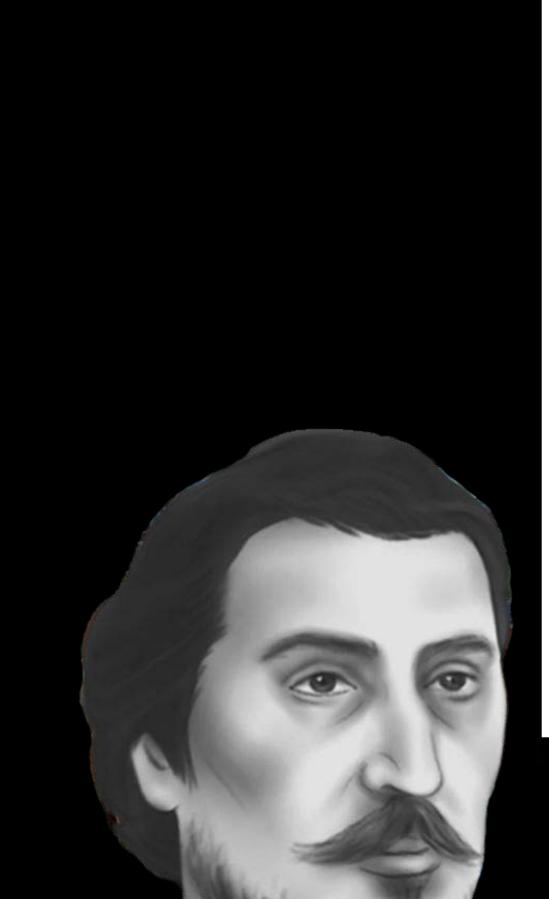
Education Training the scientists of tomorrow

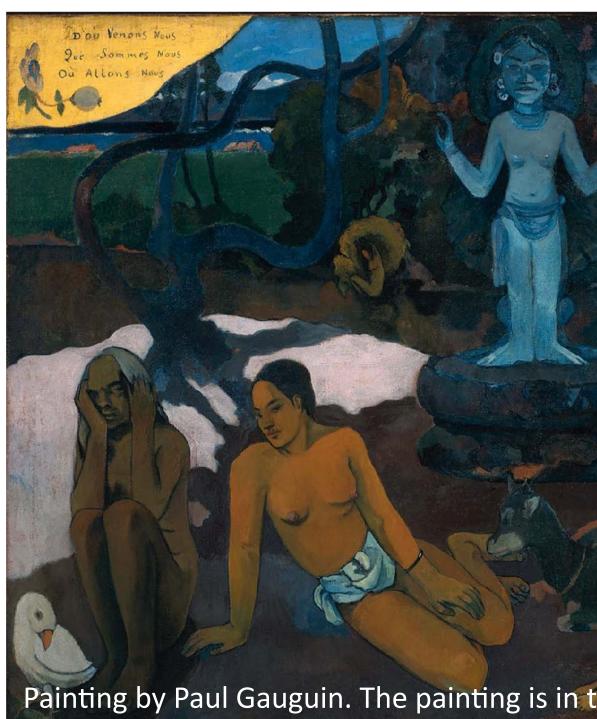
Collaborating Bringing nations together through science





Where do we come from? What are we? Where are we going?





Where Do We Come From? What Are We? Where Are We Going? Note that this painting should be read from right to left.

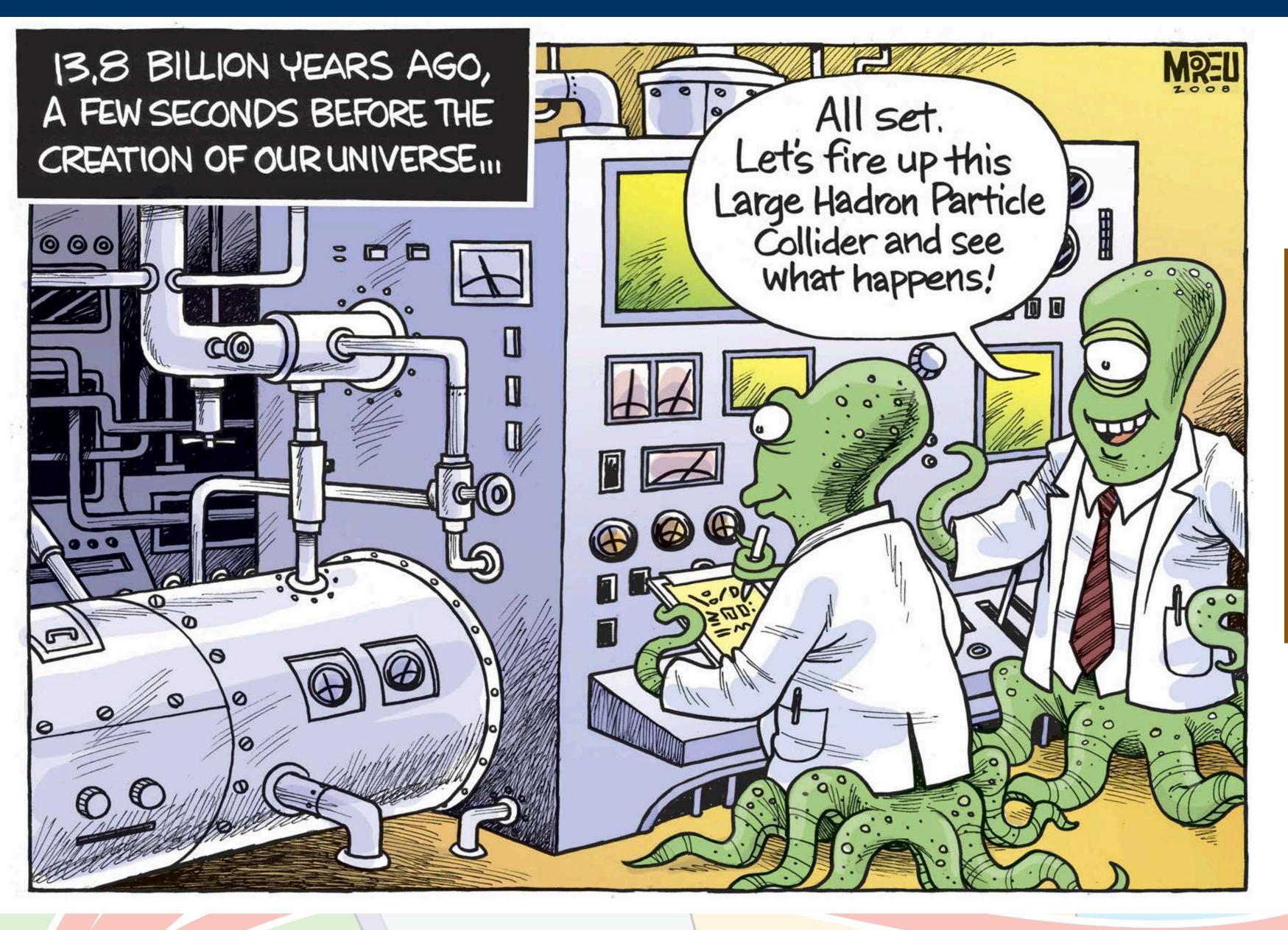
Paul Gauguin 1848 - 1903

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Painting by Paul Gauguin. The painting is in the Museum of Fine Arts in Boston, Massachusetts, US.



What are we looking for?

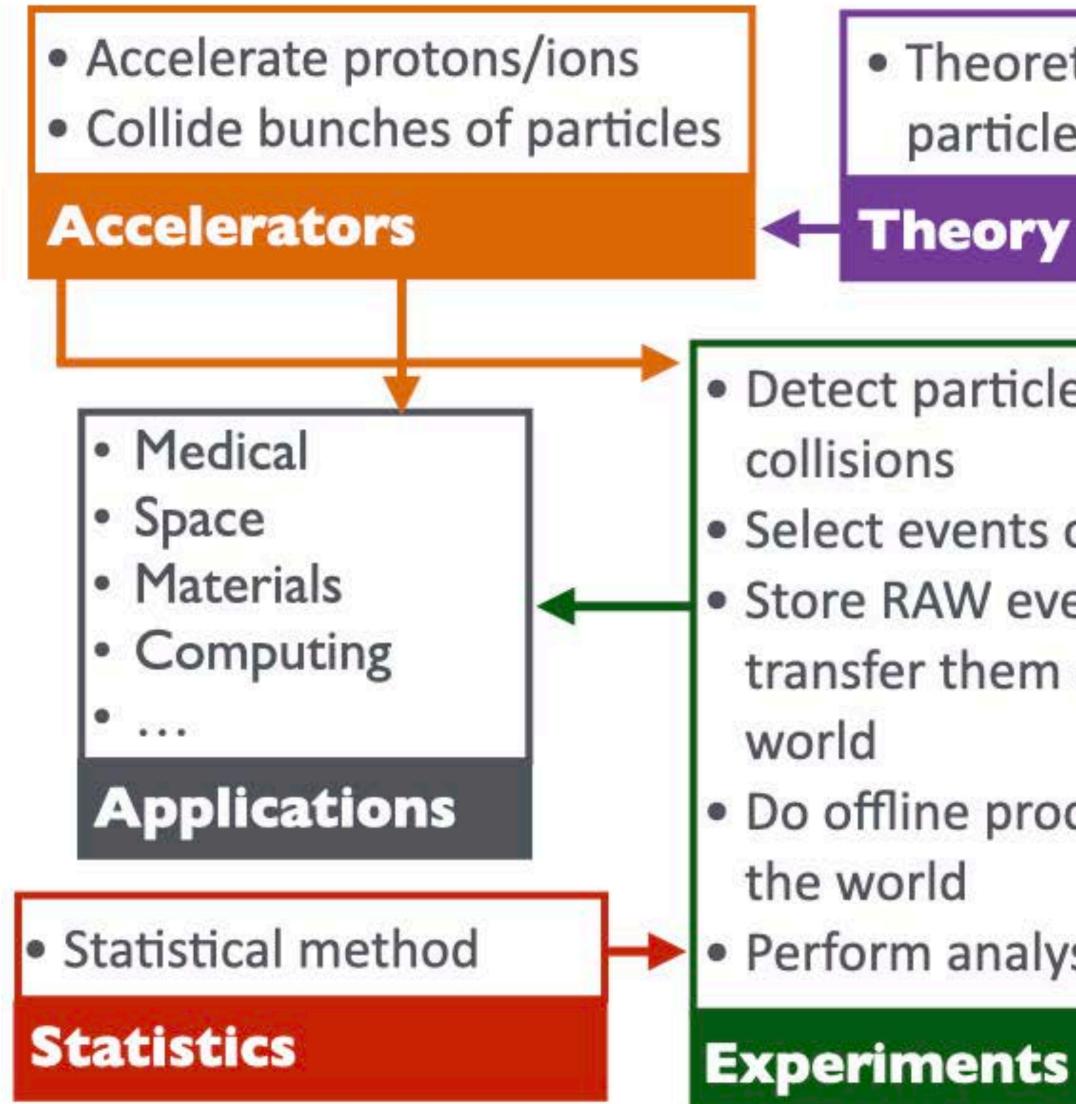


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Basic research in the field of experimental and theoretical particle physics, finding out what the Universe is made of and how it works.



What do you learn in this workshop?



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 Theoretical particle physics

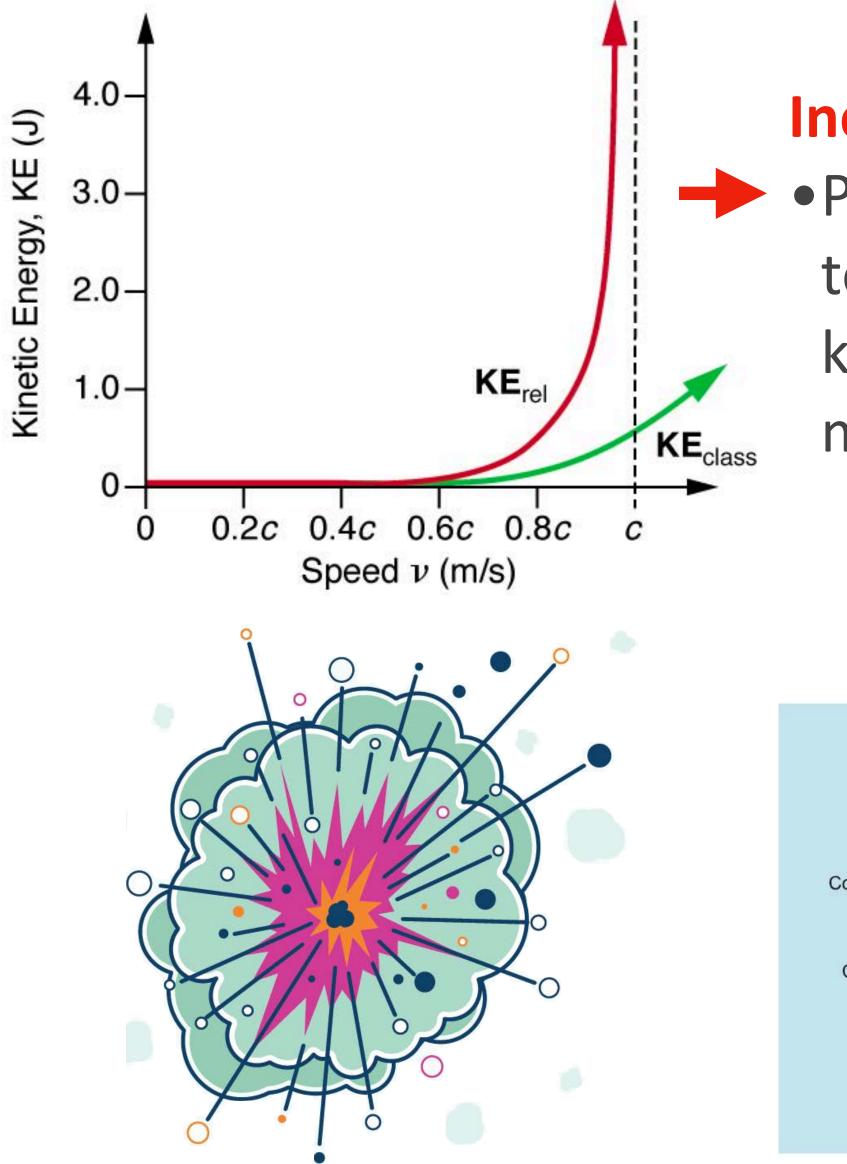
Theory

- Detect particles coming from
- Select events of interest
- Store RAW events, and
 - transfer them around the
- Do offline processing around
 - Perform analyses

 Advanced maths Mathematics Engineering Programming Machine learning Computing

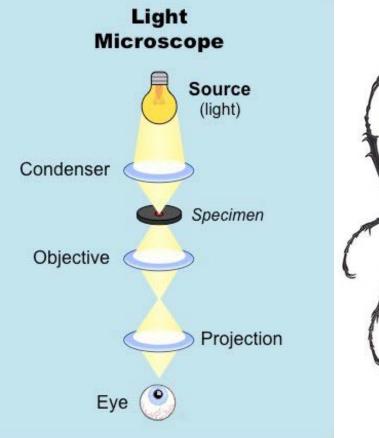


Accelerator is our tool to find answers



Increase K.E. of particle Particle speed has limit to the speed of light but kinetic energy and momentum have not

Visible light 400-500 nm



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Use high momentum particle to

create matter by collision $E^2 = p^2 c^2 + m^2 c^4$

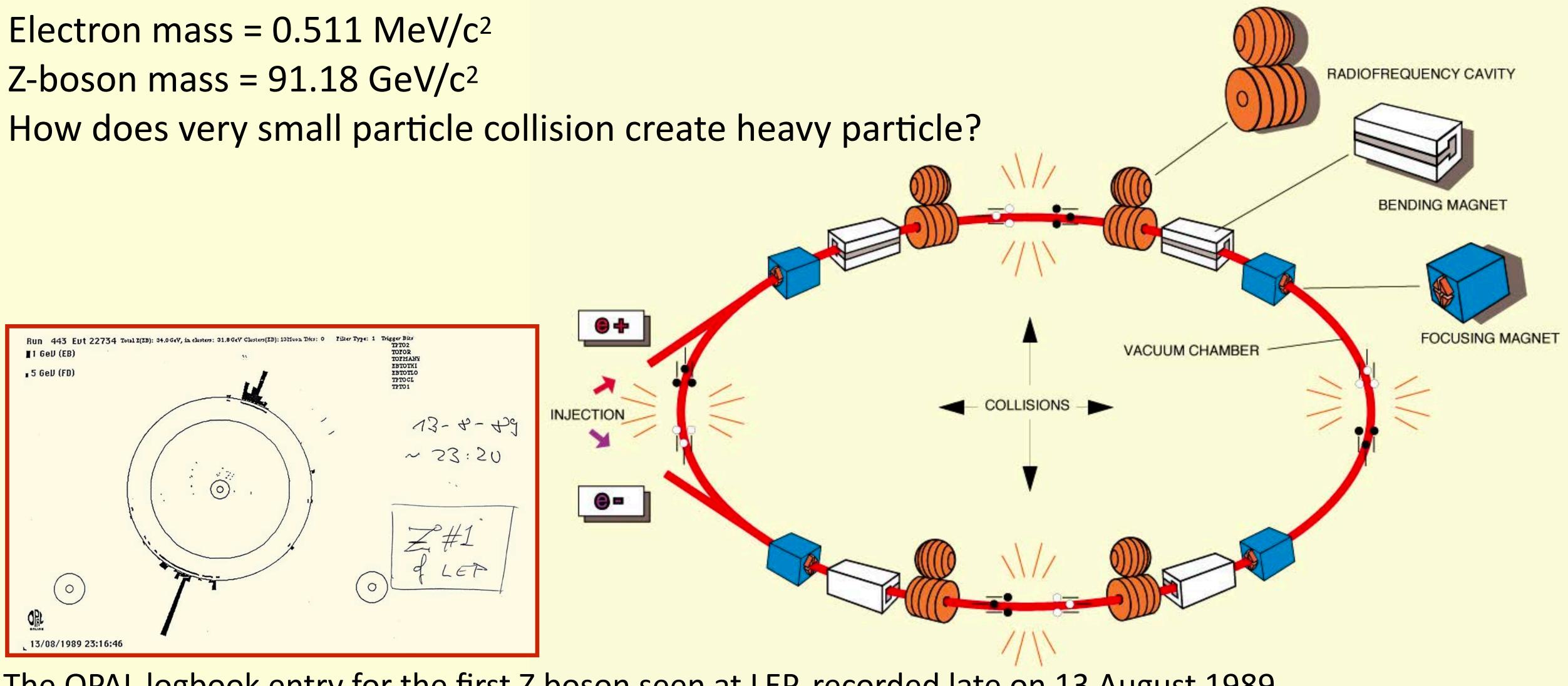
 probe insight in particle structure $\lambda = h/p$

X-Ray **Particle Accelerator** 0.01-10 nm <0.01 nm Dew Drop Electron (1/25th of Microscope Source Molecule 3 x 10⁻⁷ mm electrons) Hydrogen 0.0000001 mn or 10^{.7} mm Condenser 💿 🔵 Eye Objective Quarks and Electron Specimen detector





Accelerator is our tool to find answers



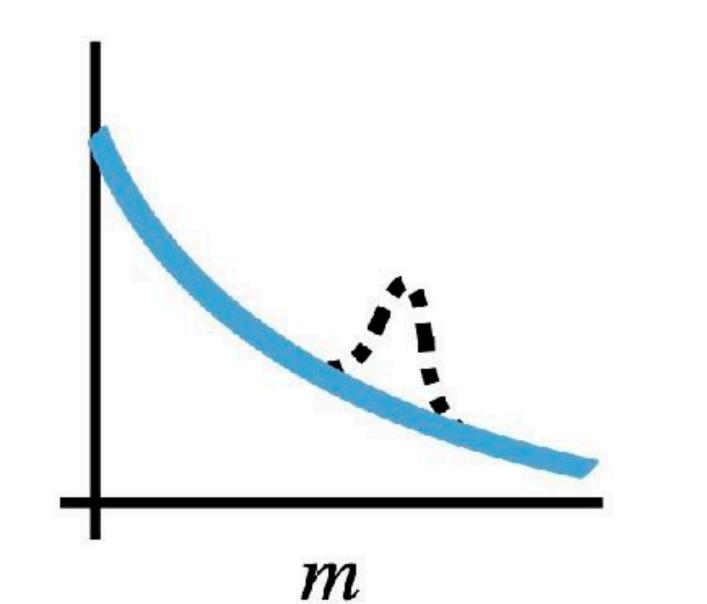
The OPAL logbook entry for the first Z boson seen at LEP, recorded late on 13 August 1989. Credit: CERN [Link]

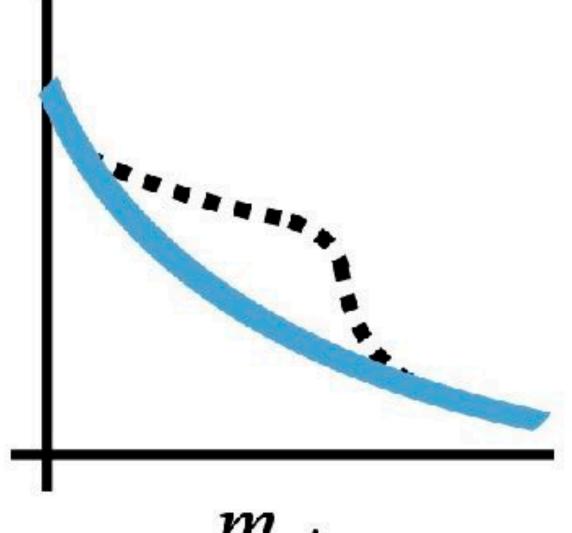
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CERN AC - E509



Searching for new physics





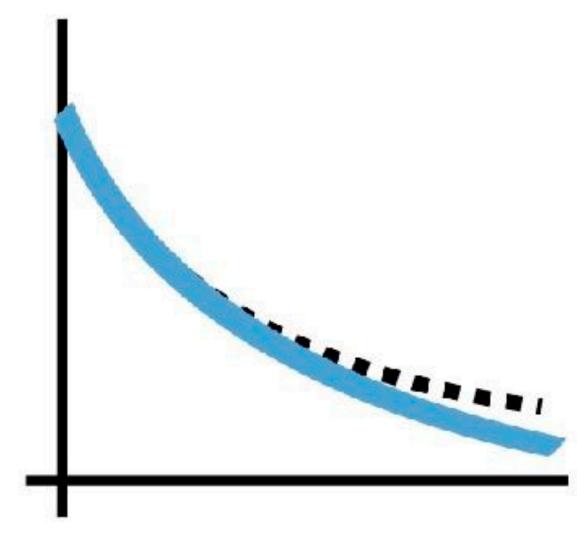
Bump(s), Resonant Di-Higgs interpretation, Z', ...

Semi-resonant, end-point Heavy charged bosons,

. . .

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 $m_{\rm vis}$



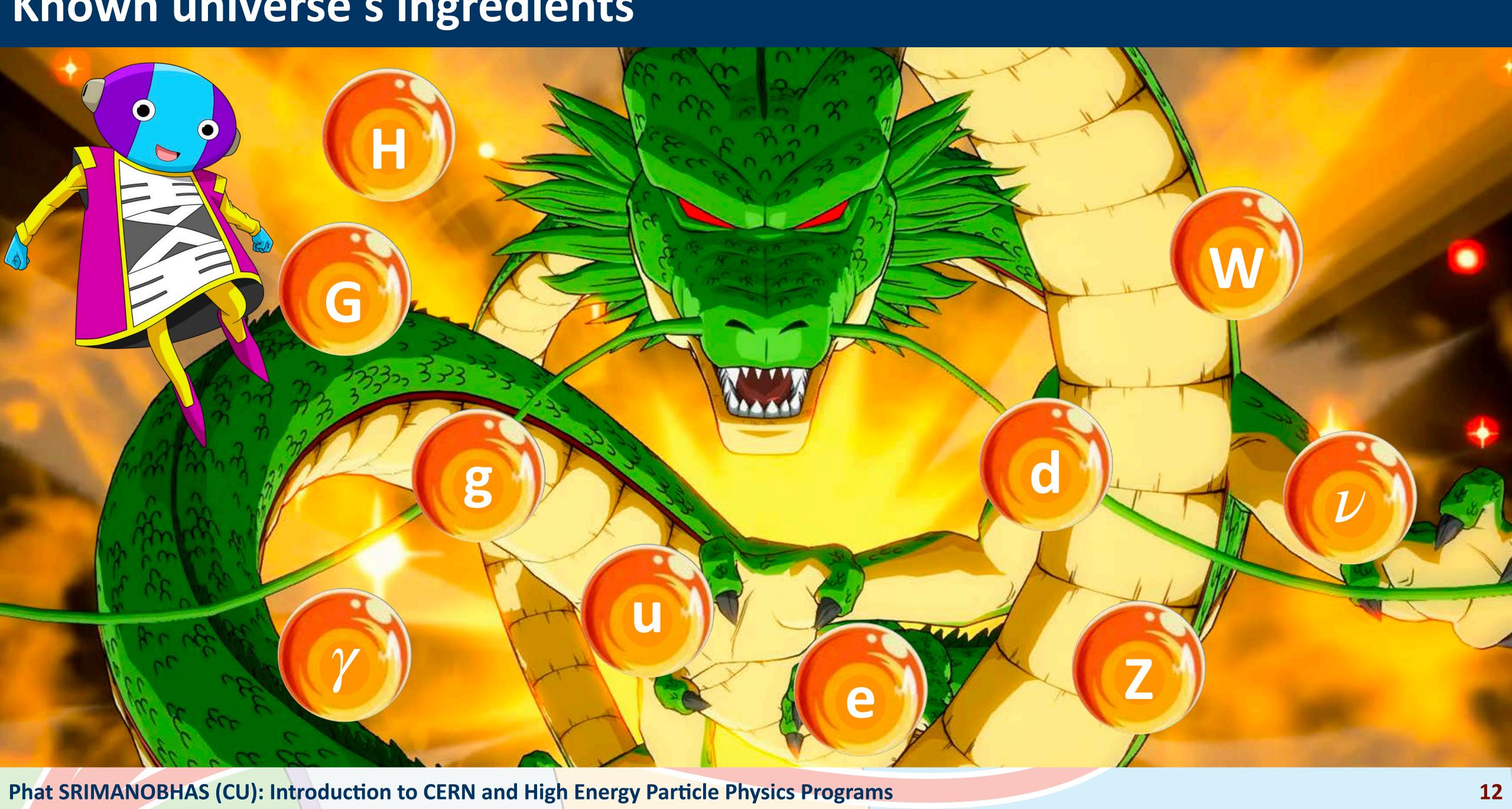
 p_T

Non-resonant, Tail *Dark matter, Extra dimensions,...*

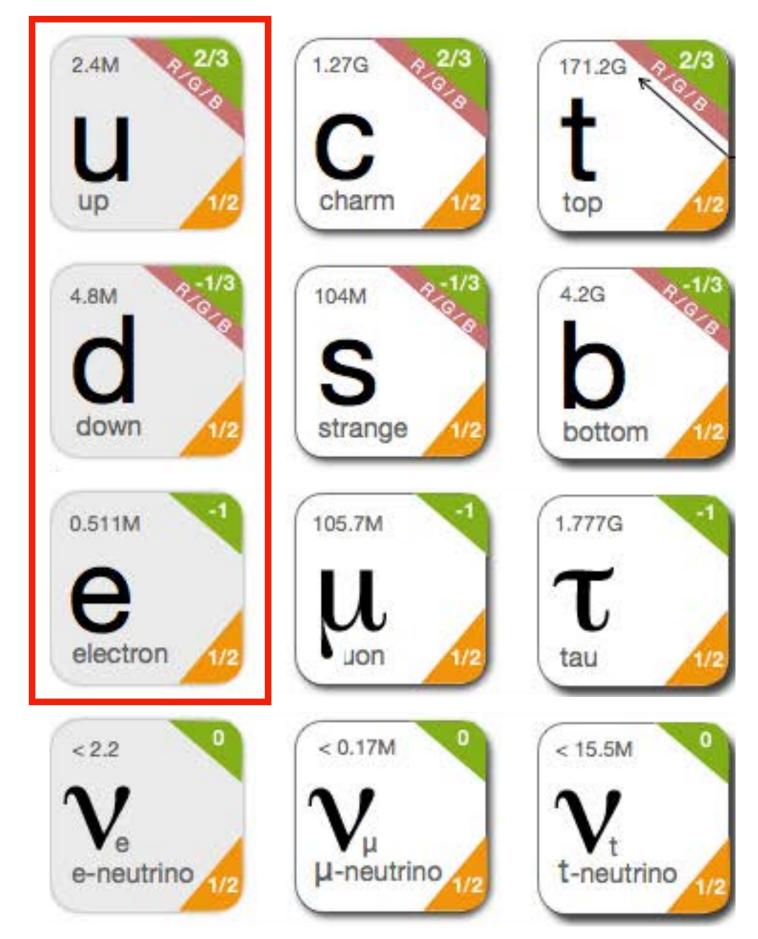


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Known universe's ingredients



Known universe's ingredients



Everything around us (the whole periodic table) is made up of the first three particles (u-quark, d-quark and electron). But somehow nature • supplies us with two extra families that are very much heavier, doesn't allow us to see free quarks (or maybe we don't see it yet), • group of 3 quarks: **Baryon** (e.g. proton), • quark-antiquark pairs: Meson (e.g. pion), describes interactions between



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these particles by three (out of four) fundamental forces.

This is so simple, compared to Periodic Table in Chemistry, or **Biological taxonomy!**





Electromagnetic force



Forming atoms and molecules

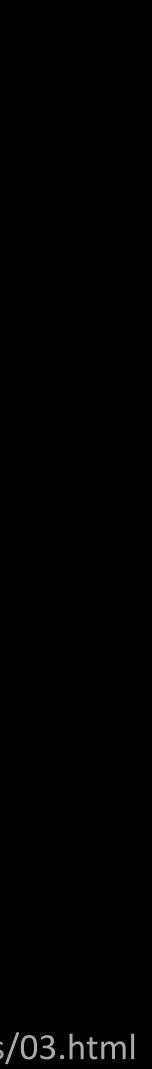
The electromagnetic force pulls negatively charged electrons into bound orbits around positively charged nuclei to form atoms and molecules. As a gas cools, electrons will find their way into the presence of atomic nuclei. Larger nuclei with a greater positve charge pull in more electrons until atoms and molecules have a balance of charges.

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Generating light

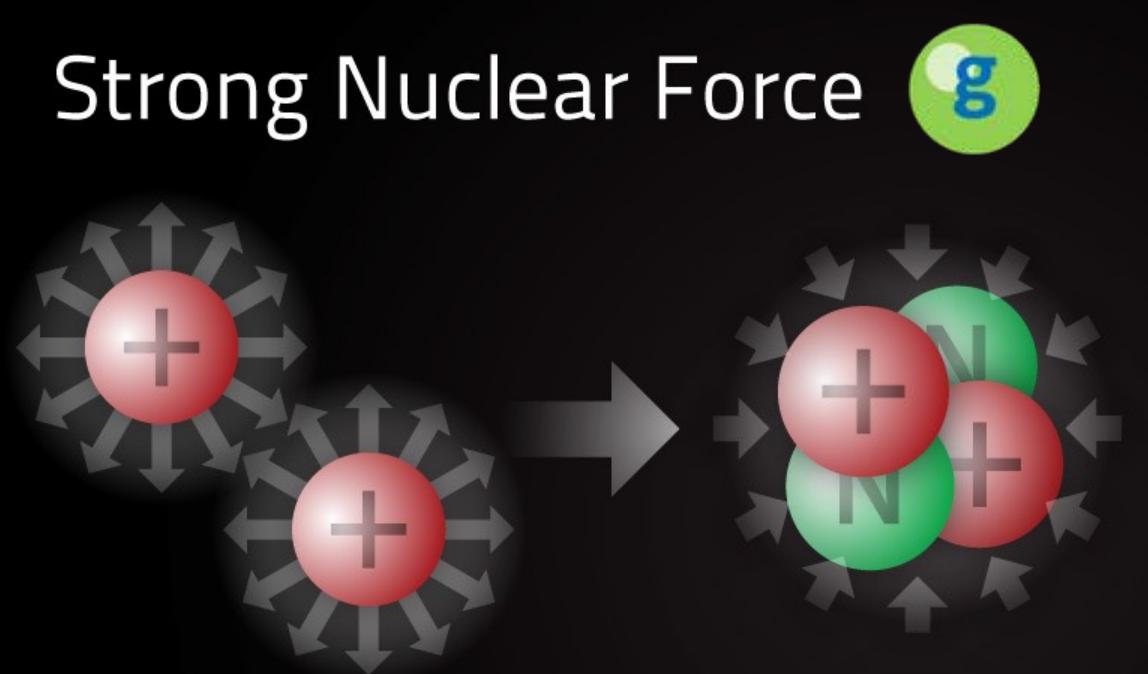
When a negative electron interacts with a positive proton, the electromagnetic force adds energy to the electron generating a photon.

http://ecuip.lib.uchicago.edu/multiwavelength-astronomy/astrophysics/03.html





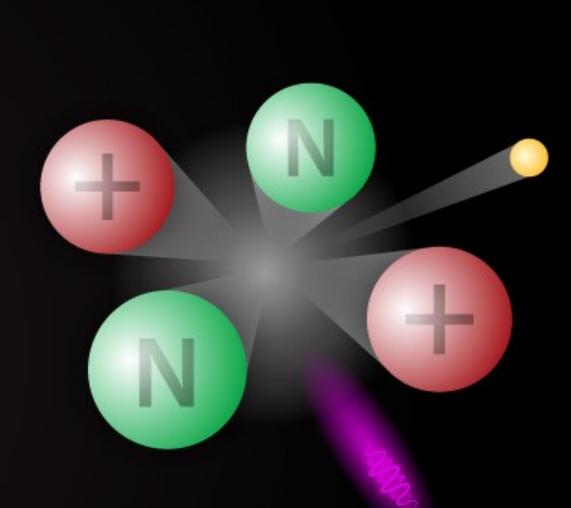
Strong nuclear force



Binding protons in atomic nuclei

Positively charged particles naturally repel each other, it takes an extreme amount of force to hold protons together. The strong nuclear force overcomes the repulsion between protons to hold together atomic nuclei. Without the strong nuclear force, complex nuclei cannot form.

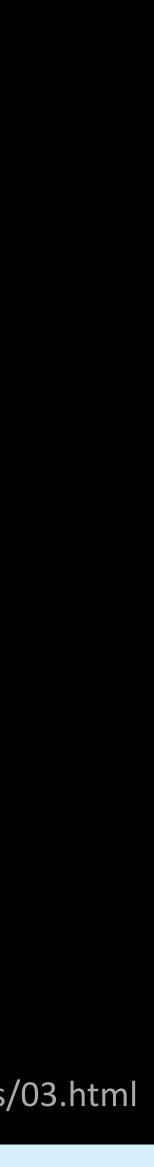
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Breaking the bond

Enormous energy is released as gamma rays and nuetrinos when the strong nuclear force is broken between protons and neutrons.

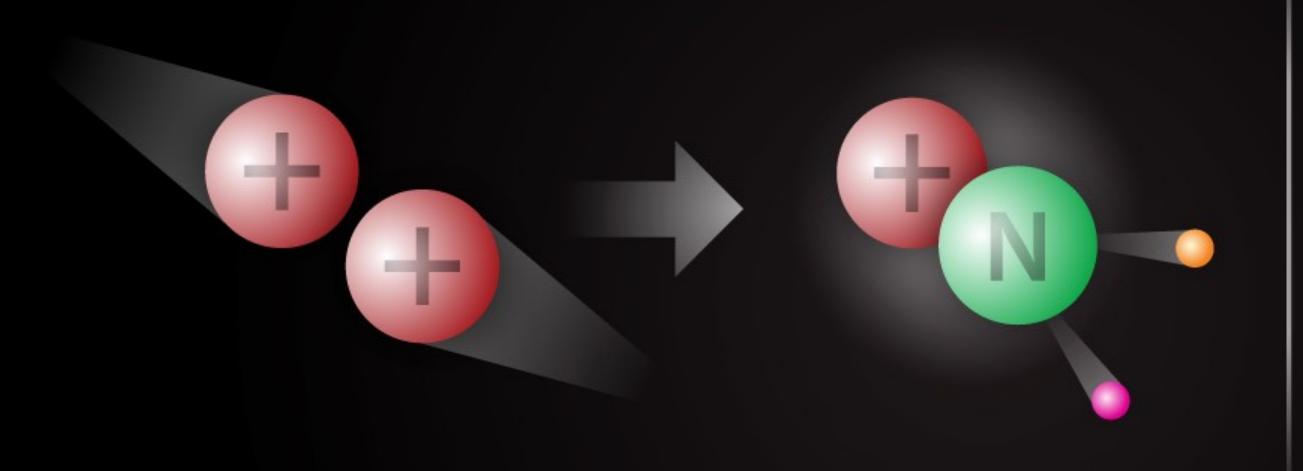
http://ecuip.lib.uchicago.edu/multiwavelength-astronomy/astrophysics/03.html



15

Weak nuclear force

Weak Nuclear Force [🗾 🚺

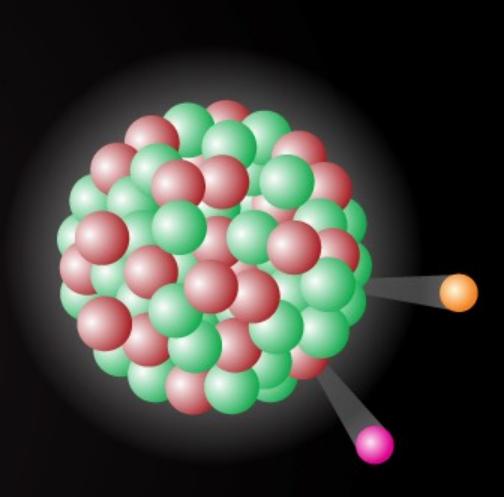


Converting protons into neutrons

When two protons collide and fuse, a disruption in the weak nuclear force emits a positron and neutrino, which converts one of the positively charged proton to a neutrally charged Nuetron. Without the weak nuclear force converting protons into nuetrons, certain complex nuclei cannot form.

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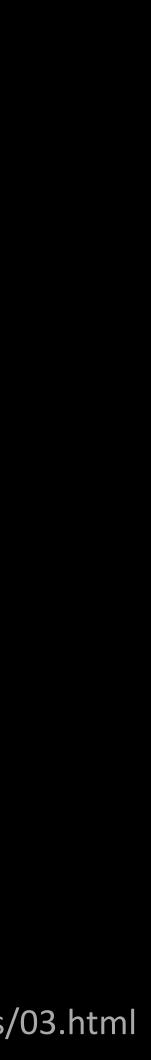




Releasing radiation

Heavy atoms have an imbalance of protons and nuetrons, so the weak nuclear force converts protons to nuetrons releasing radiation.

http://ecuip.lib.uchicago.edu/multiwavelength-astronomy/astrophysics/03.html





Gravity

Gravity

Center of Mass

The Sun and the planets all orbit a shared center of mass

Adding motion to the Universe

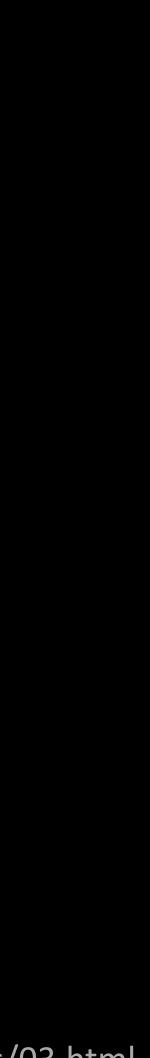
Gravity forms stars, planets, and moons, and forces these objects to spin on an axis and move along an orbital path. The planets appear to be orbiting the center of the Sun, but the Sun and planets all orbit a shared center of mass. Planets with enough mass can develop orbiting moons or rings of debris.

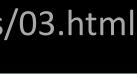
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Creating energy

Gravity is the force that creates pressure and fusion energy in the core of stars allowing them to burn for millions of years.

http://ecuip.lib.uchicago.edu/multiwavelength-astronomy/astrophysics/03.html



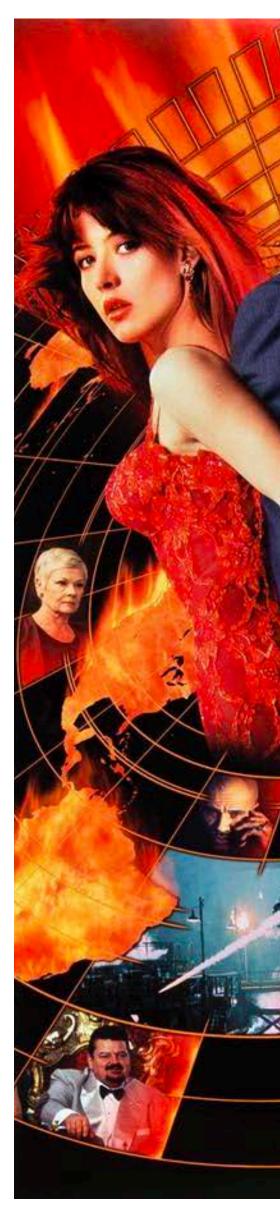




Goals of today high energy particle physics

The SM is one of the most successful models in physics, but *is it enough?*

- SM tells you how, but not why:
- Families 3 families of quark/lepton
- Number of parameters 19 free parameters
- Some phenomenon not explained by the SM:
 - Gravity not explained, why so weak?,
 SM incompatible with general relativity
 - Dark Matter & Dark Energy accounts for 95% mass of universe but not included in SM
 - Matter/anti-matter asymmetry SM does not explain the amount of matter/ anti-matter asymmetry at Big Bang



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PIERCEBROSNAN JAN (EMMES)

Standard Model IS Not FI

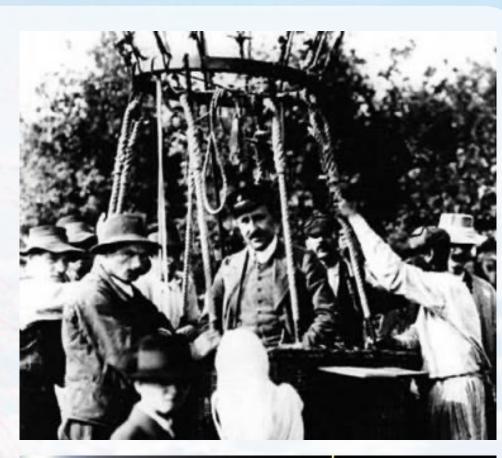


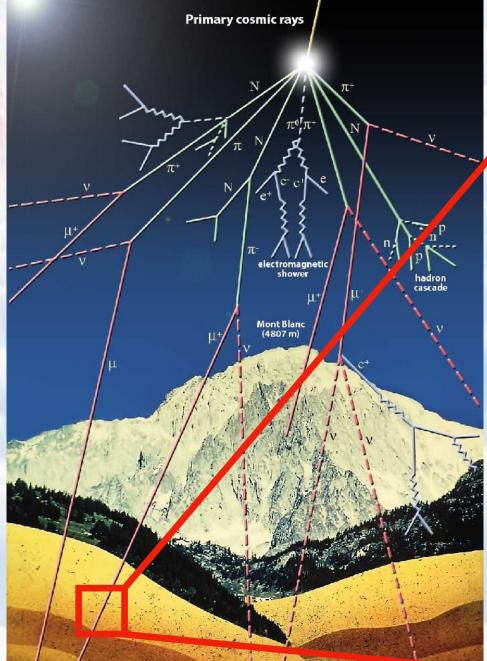


Cosmic rays: Particles from outer space accelerators

In 1912, Austrian physicist Victor Hess took an ionization chamber aloft in balloon and measured background radiation. He found that from 2000 meters to 5300 meters the amount of radiation increased, indicating the radiation came from space (Hess ruled out the Sun as the radiation's source by making a balloon ascent during a near-total eclipse). He had discovered "Cosmic Rays".

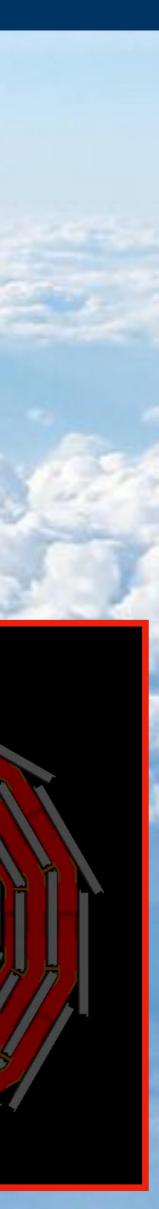
Hess shared the 1936 Nobel prize in physics for his discovery, and cosmic rays have proved useful in physics experiments – including several at CERN – since. Cosmic-ray showers were found to contain many different types of particles. Accelerators study these particles in detail.













27 Jan 1971 ... the era of hadron collider has begun





A vacuum chamber where the proton beams collide in the ISR http://cds.cern.ch/record/41966

The CERN story on hadron collider started from here

On 27 January 1971, Kjell Johnsen, who led the construction team which built the Intersecting Storage Rings (ISR), announced that the first ever interactions from colliding protons had been recorded. http://cds.cern.ch/record/39571

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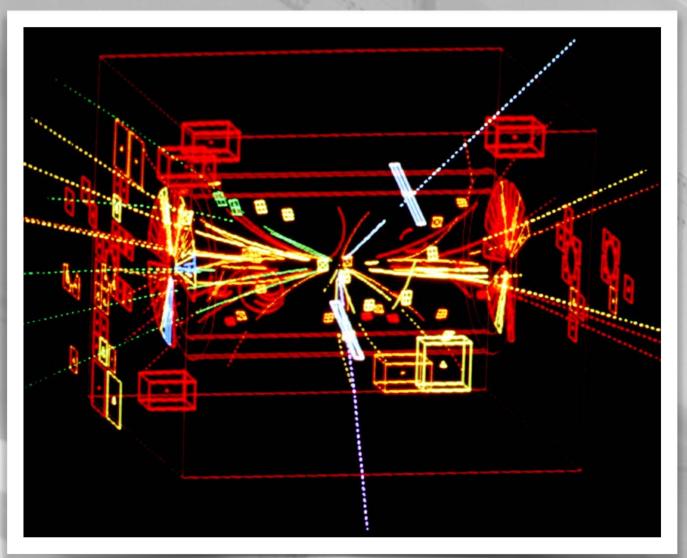
https://www.reddit.com/r/CERN/comments/8xnwz1/the_map_of_cern_and_cutaway_illustration_of_lhcb/



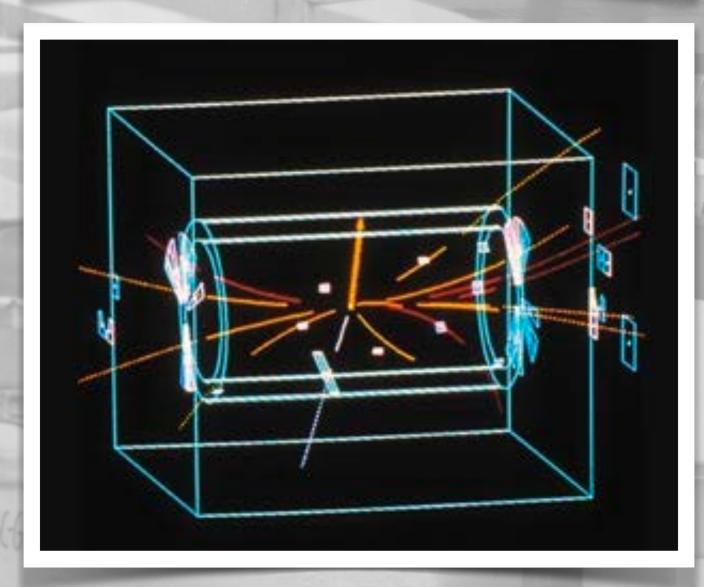




17 June 1976, SPS ... plan changed to $p\bar{p}$ collider ... discovery



30 April 1983, Image taken by the UA1 which later confirmed to be Z candidate decays to electron-positron pair



UA1 detected the W candidate event with electron and high missing energy

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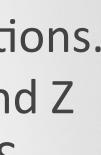
S. Weinber **Abdus Salam** S. L. Glashow (1933-202 (1926-1996)(1932-)

Around 1968, theorists came up with the electroweak theory, which unified electromagnetism and weak interactions. The theory postulated the existence of W and Z bosons. CERN decided to modify SPS to SppS.

w mass is still

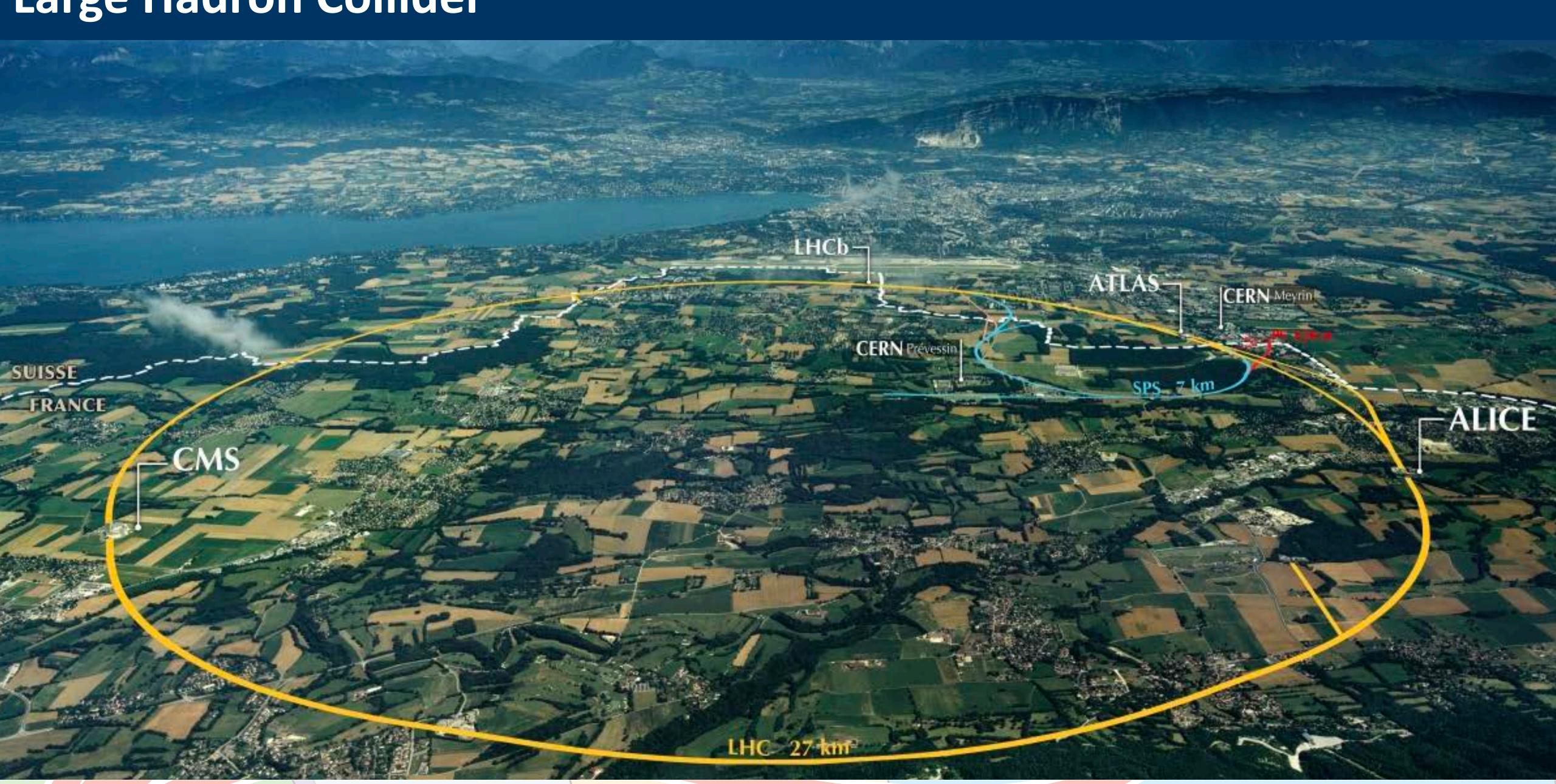
in discussion.







Large Hadron Collider





Large Hadron Collider by numbers

27KM (16 MILES) IN CIRCUMFERENCE



IN RAW DATA GENERATED BY LHC EXPERIMENTS

http://www.intelfreepress.com/news/cern-upgrades-data-center-and-restarts-large-hadron-collider/9819/

ACHIEVED BY PARTICLES

OFFOIDO INTERNAL OPERATING TEMPERATURE

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OLLISIONS

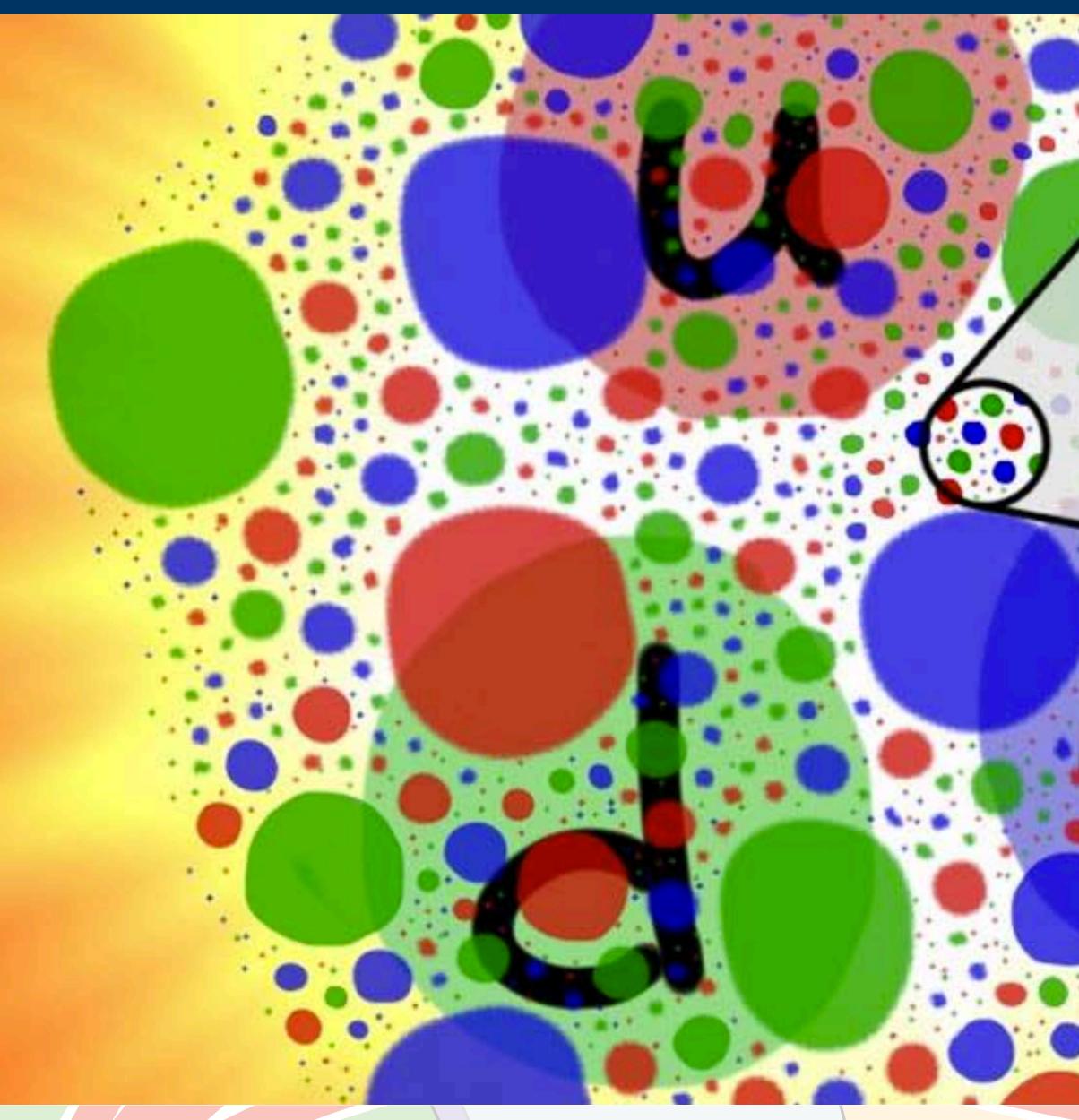
OCCUR PER SECOND

HEAT GENERATED BY COLLISIONS

CERN'S OPENSTACK CLOUD ACROSS TWO DATA CENTERS

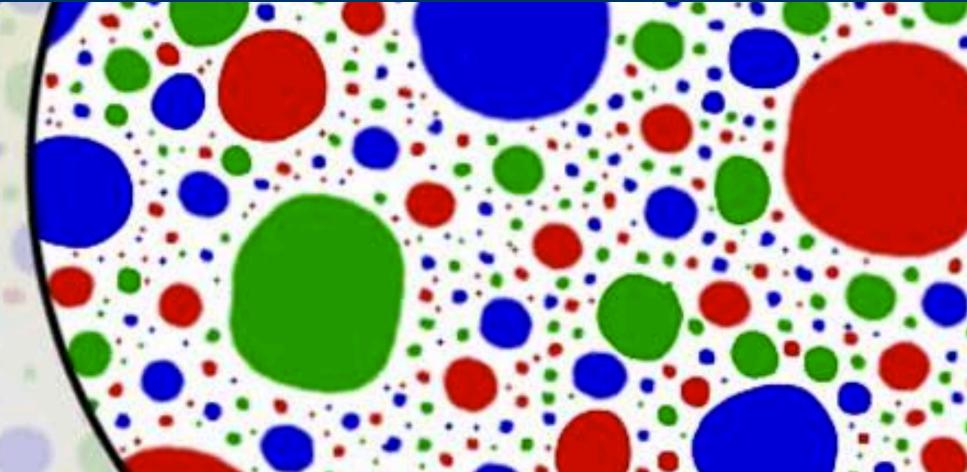


Parton - Proton



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http://www.fnal.gov/pub/today/archive/archive_2012/today12-05-04_NutshellReadMore.html

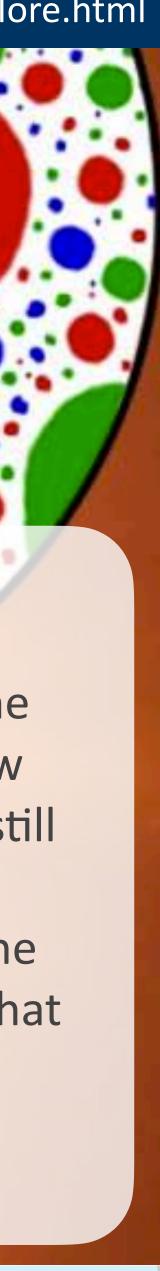


Parton

1964, Murray Gell-Mann and George Zweig

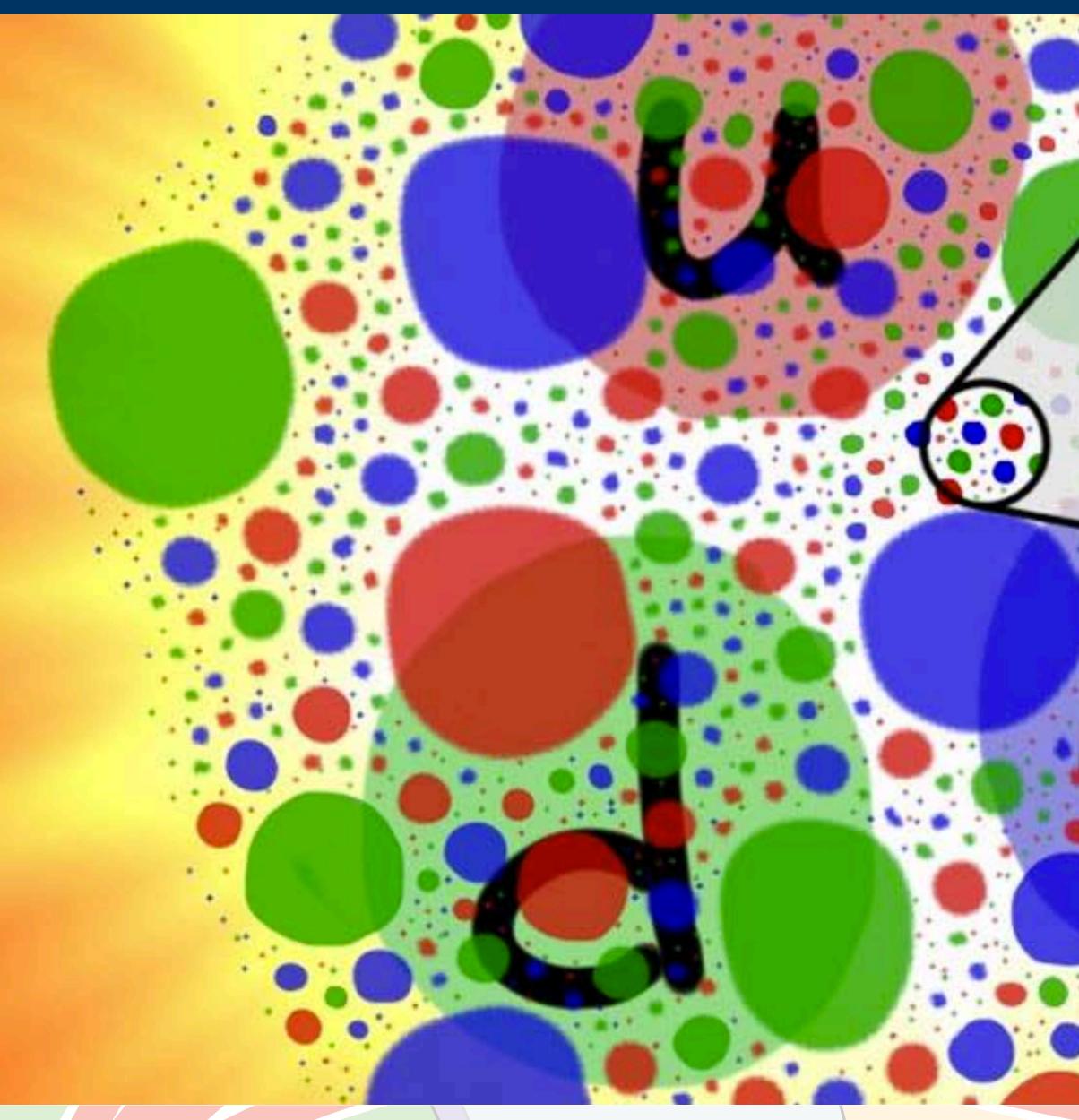
independently proposed that the proton (and also the neutron) consisted of three smaller particles. We now use Gell-Mann's name for them: quarks. They were still theoretical.

1968, SLAC use electron to probe the structure of the proton. However initially it was impossible to show that these proton constituents were quarks. Richard Feynman called them as parton in 1969. Referred today as quarks and gluons.



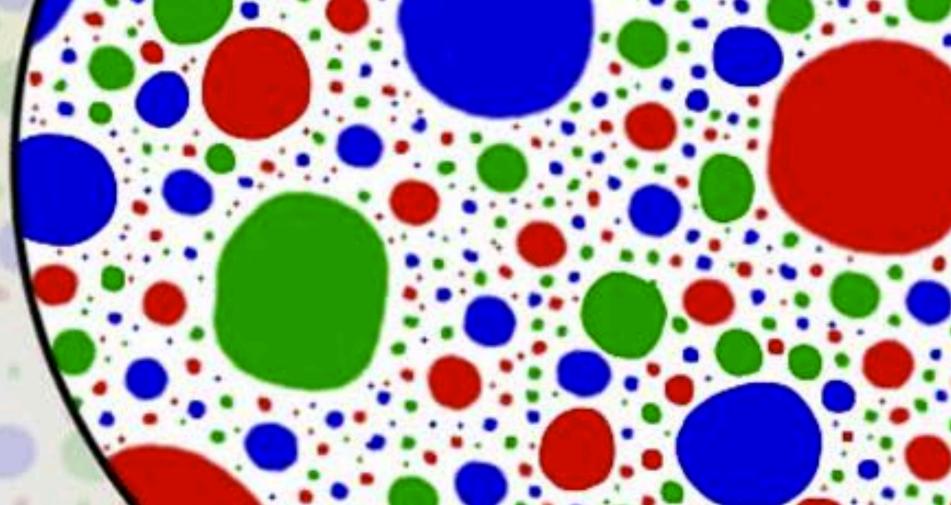


Parton - Proton



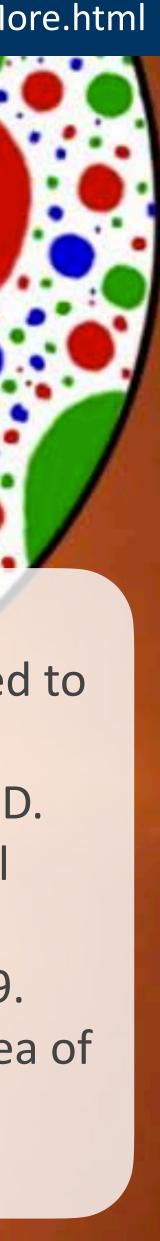
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http://www.fnal.gov/pub/today/archive/archive_2012/today12-05-04_NutshellReadMore.html



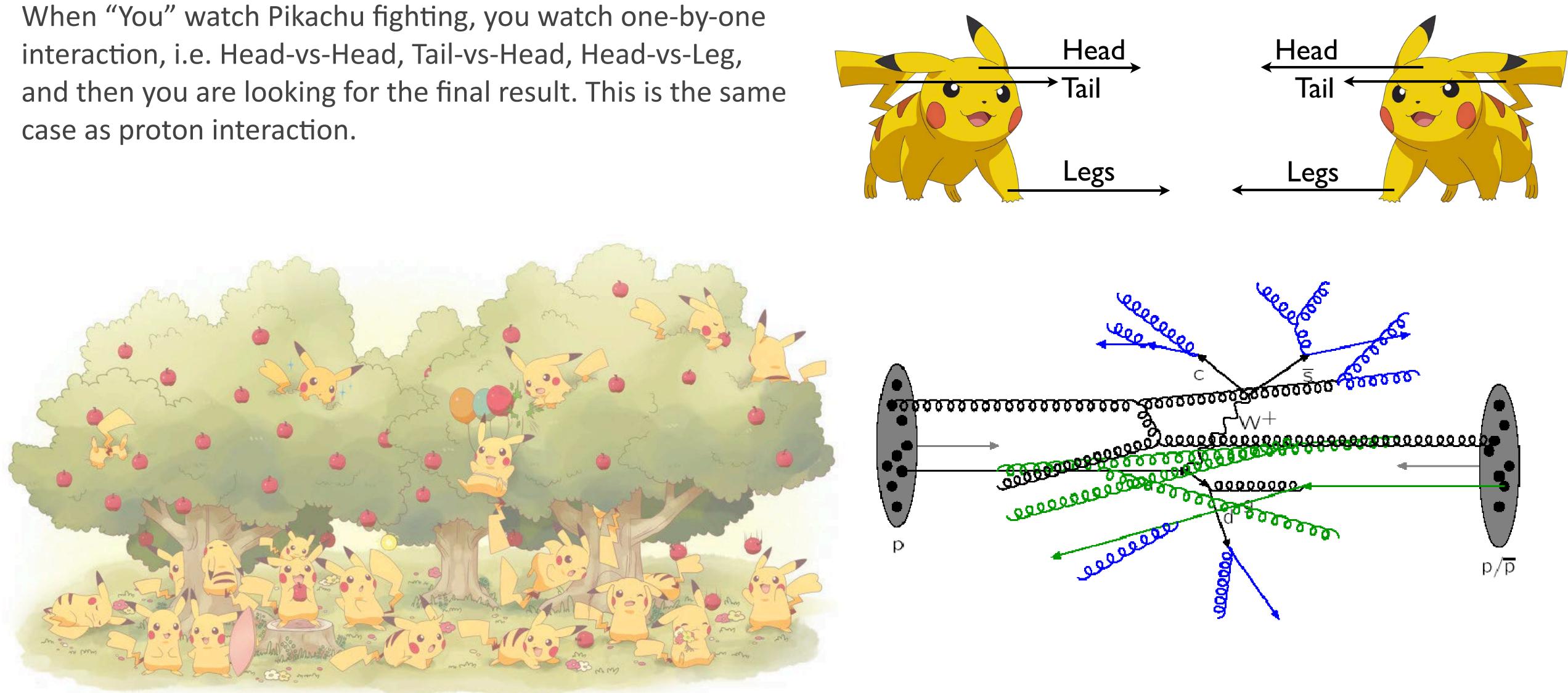
Proton

Attempts to understand partons inside the proton led to the generation of the current theory of strong force interactions, called quantum chromodynamics, or QCD. This theory postulated that there would be additional particles in the proton called gluons. Gluons were observed at the HERA accelerator in Germany in 1979. Current model: valence quarks are imbedded in a sea of virtual quark-antiquark pairs generated by the gluons which hold the quarks together in the proton.





Proton collision





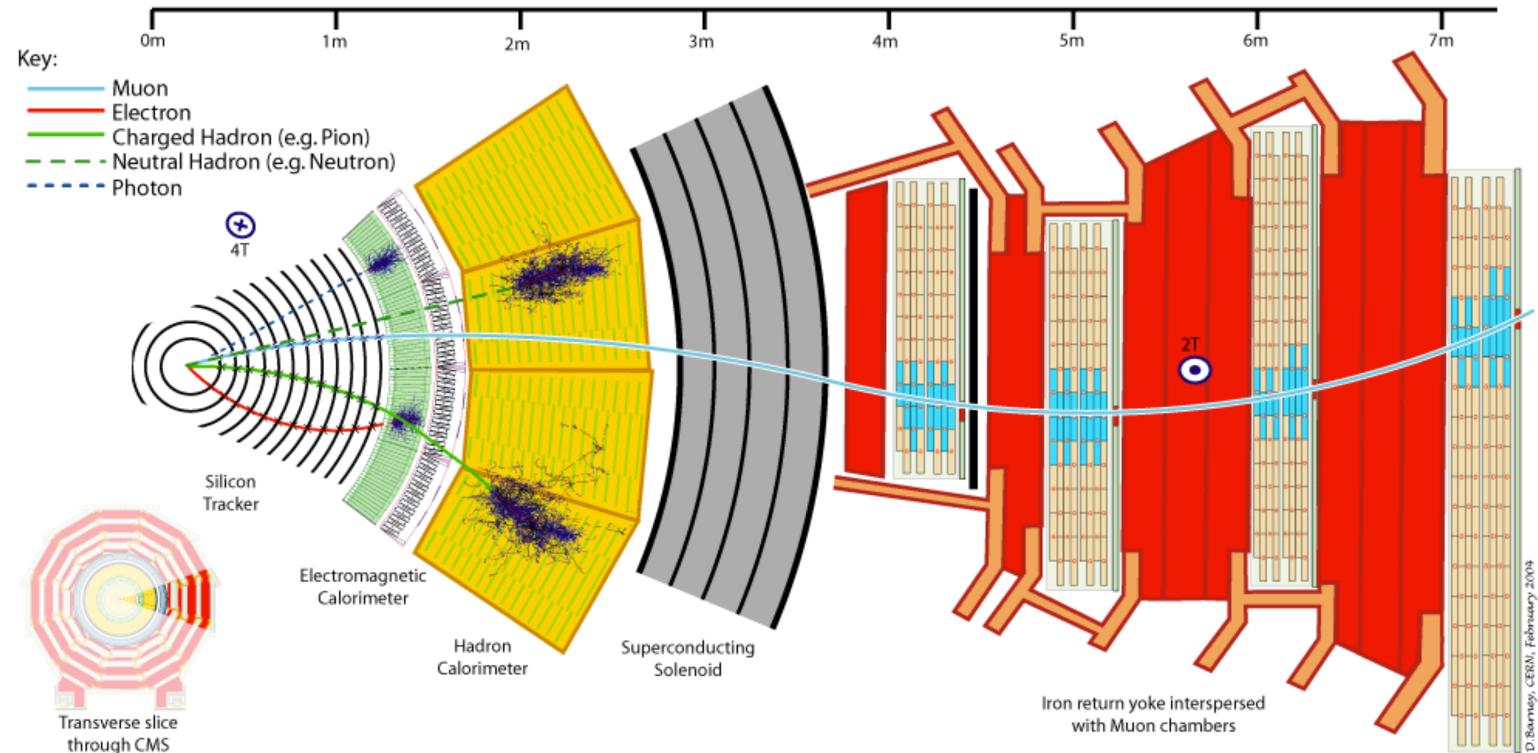
Particle detection

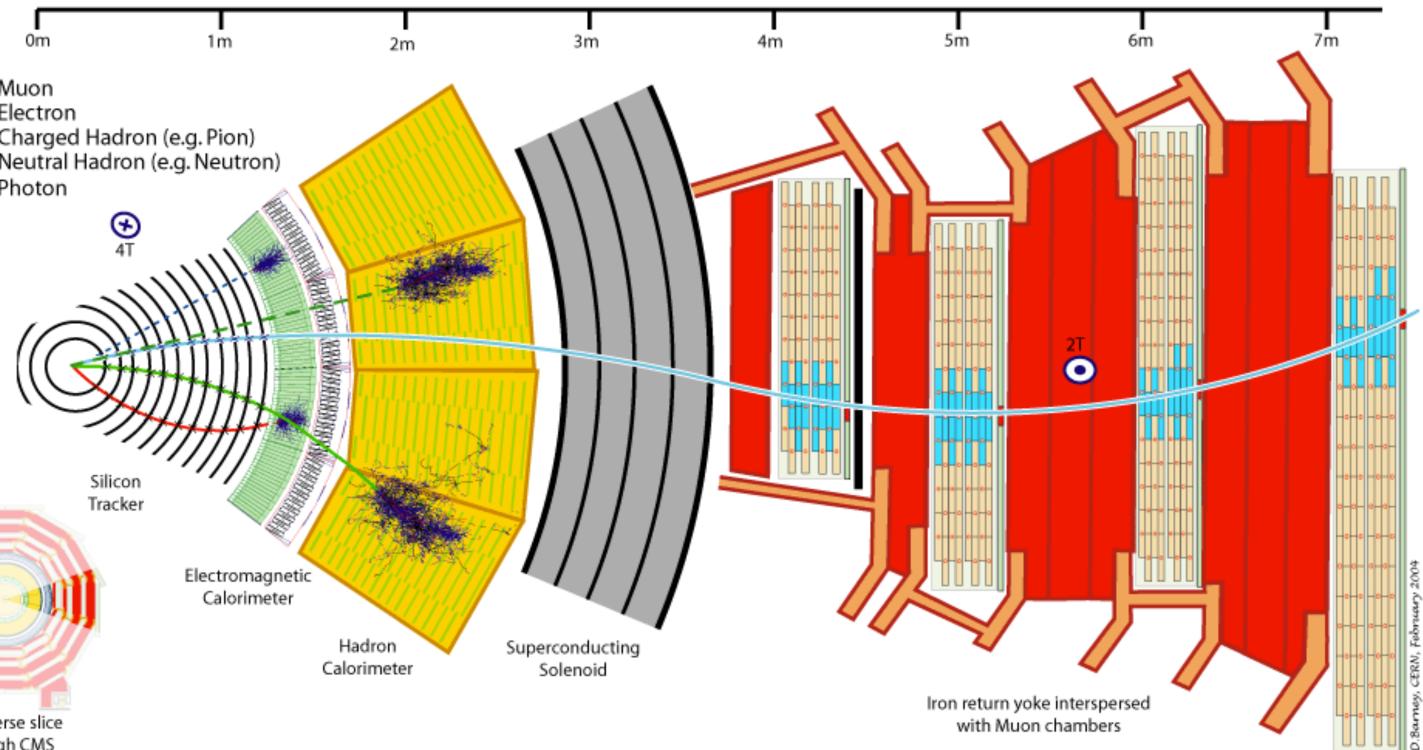
In experimental physics, a particle detector or radiation detector is an instrument used to detect, to track and to identify elementary particles by measuring one or more properties of them.

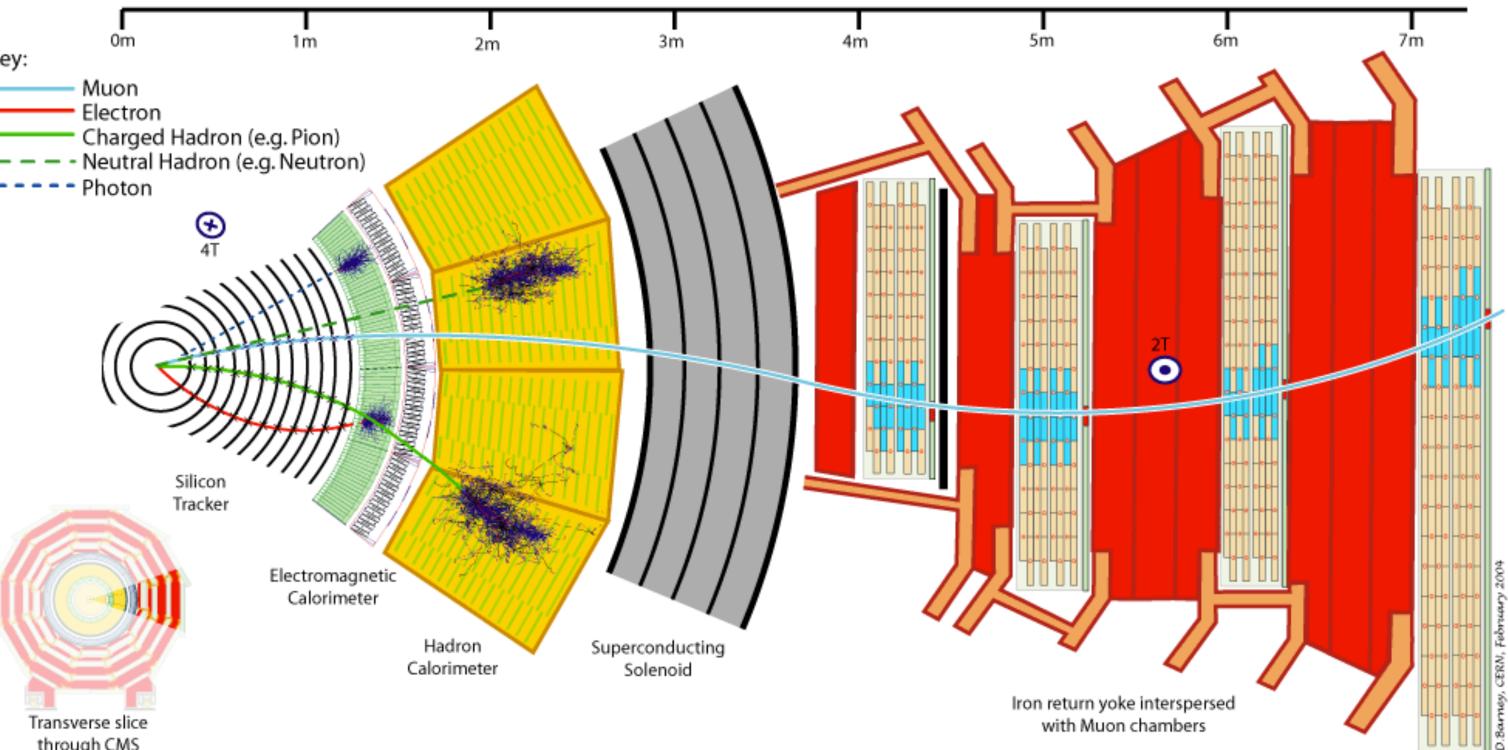
Particle detectors are devices producing an observable signal when they are crossed by a particle. Usually they are made by an active element (such that there is some interaction with the particle) and by a readout system ("forming" the signal and sending it to the data acquisition chain)

Aim: to detect as many of the stable and longlived particles produced in a particle collision.

Need to measure: charge, mass, energy, direction.









Standard Model

The Standard Model (SM) of particle physics is the theory describing three of the four known fundamental forces (the electromagnetic, weak, and strong interactions, and not including the gravitational force) in the universe, as well as classifying all known elementary particles.



Steven Weinberg



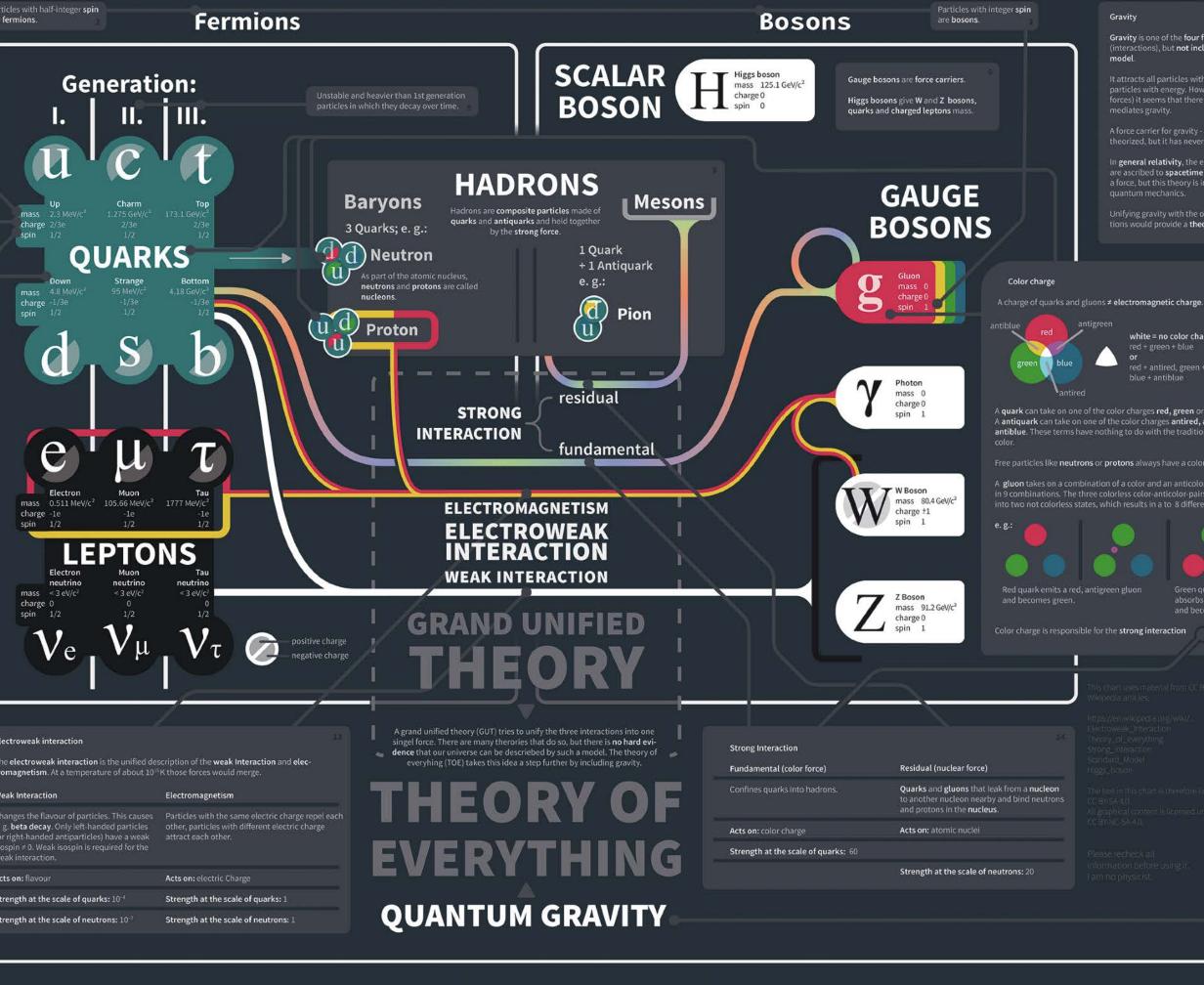
Sheldon Glashow

Abdus Salam

THE STANDARD MODEL OF PARTICLE PHYSICS

ee electric charge

β⁺ decay



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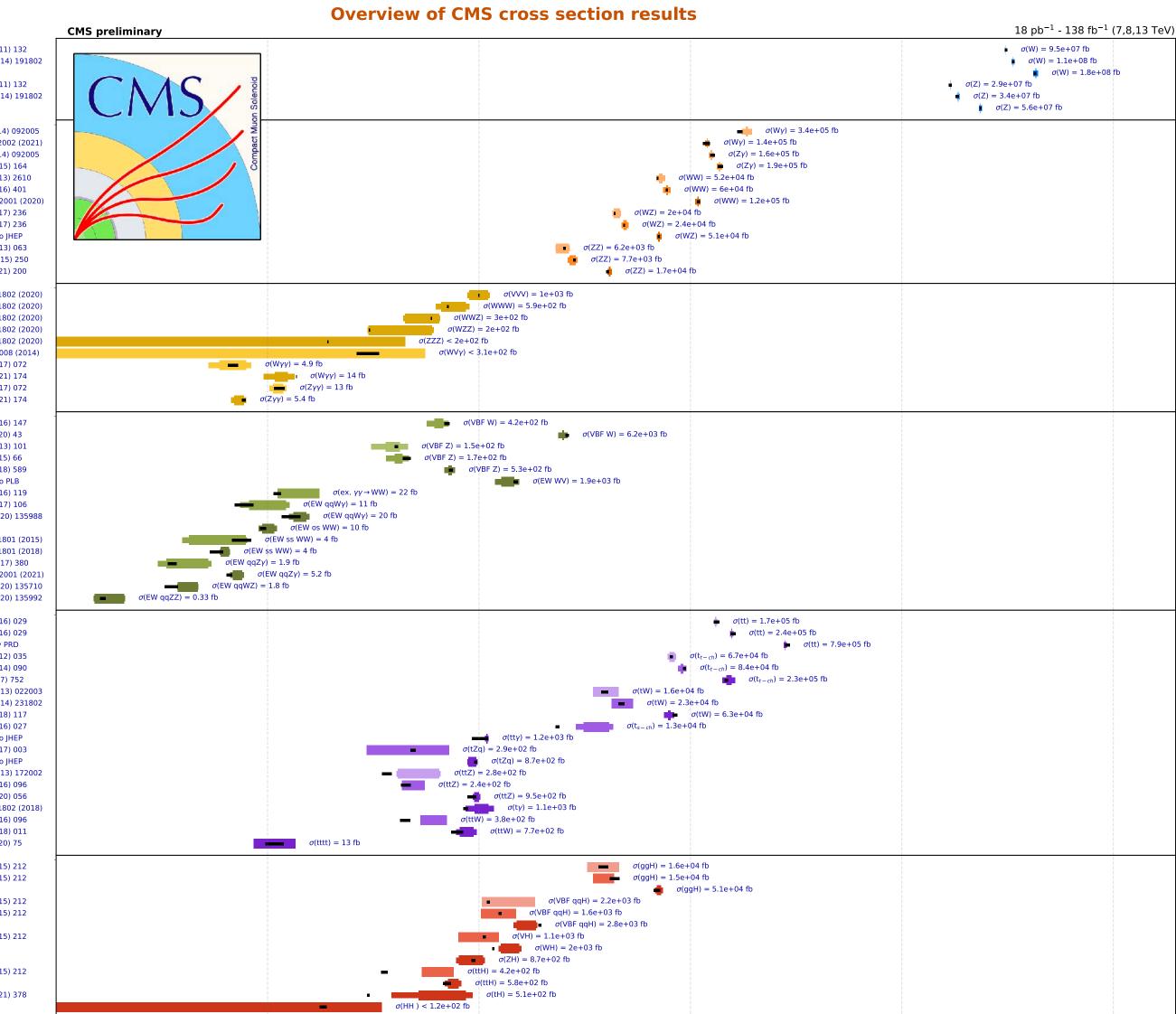
LHC with Standard Model

Since the 7 TeV data collection started in 2011, until now, LHC physics program shows that SM (still) works very well and Higgs (with mass 125 GeV) is there. However, one need to remember than studying known processes is very challenging:

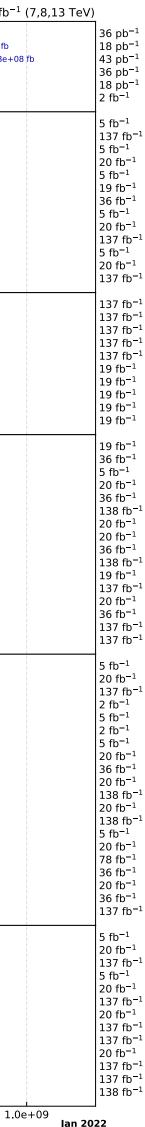
- Excellent performance and calibration
- Probe uncovered parameter space
- Try new techniques to enhance Signal/ Background separation, i.e. new machine learning techniques

¥	W	7 TeV	JHEP 10 (2011
troweak	W W	8 TeV 13 TeV	PRL 112 (2014 SMP-15-004
tro	Z	7 TeV	JHEP 10 (2011
Elec	Z	8 TeV	PRL 112 (2014
_	Z	13 TeV	SMP-15-011
	\mathbb{W}_{Y} \mathbb{W}_{Y}	7 TeV 13 TeV	PRD 89 (2014 PRL 126 2520
	Zγ	7 TeV	PRD 89 (2014
	Zγ	8 TeV	JHEP 04 (2015
_	WW	7 TeV	EPJC 73 (2013
sor	WW	8 TeV	EPJC 76 (2016
di-Boson	ww wz	13 TeV 7 TeV	PRD 102 0920 EPJC 77 (2017
9	WZ	8 TeV	EPJC 77 (2017
	WZ	13 TeV	Submitted to
	ZZ	7 TeV	JHEP 01 (2013
	ZZ ZZ	8 TeV 13 TeV	PLB 740 (201 EPJC 81 (2021
	VVV	13 TeV	PRL 125 1518
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	VBF Z	8 TeV	EPJC 75 (2015
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VBF and VBS	EW qqW γ	8 TeV	JHEP 06 (2017
Far	EW qqWγ	13 TeV	PLB 811 (2020
VB	EW os WW EW ss WW	13 TeV 8 TeV	SMP-21-001 PRL 114 0518
	EW ss WW	13 TeV	PRL 120 0818
	EW qqZ γ	8 TeV	PLB 770 (201
	EW qqZγ	13 TeV	PRD 104 0720
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	tt	7 TeV	JHEP 08 (2016
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	t_{t-ch}	13 TeV	PLB 72 (2017)
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	tW	8 TeV	PRL 112 (2014
	tW t _{s – ch}	13 TeV 8 TeV	JHEP 10 (2018 JHEP 09 (2016
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		8 TeV	JHEP 07 (2017
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	tZq	13 TeV	Submitted to
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	tZq ttZ	<mark>13 TeV</mark> 7 TeV	Submitted to PRL 110 (2013 JHEP 01 (2016 JHEP 03 (2020
	tZq ttZ ttZ ttZ tγ ttW	13 TeV 7 TeV 8 TeV 13 TeV 13 TeV 8 TeV	Submitted to PRL 110 (201 JHEP 01 (2016 JHEP 03 (2020 PRL 121 2218 JHEP 01 (2016
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	tZq ttZ ttZ ttZ tγ ttW ttW ttW	13 TeV 7 TeV 8 TeV 13 TeV 13 TeV 8 TeV 13 TeV 13 TeV 13 TeV	Submitted to PRL 110 (2012) JHEP 01 (2016) JHEP 03 (2020) PRL 121 2218 JHEP 01 (2016) JHEP 08 (2018) EPJC 80 (2020)
	tZq ttZ ttZ ttZ tγ ttW ttW	13 TeV 7 TeV 8 TeV 13 TeV 13 TeV 8 TeV 13 TeV	Submitted to PRL 110 (2012) JHEP 01 (2016) JHEP 03 (2020) PRL 121 2218 JHEP 01 (2016) JHEP 08 (2018) EPJC 80 (2020) EPJC 75 (2015)
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s66	tZq ttZ ttZ ttZ tγ ttW ttW ttW tttt ggH ggH ggH VBF qqH	13 TeV 7 TeV 8 TeV 13 TeV 13 TeV 8 TeV 13 TeV 13 TeV 7 TeV 7 TeV 7 TeV 7 TeV 7 TeV	Submitted to PRL 110 (2012) JHEP 01 (2016) JHEP 03 (2020) PRL 121 2218 JHEP 01 (2016) JHEP 08 (2018) EPJC 80 (2020) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) HIG-19-005
Higgs	tZq ttZ ttZ ttZ ttW ttW ttW ttt ggH ggH ggH ggH VBF qqH VBF qqH VBF qqH VBF qqH VH WH	13 TeV 7 TeV 8 TeV 13 TeV 13 TeV 13 TeV 13 TeV 7 TeV 8 TeV 13 TeV 8 TeV 13 TeV 8 TeV 13 TeV 8 TeV 13 TeV 13 TeV	Submitted to PRL 110 (2012) JHEP 01 (2016) JHEP 03 (2020) PRL 121 2218 JHEP 01 (2016) JHEP 08 (2018) EPJC 80 (2020) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) HIG-19-005 EPJC 75 (2015) HIG-19-005
Higgs	tZq ttZ ttZ ttZ ty ttW ttW ttW ttW tttt ggH ggH ggH ggH VBF qqH VBF qqH VBF qqH VBF qqH VBF qqH VH WH ZH	13 TeV 7 TeV 8 TeV 13 TeV 13 TeV 8 TeV 13 TeV 13 TeV 8 TeV 13 TeV 8 TeV 13 TeV 8 TeV 13 TeV 13 TeV 13 TeV 13 TeV	Submitted to PRL 110 (2012) JHEP 01 (2016) JHEP 03 (2020) PRL 121 2218 JHEP 01 (2016) JHEP 08 (2018) EPJC 80 (2020) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) HIG-19-005 HIG-19-005 HIG-19-005
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Higgs	tZq ttZ ttZ ttZ ty ttW ttW ttW ttW tttt ggH ggH ggH ggH VBF qqH VBF qqH VBF qqH VBF qqH VBF qqH VH WH ZH	13 TeV 7 TeV 8 TeV 13 TeV 13 TeV 8 TeV 13 TeV 13 TeV 8 TeV 13 TeV 8 TeV 13 TeV 8 TeV 13 TeV 13 TeV 13 TeV 13 TeV	Submitted to PRL 110 (2012) JHEP 01 (2016) JHEP 03 (2020) PRL 121 2218 JHEP 01 (2016) JHEP 08 (2018) EPJC 80 (2020) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) EPJC 75 (2015) HIG-19-005 EPJC 75 (2015) HIG-19-005

Measured cross sections and exclusion limits at 95% C.L. See here for all cross section summary plots



1.0e-01 1.0e+01 1.0e+031.0e+05 1.0e+07 Inner colored bars statistical uncertainty, outer narrow bars statistical+systematic uncertainty σ [fb] Light colored bars: 7 TeV, Medium bars: 8 TeV, Dark bars: 13 TeV Black bar theory prediction

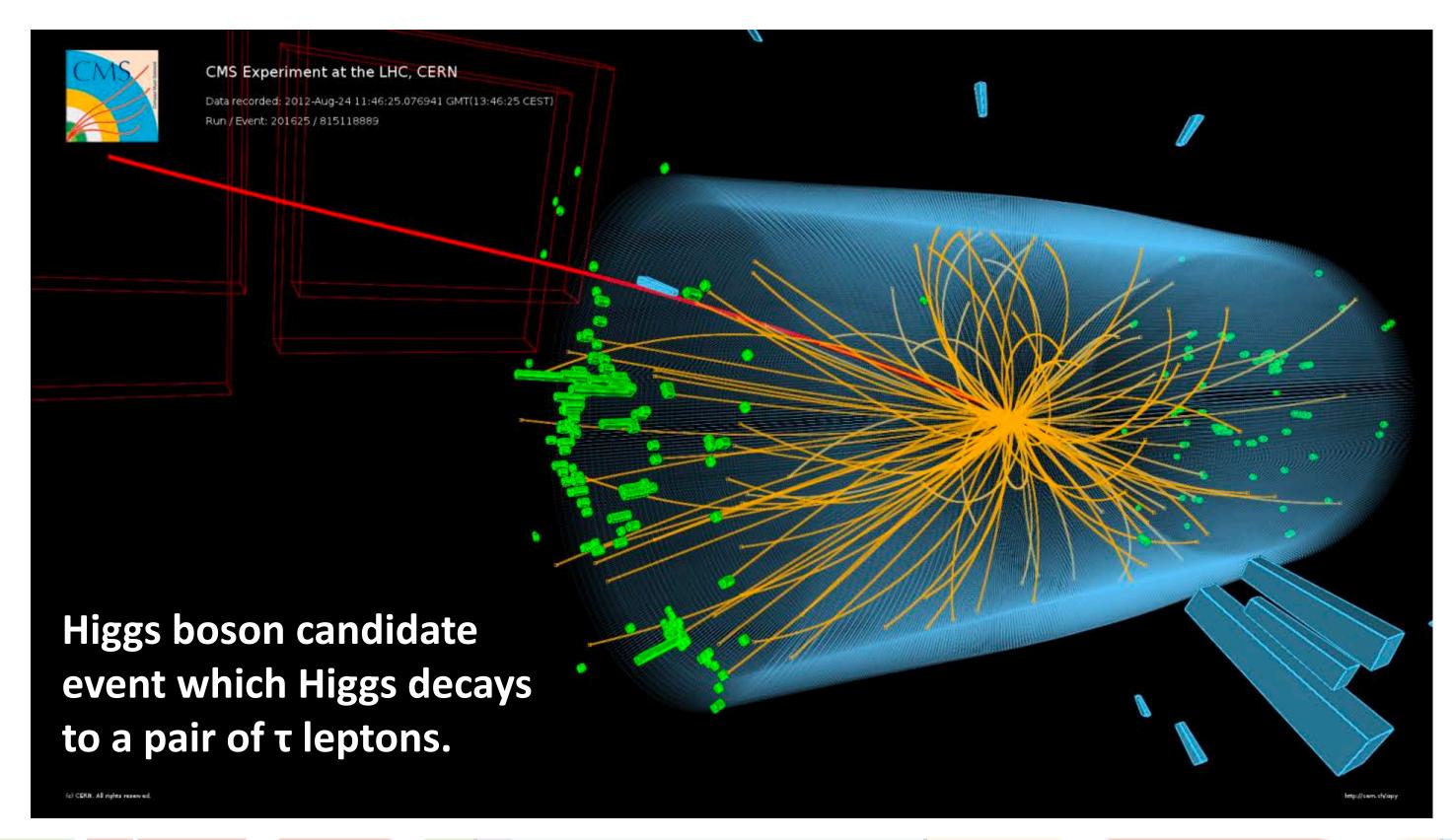




Mass of elementary particles?

One missing piece was the way to explain <u>masses of elementary</u> particles. In the mid 1960s, the mechanism to explain the mass generation came out by three independent groups,

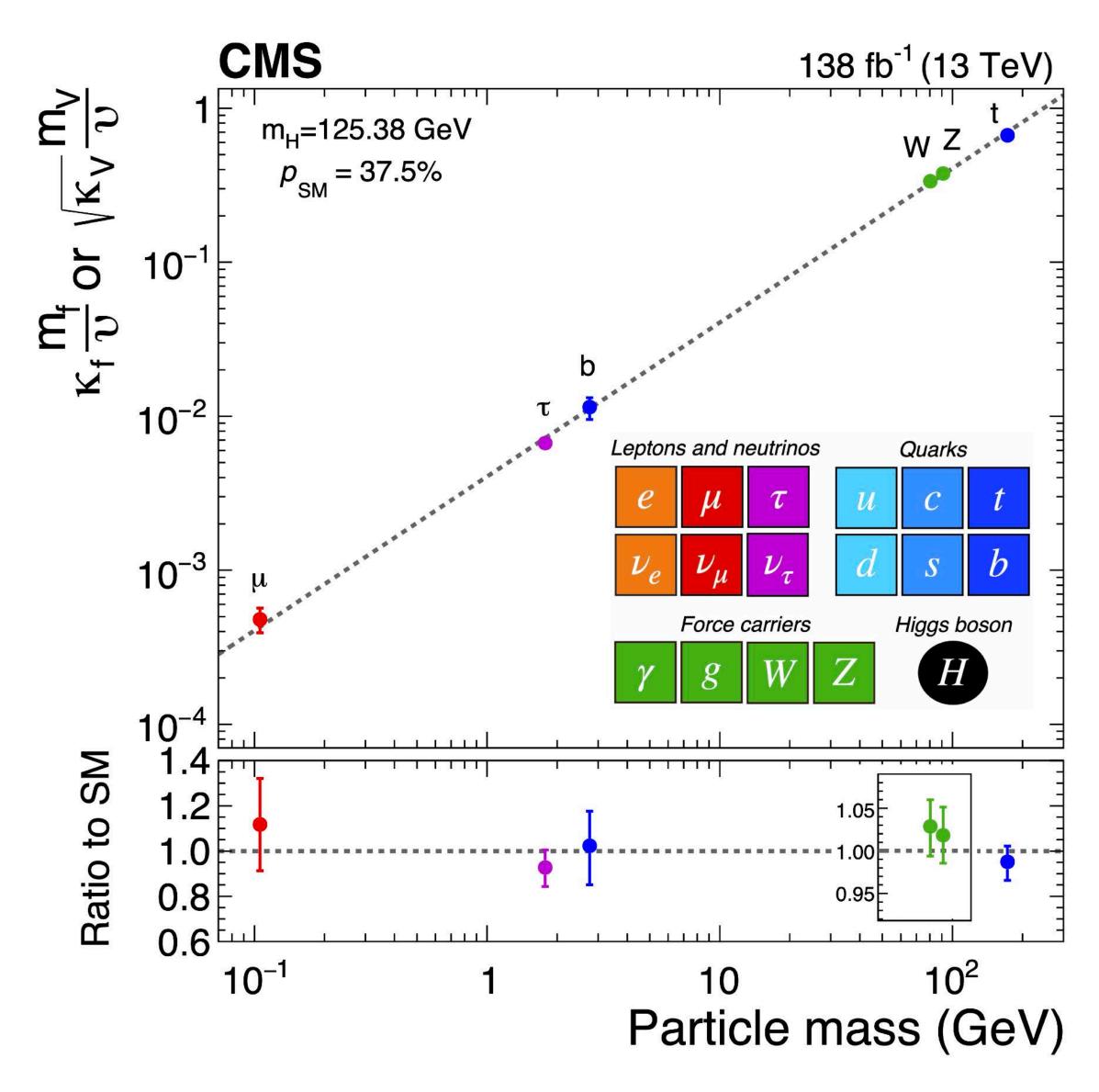
- by Robert Brout and François Englert [Phys. Rev. Lett. 13, 321];
- by **Peter Higgs** [Phys. Rev. Lett. 13, 508];
- by Gerald Guralnik, C. R. Hagen, and Tom Kibble [Phys. Rev. Lett. 13, 585].







Mass of elementary particles?



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The Higgs boson is predicted to couple to particles (or decay into them) with a strength depending on their masses in a welldefined way: the higher the mass, the stronger the coupling. The measurement of the many ways this Higgs boson decays and the measurement of its couplings to different particles, shown in the figure, provide a crucial test of the validity of existing theories.

All the measurements of the properties of the Higgs boson presently agree with the theoretical predictions within the measurement and prediction uncertainties.

[Link]

31

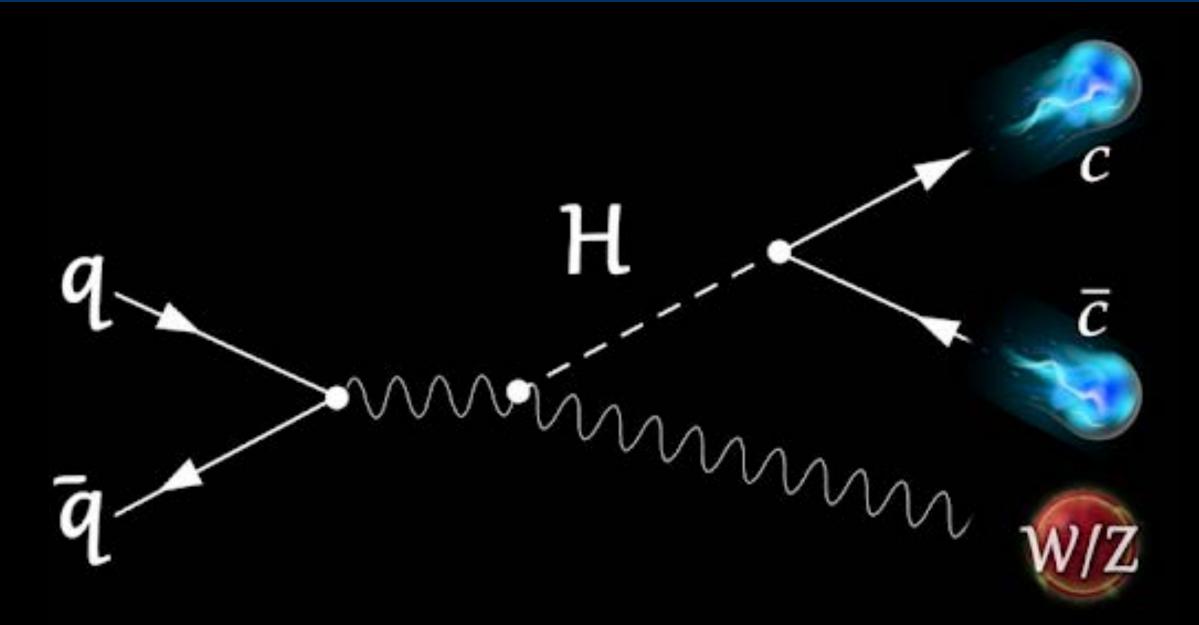
Higgs ... from search to precision measurement



CMS Experiment at the LHC, CERN Data recorded: 2018-Aug-05 09:43:33.747957 GMT Run / Event / LS: 320854 / 196048575 / 115

We know that Higgs is there, why do we need precise measurements?

Because the lack of direct observations of new particles at the LHC!, we need an alternative approach which we consider that BSM physics interfere with standard model particles and subsequently leave an imprint on their properties. With precise measurement, we may see hint(s) of new physics.

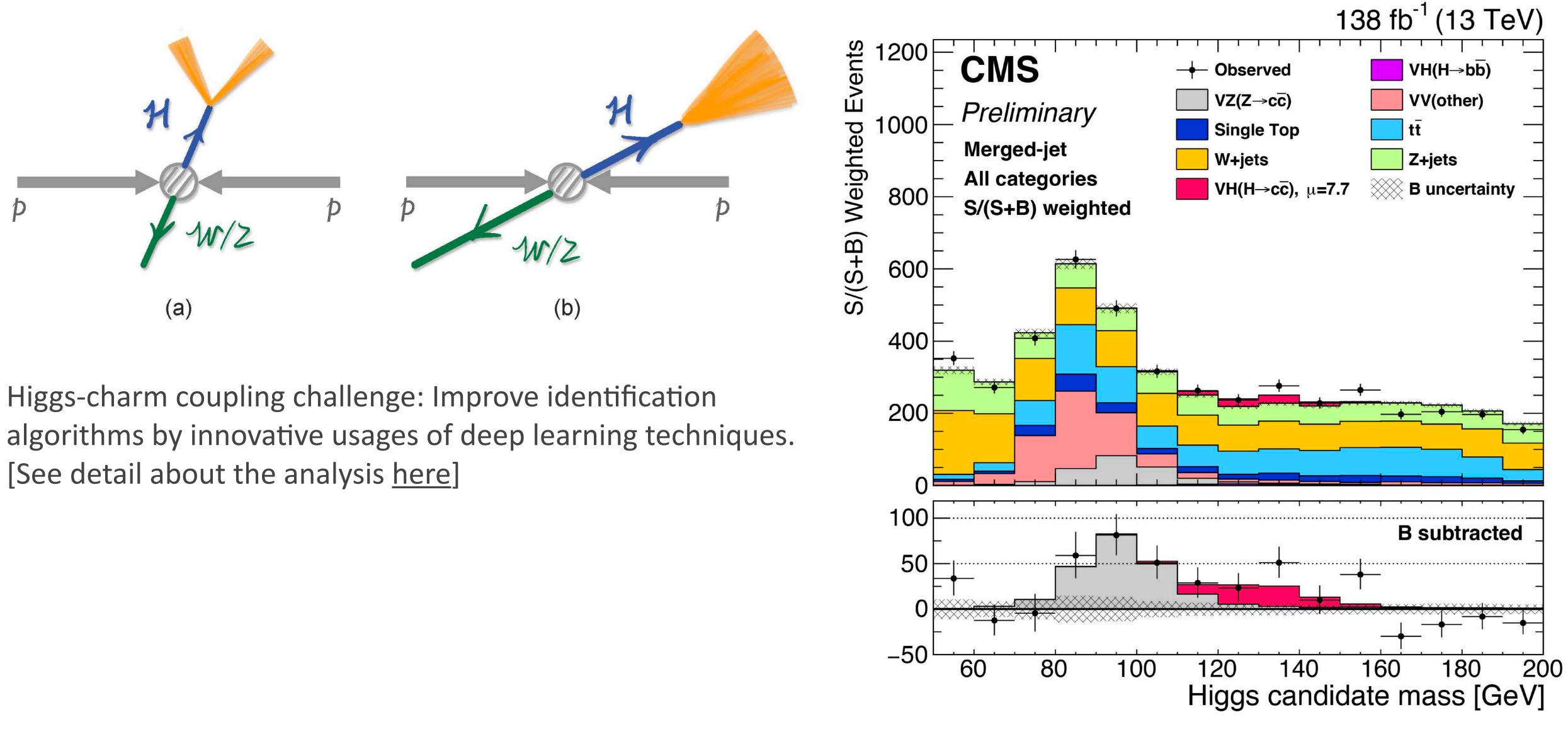








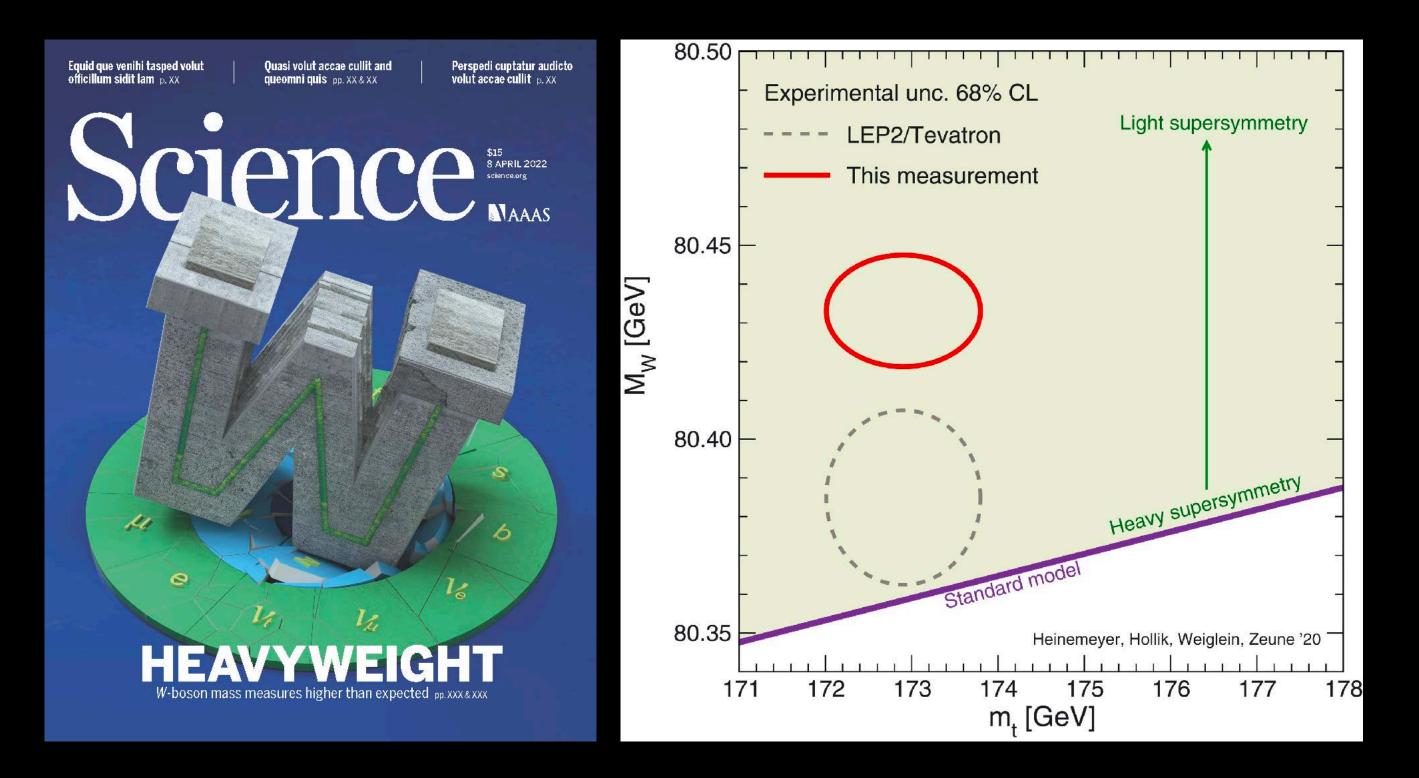
Higgs ... from search to precision measurement





Top quark ... a tool for discoveries

- Discover in 1995 by CDF and D0 at Tevatron
- To predict the top quark mass, need to know accurately the W boson and Higgs boson masses
- Consequently, use top and Higgs masses to produce W boson mass ... reported on 7 April 2022 by CDF that W boson mass extracted from data taken at the Tevatron (2002-2011)



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CMS Experiment at the LHC, CERN Data recorded: 2016-Aug-17 08:01:23.065024 GMT Run / Event / LS: 278969 / 229126383 / 184

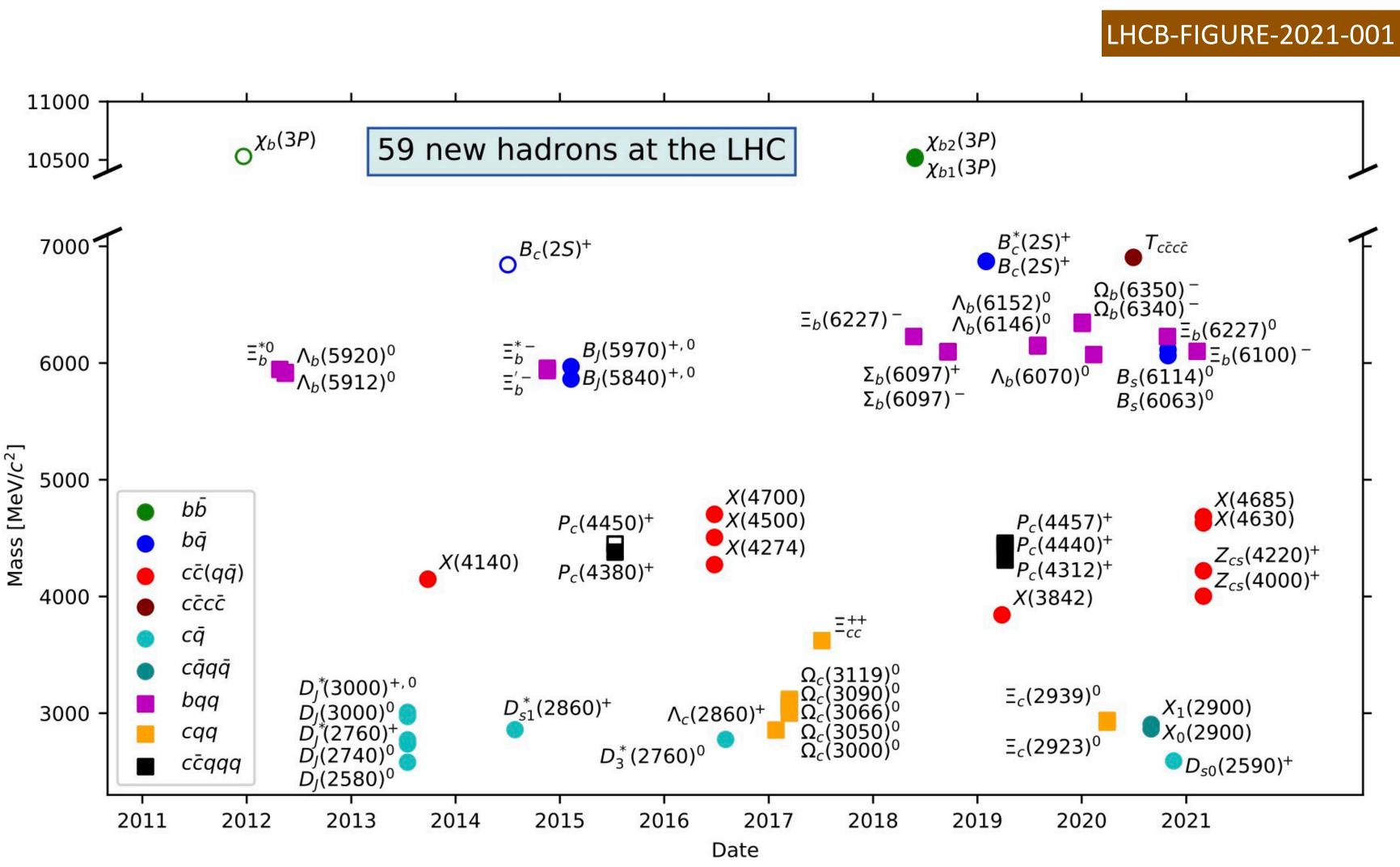




10 years (2011-2021) of LHC, 59 new hadrons

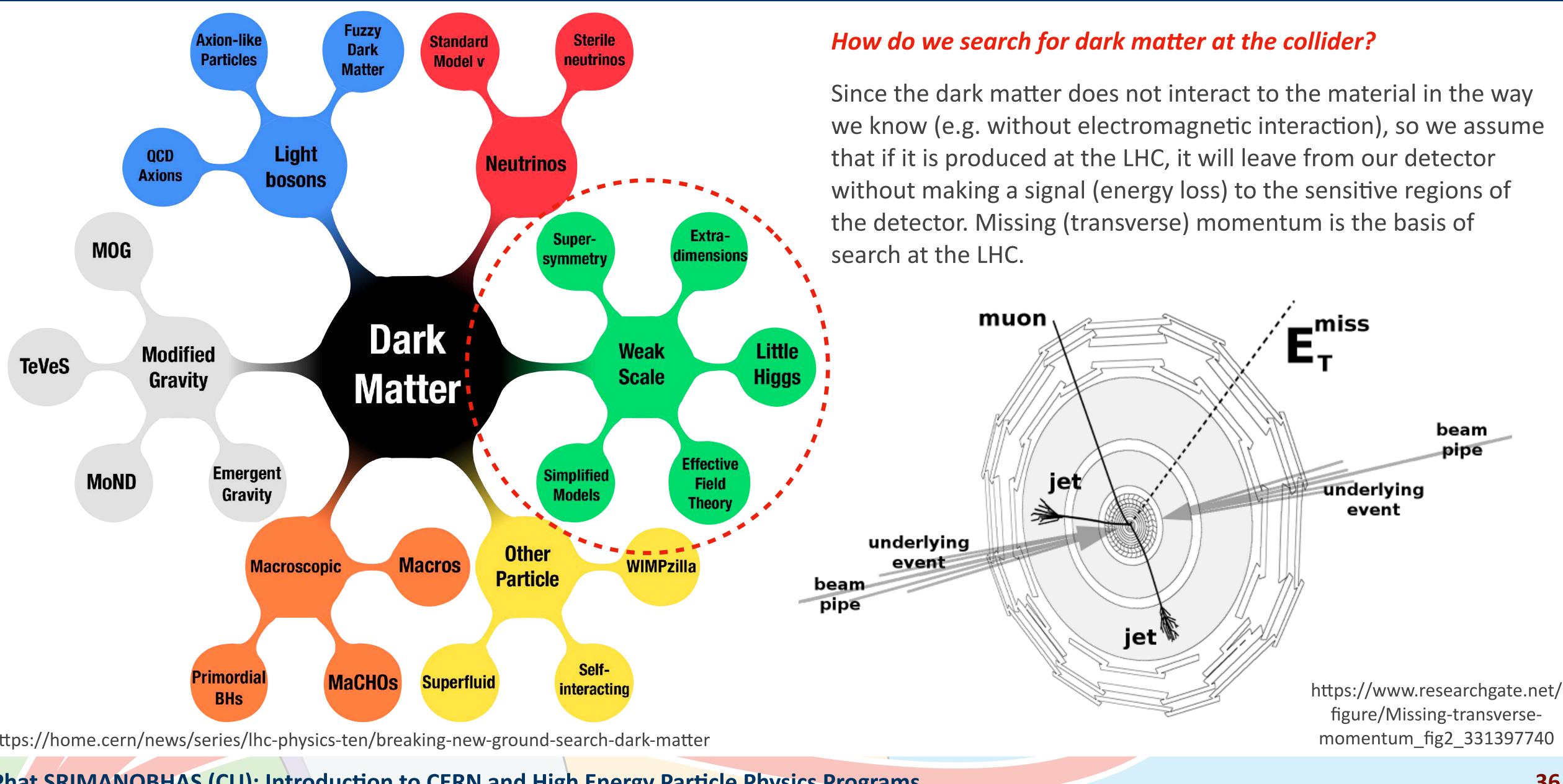
Quantum chromodynamics (or QCD) is still on very solid footing

- QCD is the theory to describe the strong interaction that holds quarks together inside hadrons.
- Experiments continue to discover hadrons including combination of
- mesons: quarks and antiquarks,
- **baryons**: three quarks,
- antibaryons: three antiquarks,
- tetraquarks: two quarks and two antiquarks,
- pentaquarks: four quarks and one antiquark.
- Reminder: we cannot (yet) prove theoretically that quarks can't stay alone. In addition, we can't also calculate of which combinations of quarks would be/ would not be viable in nature.





Exotic searches: Dark Matter searches at collider



https://home.cern/news/series/lhc-physics-ten/breaking-new-ground-search-dark-matter

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Since the dark matter does not interact to the material in the way we know (e.g. without electromagnetic interaction), so we assume that if it is produced at the LHC, it will leave from our detector without making a signal (energy loss) to the sensitive regions of the detector. Missing (transverse) momentum is the basis of



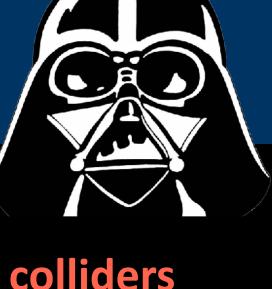
Exotic searches: Darth Vader Dark Matter



Run: 337215 Event: 2546139368 2017-10-05 10:36:30 CEST

 $E_{\pi}^{miss} = 1.9 \text{ TeV}$ jet $p_{T} = 1.9 \text{ TeV}$

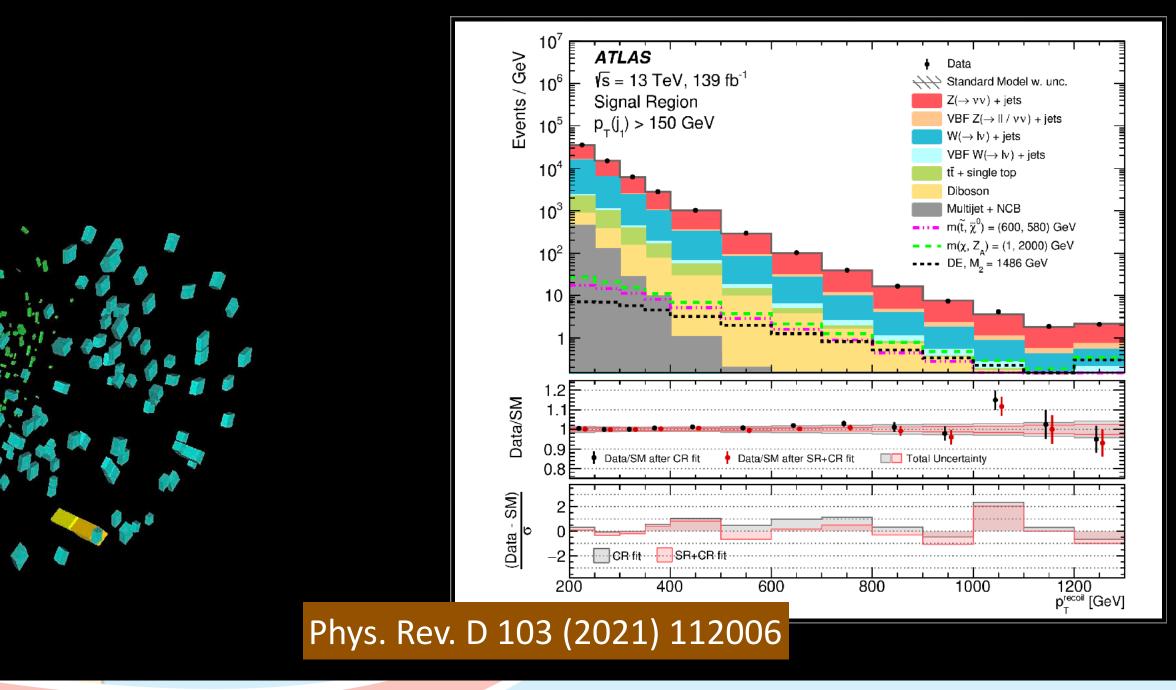




Type of DM-candidate events we study at the colliders

• Simplified models, one s-channel mediator (Mono-X)

- A pair of WIMPs that recoil against X (a visible SM particle, or a set of SM particles). Currently, X includes hadronic jet, heavy-flavor quarks, a photon, or a W or Z boson, even Higgs boson
- An imbalance in the total momentum in the plane transverse to the colliding beams as reconstructed in the detector









Exotic searches: Darth Vader Dark Matter

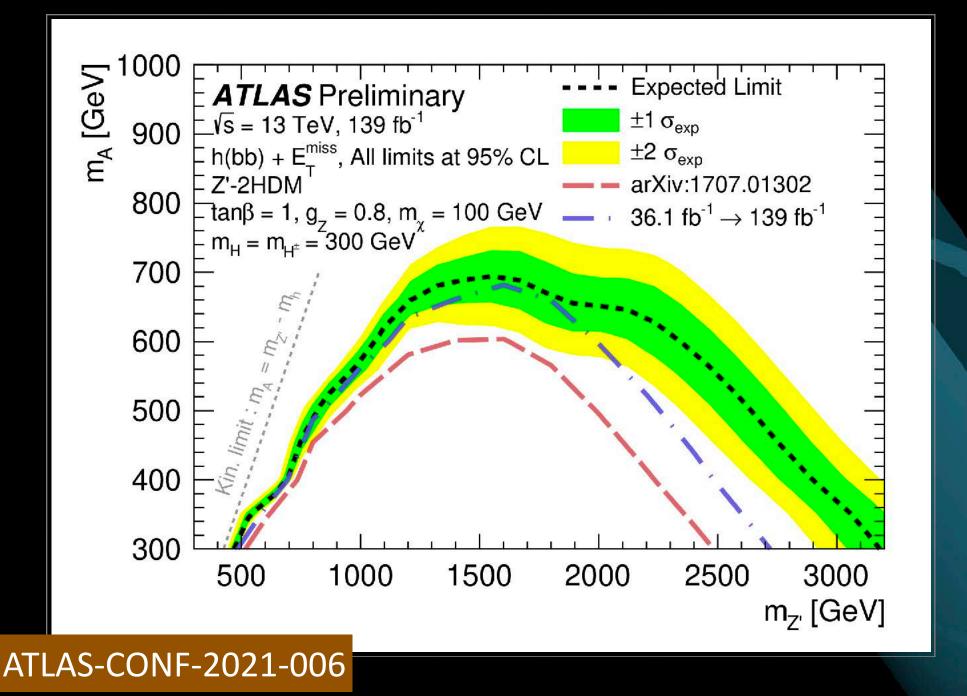


Type of DM-candidate events we study at the colliders

• two-Higgs-doublet model (2HDM)

Run: 349309 Event: 769175011 2018-05-01 13:57:22 CEST

• Example: extension with Z' then $Z' \rightarrow Ah; A \rightarrow DM+DM$



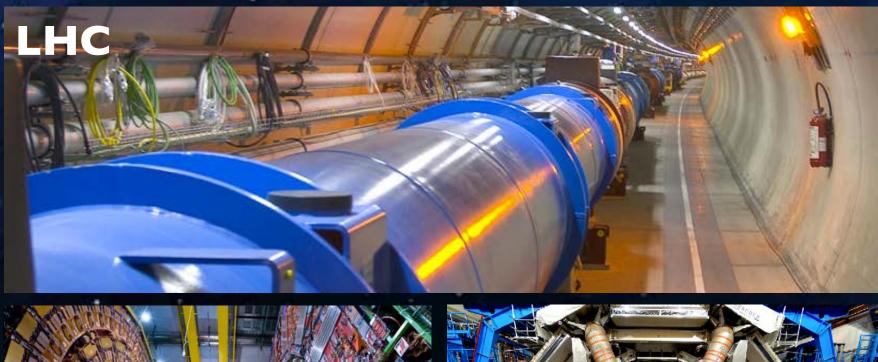


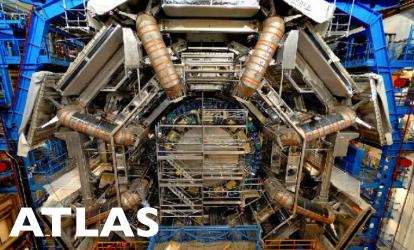




Dark matter searches

- Identify DM is one of the most important in physics.
 DM is likely to be (direct-)undetected particles.
- Laboratory production of DM particles





Collider

Dark Matter-nucleus scattering COUPP

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Observe DM annihilation products

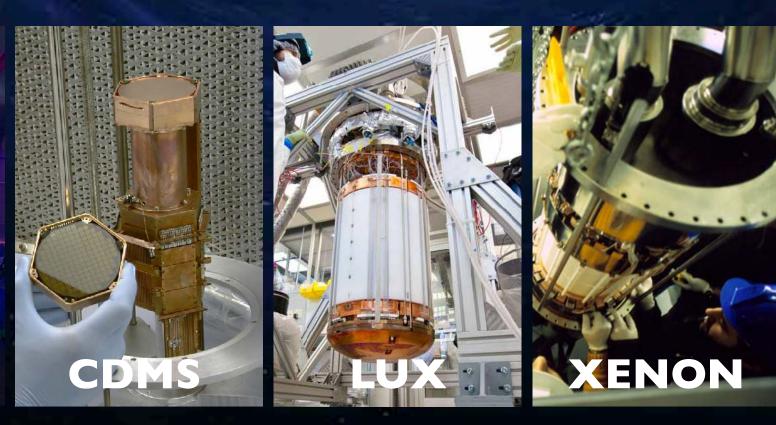






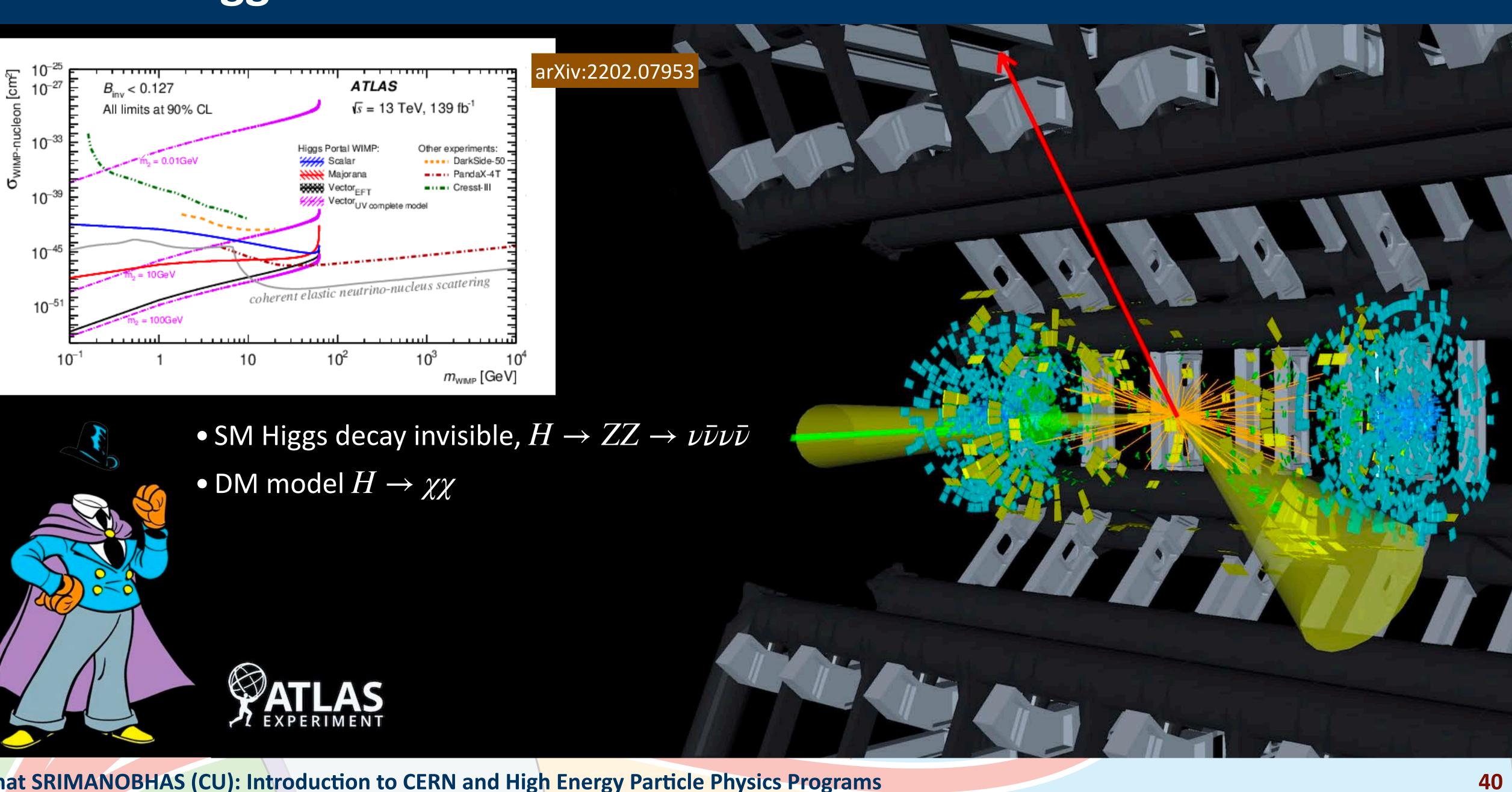
Indirect detection

Direct detection





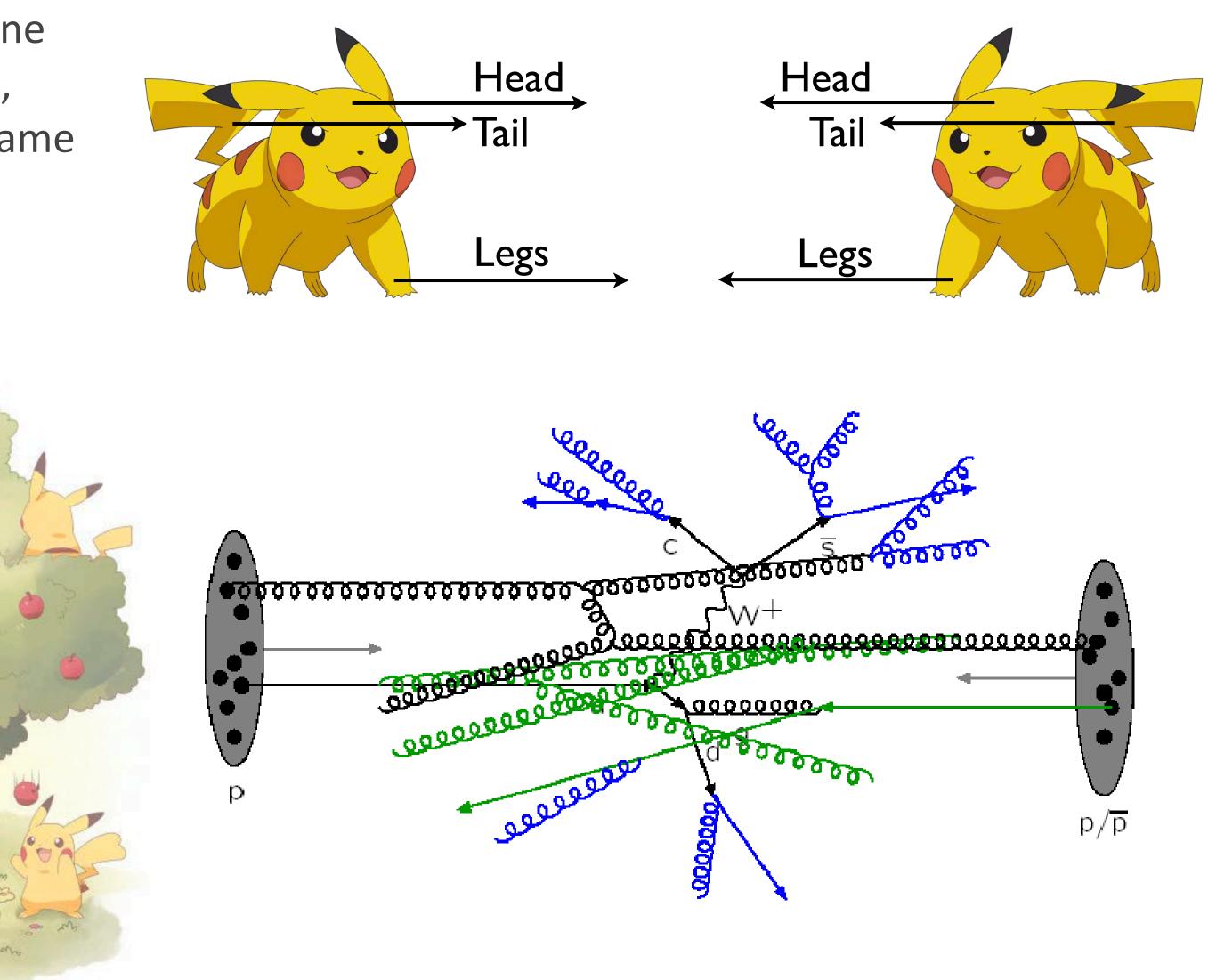
Invisible Higgs





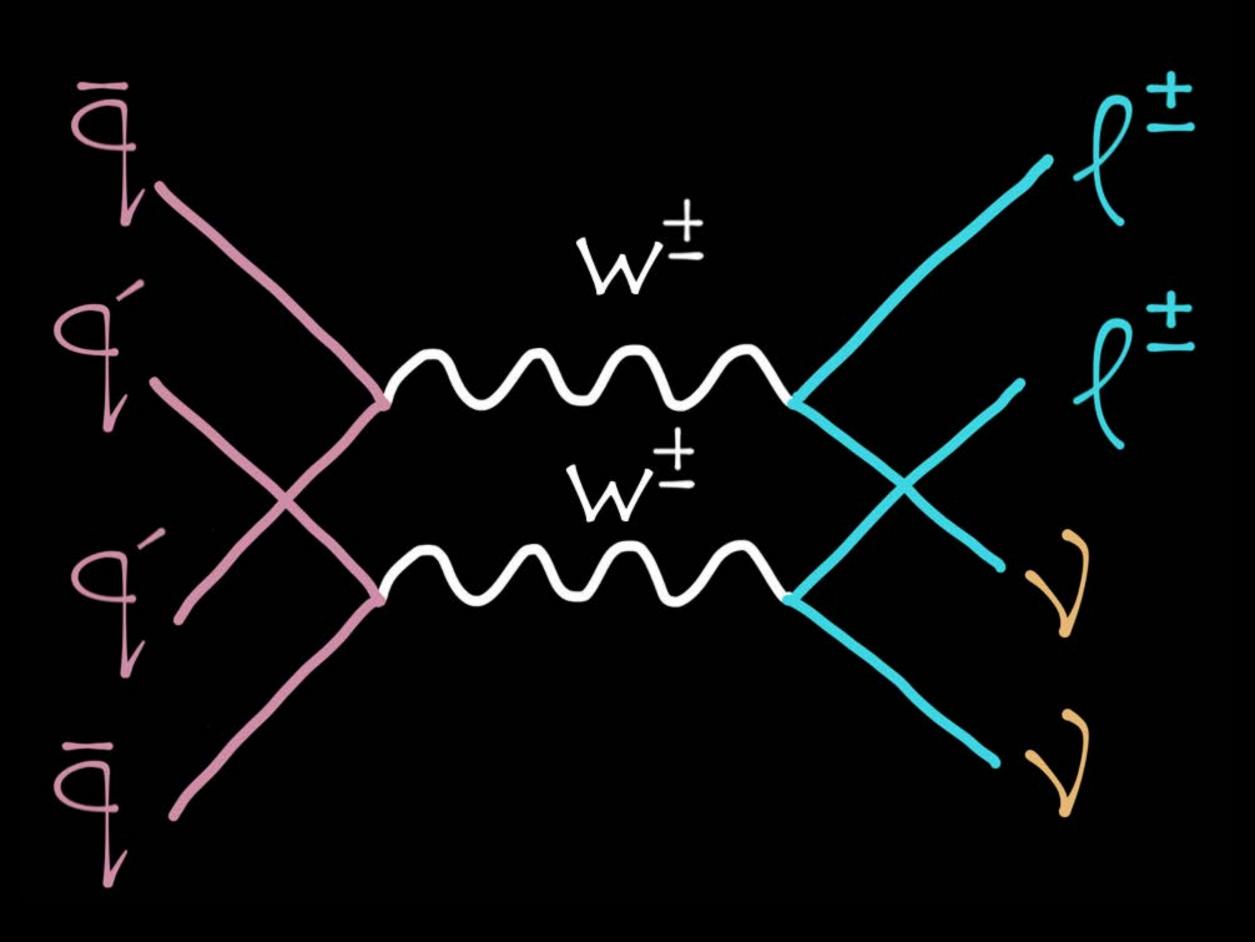
Proton collision

When "You" watch Pikachu fighting, you watch one-by-one interaction, i.e. Head-vs-Head, Tail-vs-Head, Head-vs-Leg, and then you are looking for the final result. This is the same case as proton interaction.





Double parton scattering



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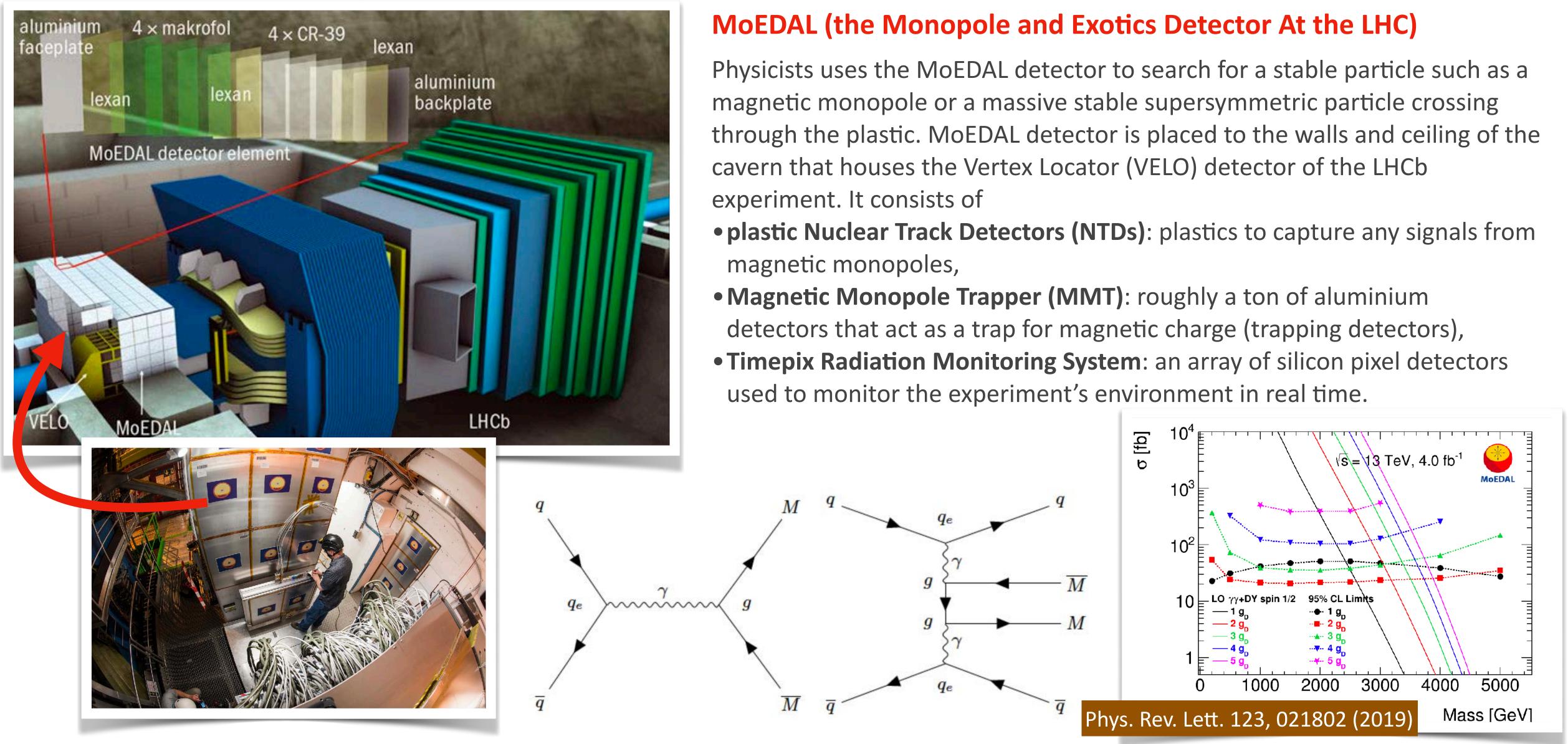
CMS Experiment at the LHC, CERN Data recorded, 2017-Nov-10 10:58:32 136704 GMT Run / Event / LS: 306459 / 2221501824 / 2004







Small experiments but big Physics potentials at the LHC

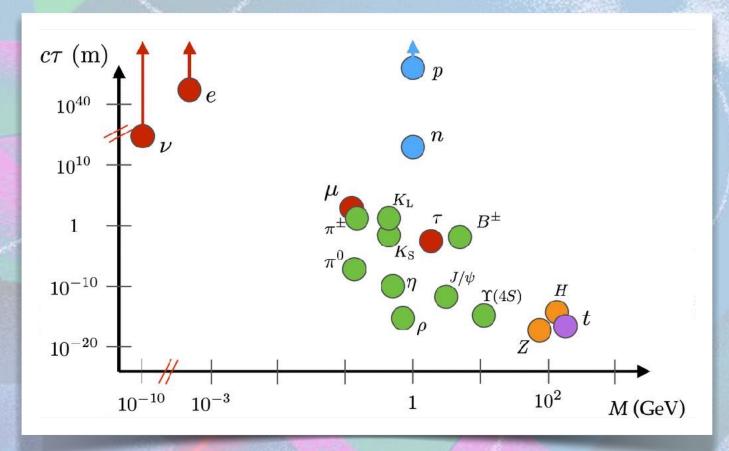




Exotic searches: Long-lived particles

When produced particles will not decay immediately ...

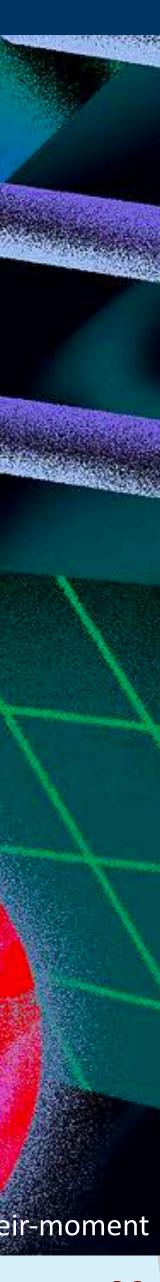
In high energy particle physics experiments, physicists often assumes that new massive particles produced in particle collisions would decay immediately, closed to their points of origin, e.g. Higgs boson. However, we also know that there are particles which have long lifetimes, e.g. muons which can travel several kilometers (with the help of special relativity) before transforming into electrons and neutrinos.



What if new particles we are hunting for has long lifetimes and traveled centimeters—even kilometers -before transforming into something physicists could detect?

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https://www.symmetrymagazine.org/article/long-lived-particles-get-their-moment



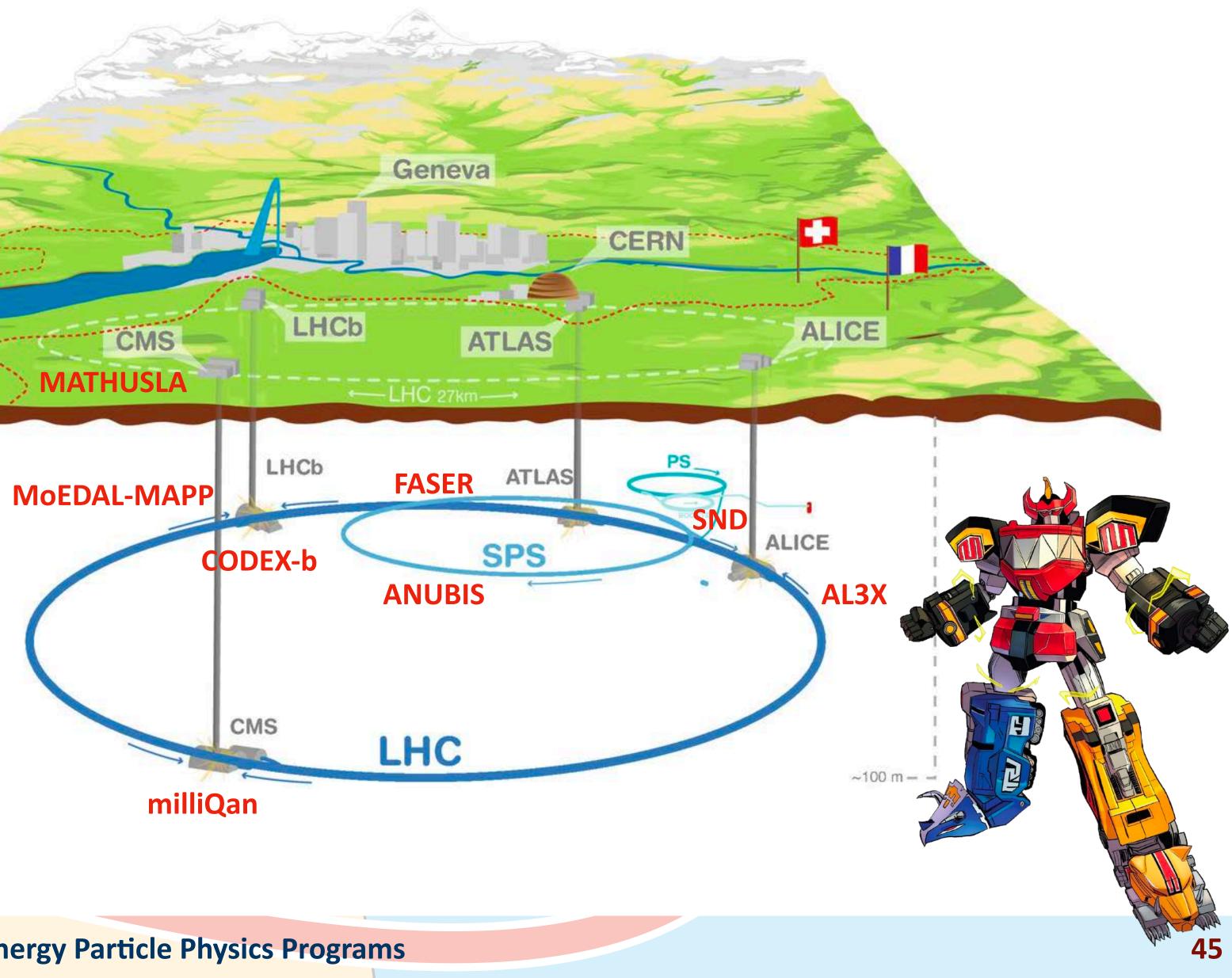


Power Detector Rangers: To search for Long-lived particles

Several small experiments around LHC to search for LLPs

- FASER (Approved): with a volume of ~1 m³ will be installed 480 m downstream from the ATLAS interaction point
- MATHUSLA: large scale surface detector instrumenting ~8×10⁵ m³ above ATLAS or CMS
- **CODEX-b**: ~10³ m³ detector to be installed in the LHCb cavern
- AL3X to use a cylindrical ~900 m³ detector inside the L3 magnet and the time-projection chamber of the ALICE experiment
- MilliQan to search for millicharged particles in the drainage gallery of CMS
- **MoEDAL** to look for highly ionizing particles like magnetic monopoles at LHCb alongside **MAPP**

.... Why do need need several experiments?



Power Detector Rangers: To search for Long-lived particles

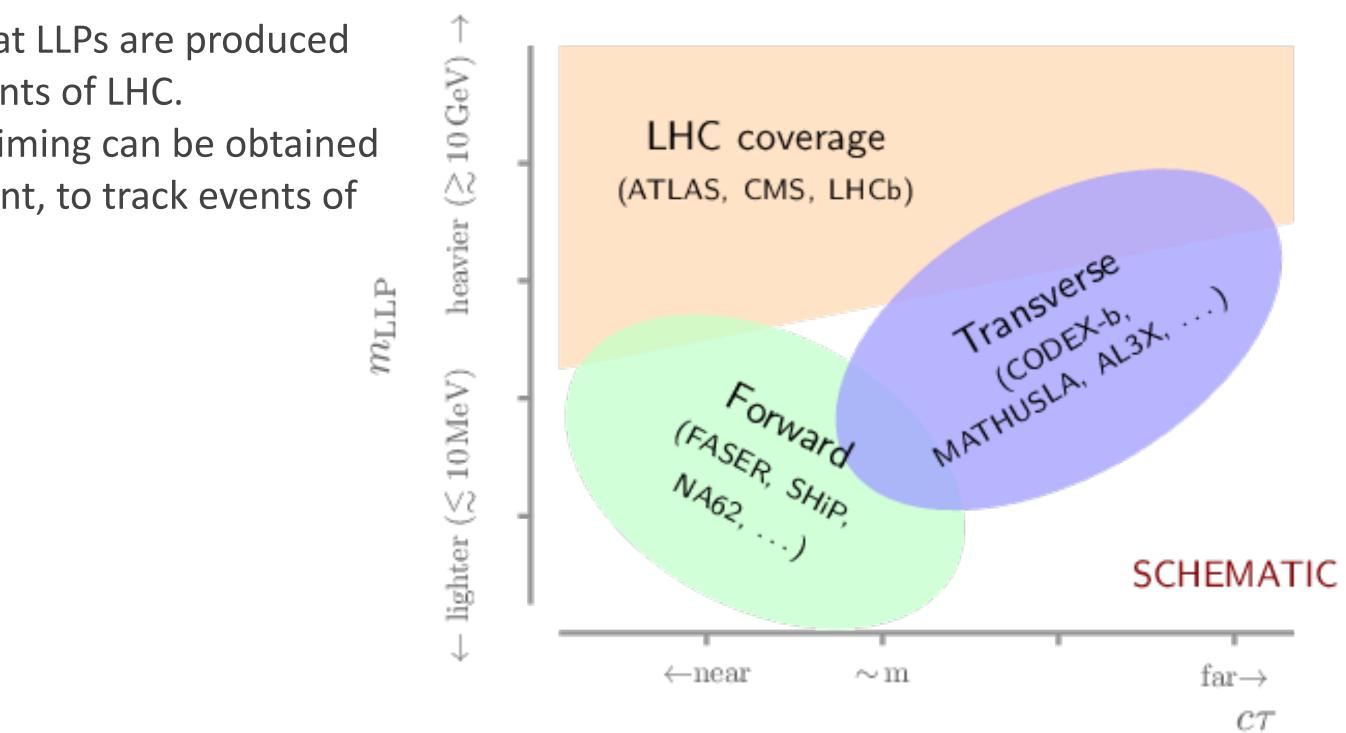


Why do we need detector rangers to search for LLP?

their masses ... don't know their lifetimes ... don't know

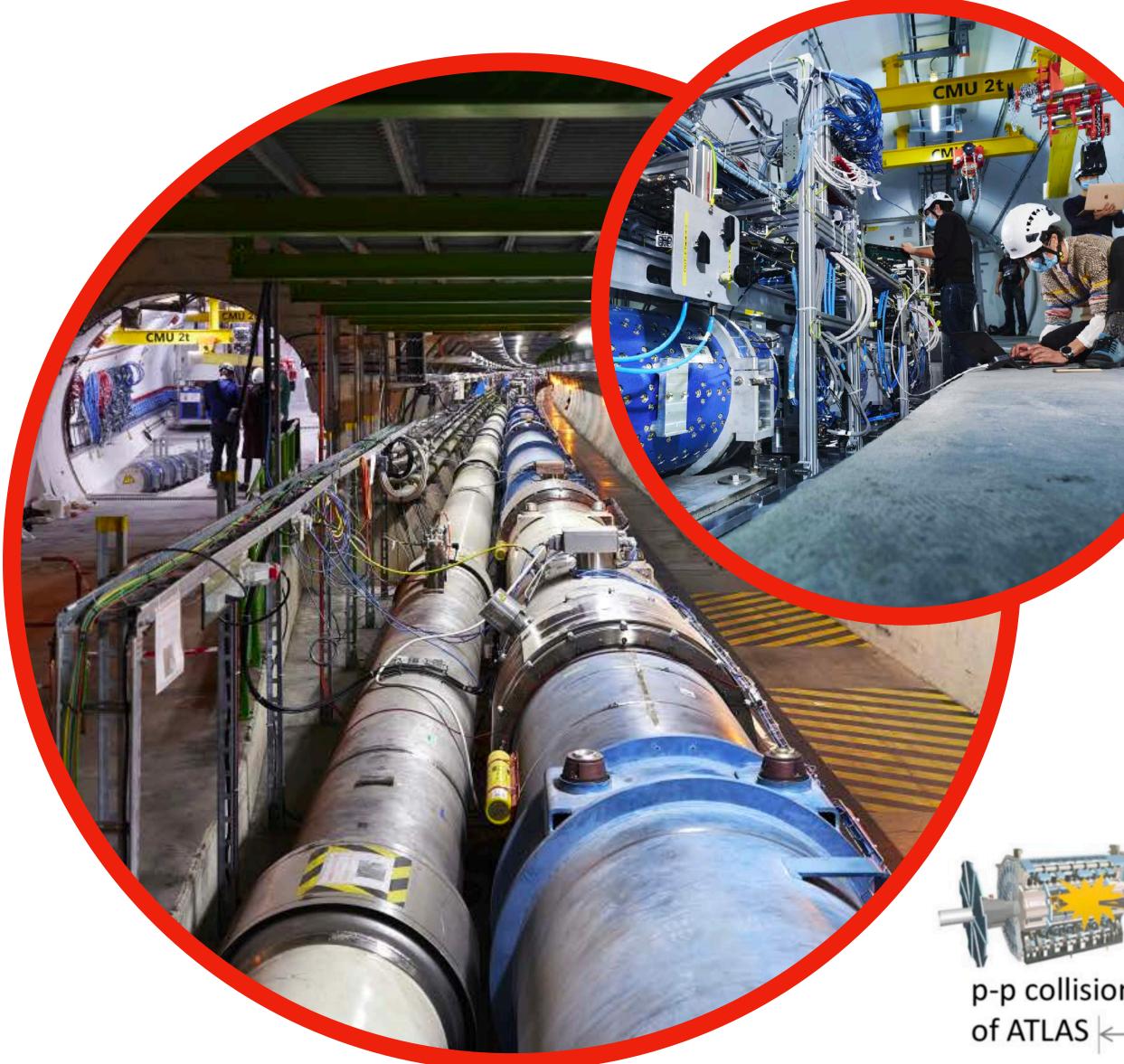
Properties of LLPs may span many orders of magnitude. This makes us impossible from first principles to construct a single detector which would have the ultimate sensitivity to all possible LLP signatures. Multiple complementary experiments are necessary. *However, we still need to connect* to main-big detector. Why?

Because we assume that LLPs are produced at the four collision points of LHC. Information including timing can be obtained from big four experiment, to track events of interest.





ForwArd Search ExpeRiment (FASER)



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particles • New particles produced in decays of light mesons • Travel at ~zero angle • Escaping detection in ATLAS (for FASER)/CMS • $pp \rightarrow X + LLP$, then LLP travels for ~480m, $LLP \rightarrow e^+e^-, \mu^+\mu^-, \gamma\gamma, \dots$ • Several models including dark photon, axion-like particles (ALPs), heavy neutral leptons (HNLs), and dark Higgs bosons Scinti. Scinti. 0.5T magnet Scinti. 0.5T magnet 0.5T magnet Decaying to e⁺e⁻ pair Tracker Calorimeter Tracker Tracker LHC tunnel charged particles (P<7 TeV) forward jets neutrino, dark photon LHC magnets 100 m of rock p-p collision at IP 480 m



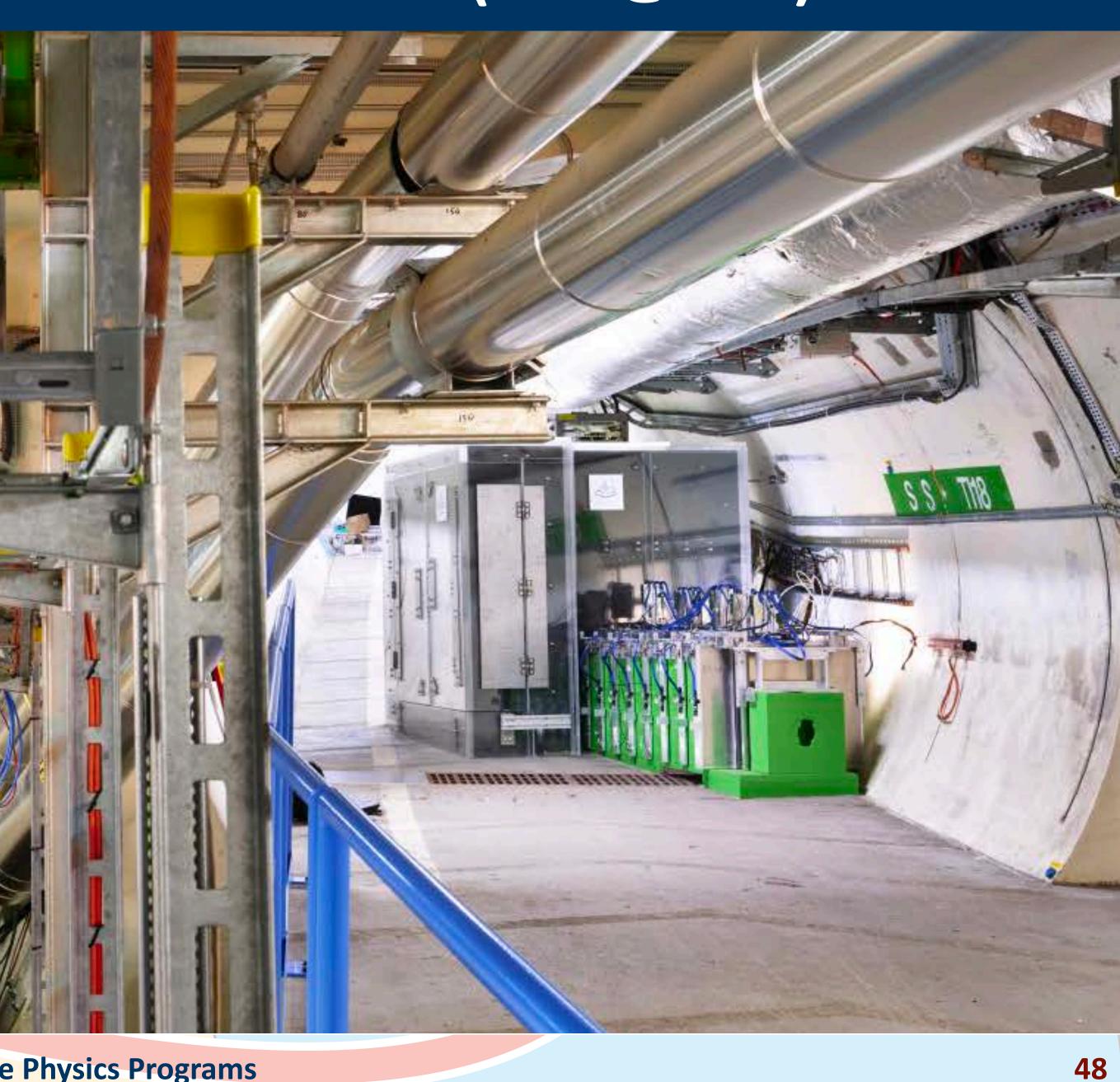




The Scattering and Neutrino Detector at the LHC (SND@LHC)

The SND@LHC is located underground close to the ATLAS experiment, in an unused tunnel that links the LHC to the Super Proton Synchrotron. Positioned slightly off the LHC's beamline, it will be able to detect neutrinos produced in the LHC collisions at small angles with respect to the beamline.

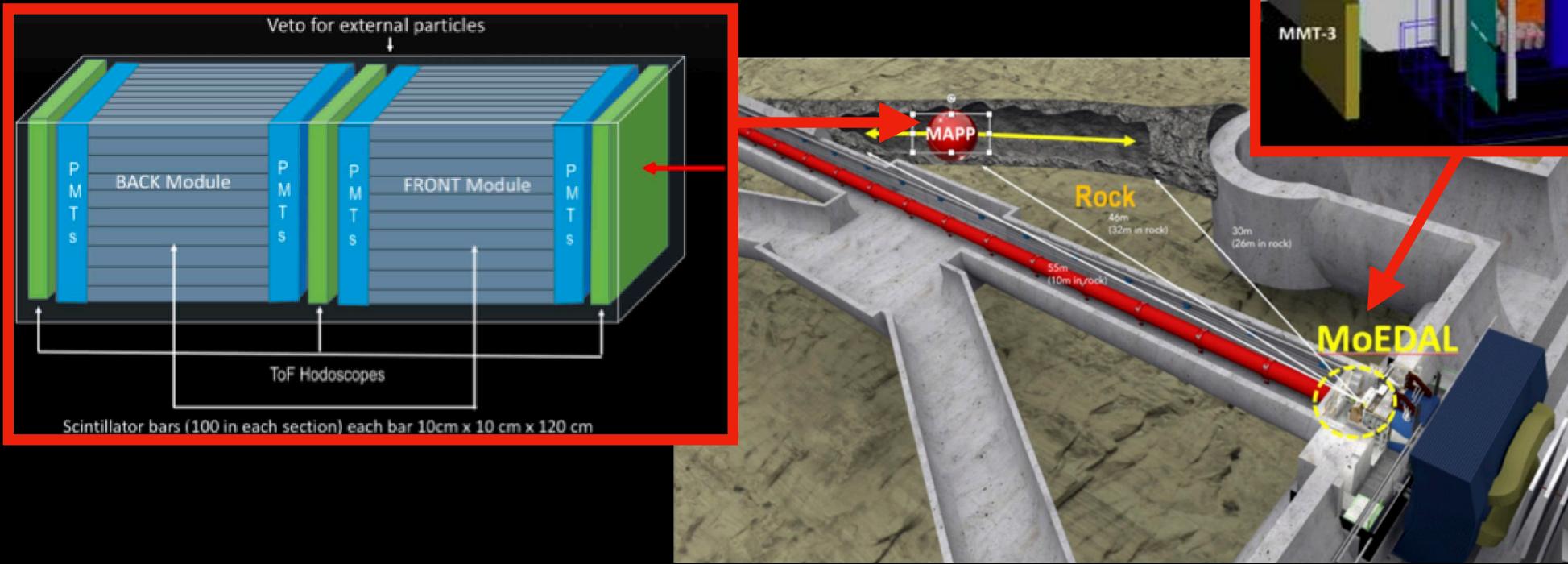
SND@LHC consists of a neutrino target followed downstream by a device to measure the neutrino energy and to detect muons (the heavier cousins of electrons) that are produced when neutrinos interact with the target.

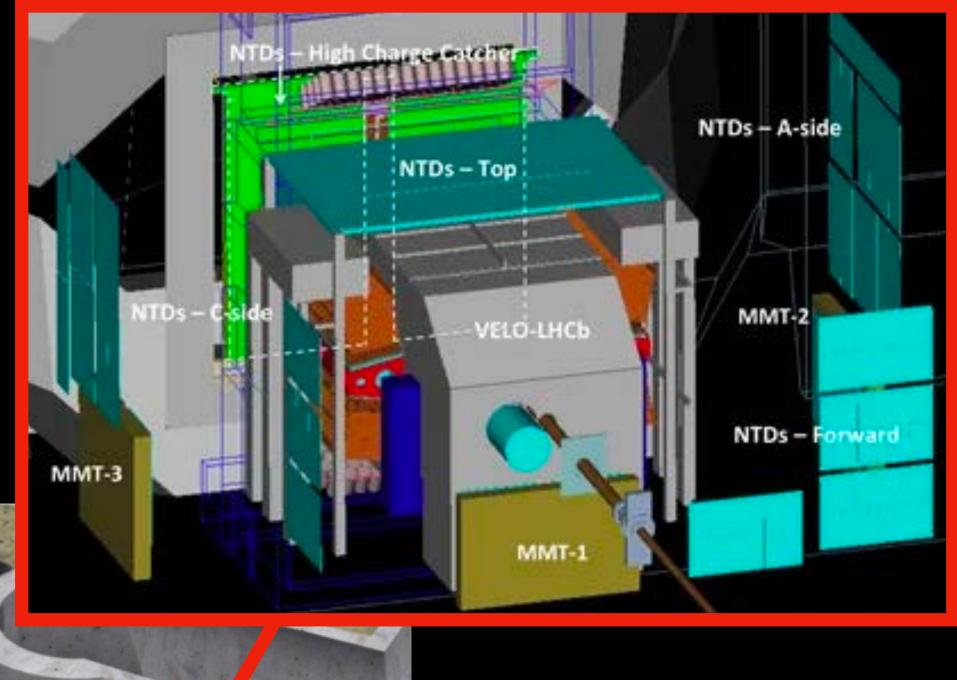


MoEDAL Apparatus for Penetrating Particles (MAPP)

Several small experiments around LHC to search for LLPs

- For Phase-1 (LHC RUN-3): The baseline MoEDAL detector will be reinstalled and two MAPP sub detectors for two class of particles:
- MAPP-LLP: nw pseudo-stable weakly interactive neutral particles with long lifetime
- MAPP-mQP: mini-charged particle detector
- Positioned at an angle of 5-degree w.r.t. beam axis





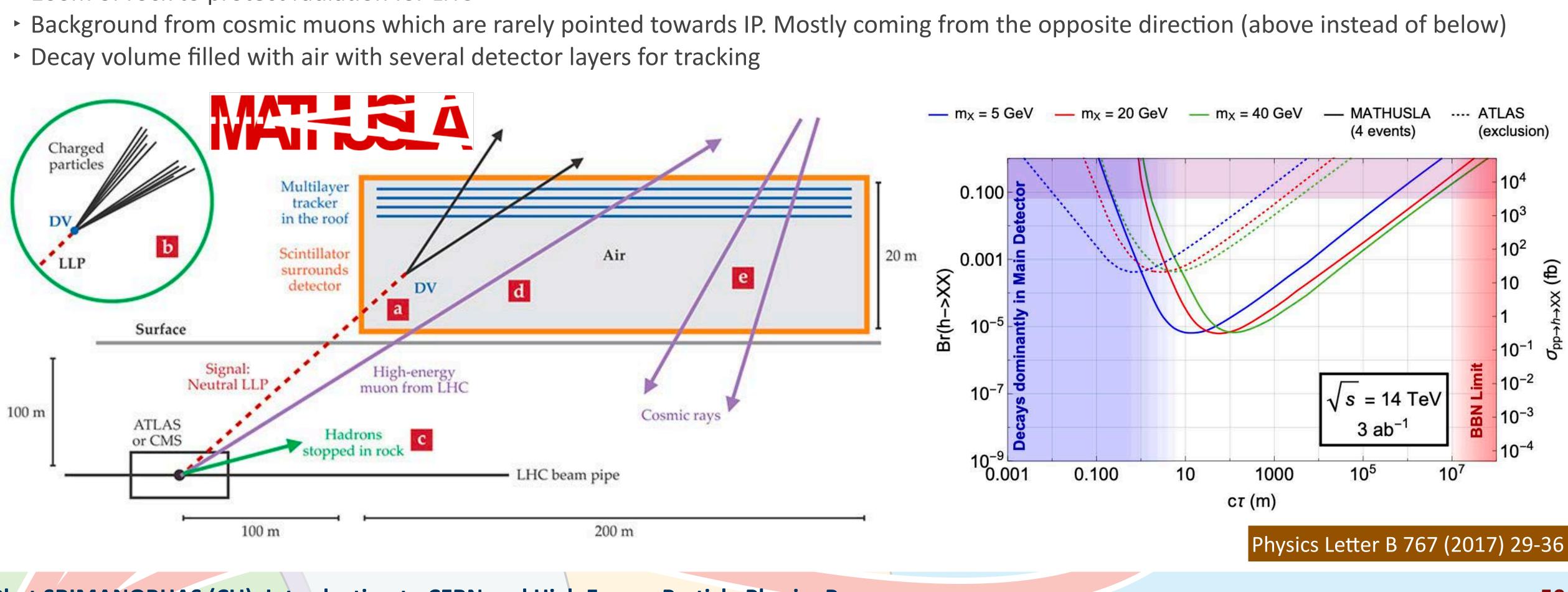


MAssive Timing Hodoscope for Ultra-Stable neutral pArticles (MATHUSLA)

Ultra-long-lived particles (ULLP), with surface detector

Requirement: New detector that minimizes background as much as possible; Maintain reasonably large radial detector size and solid angle coverage relative to the main interaction point (IP); Cheap. The proposal is to go for surface detector:

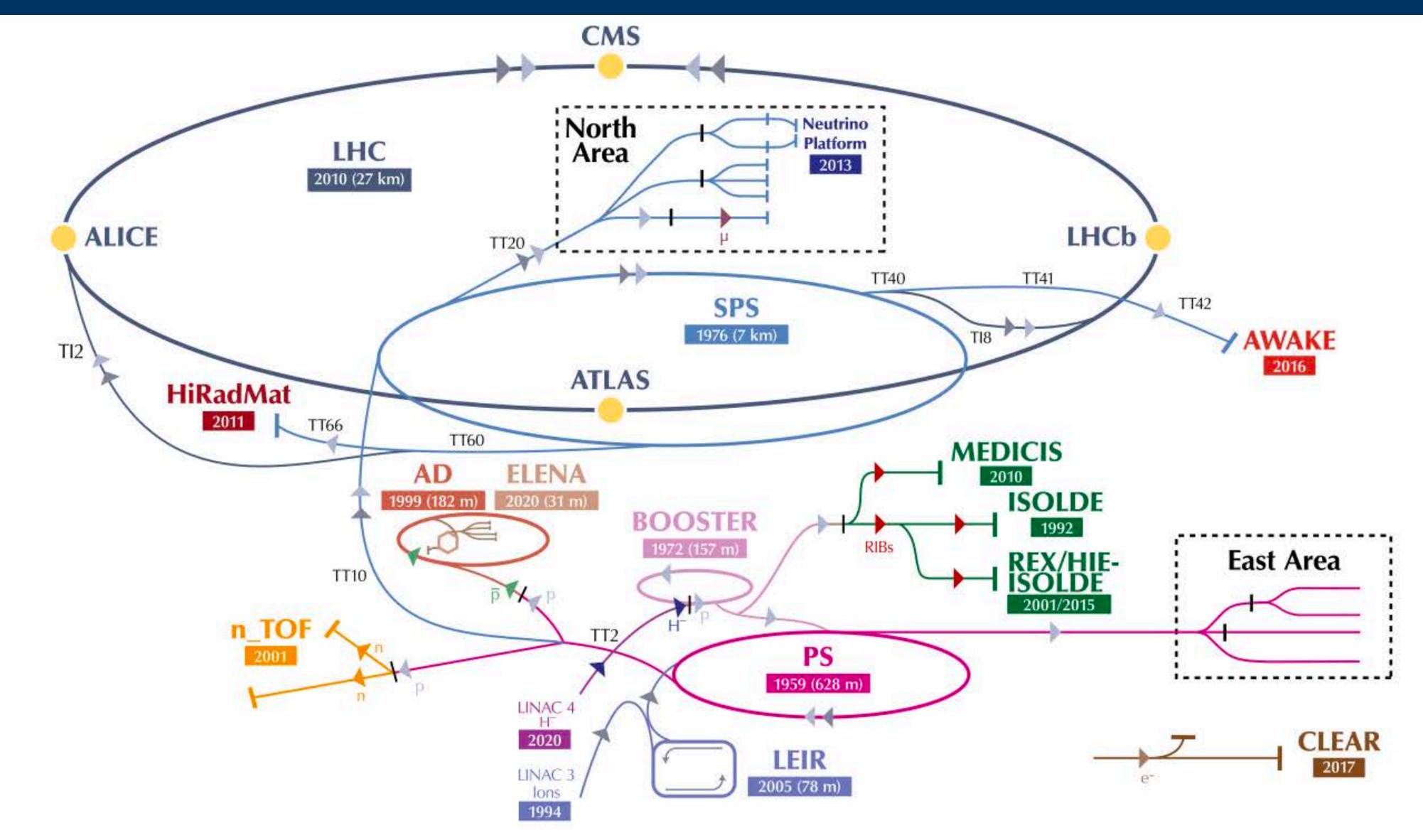
- No additional tunnel
- 100m of rock to protect radiation for LHC







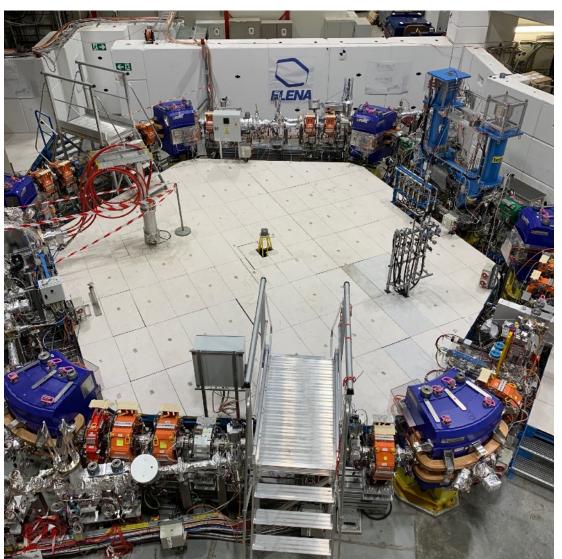
CERN accelerator complex: Not only LHC





Not only accelerate, but also decelerate



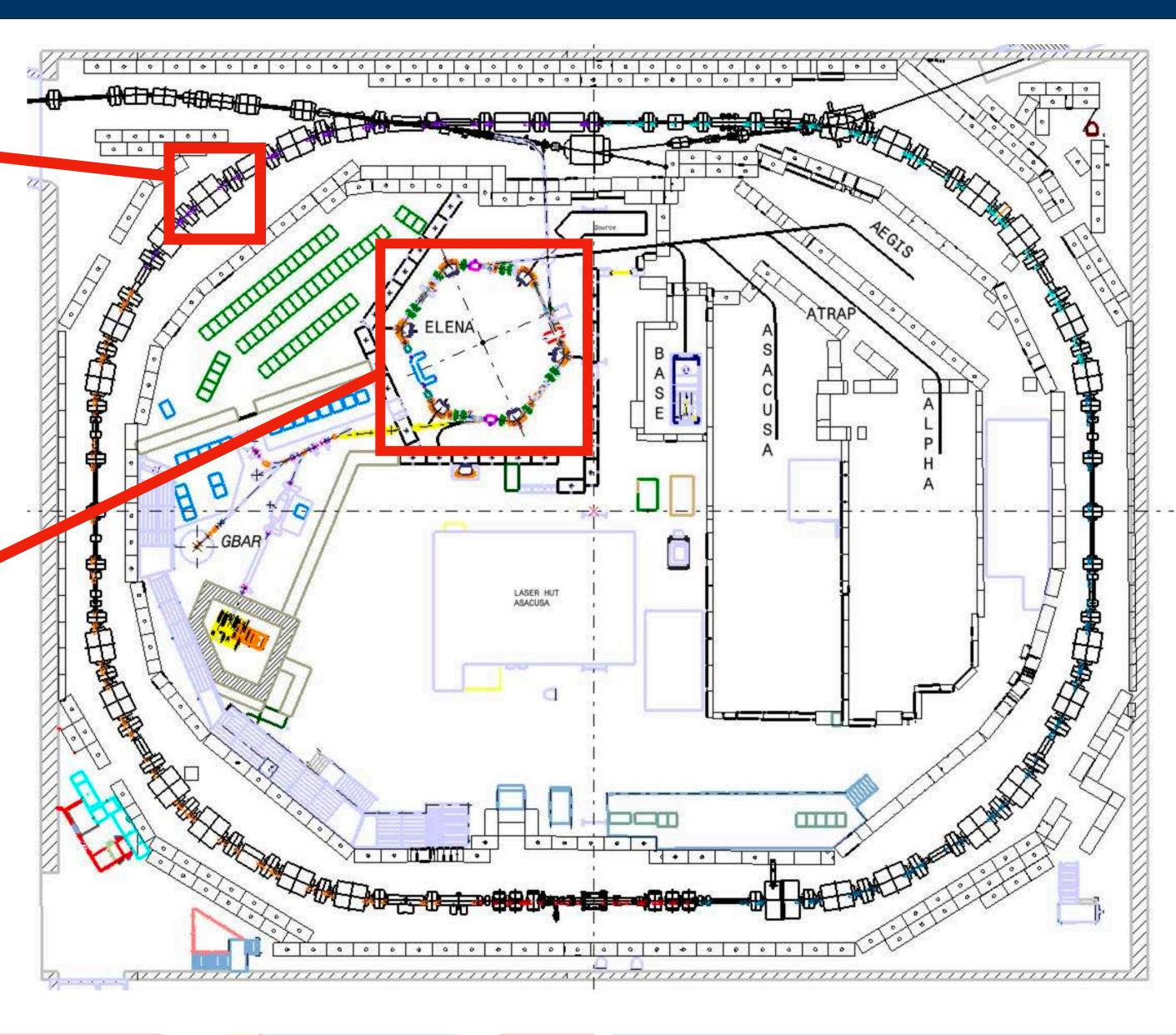


Antiproton Decelerator (AD)

A machine that produces low-energy antiprotons for studies of antimatter, and also creates antiatoms.

Extra Low ENergy **Antiproton (ELENA)**

A machine to slow more the antiprotons from AD. This is to improve the efficiency of the experiments



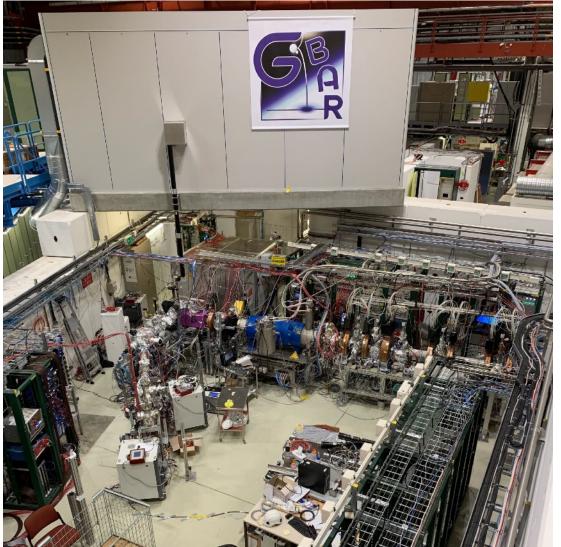


To study anti-matter



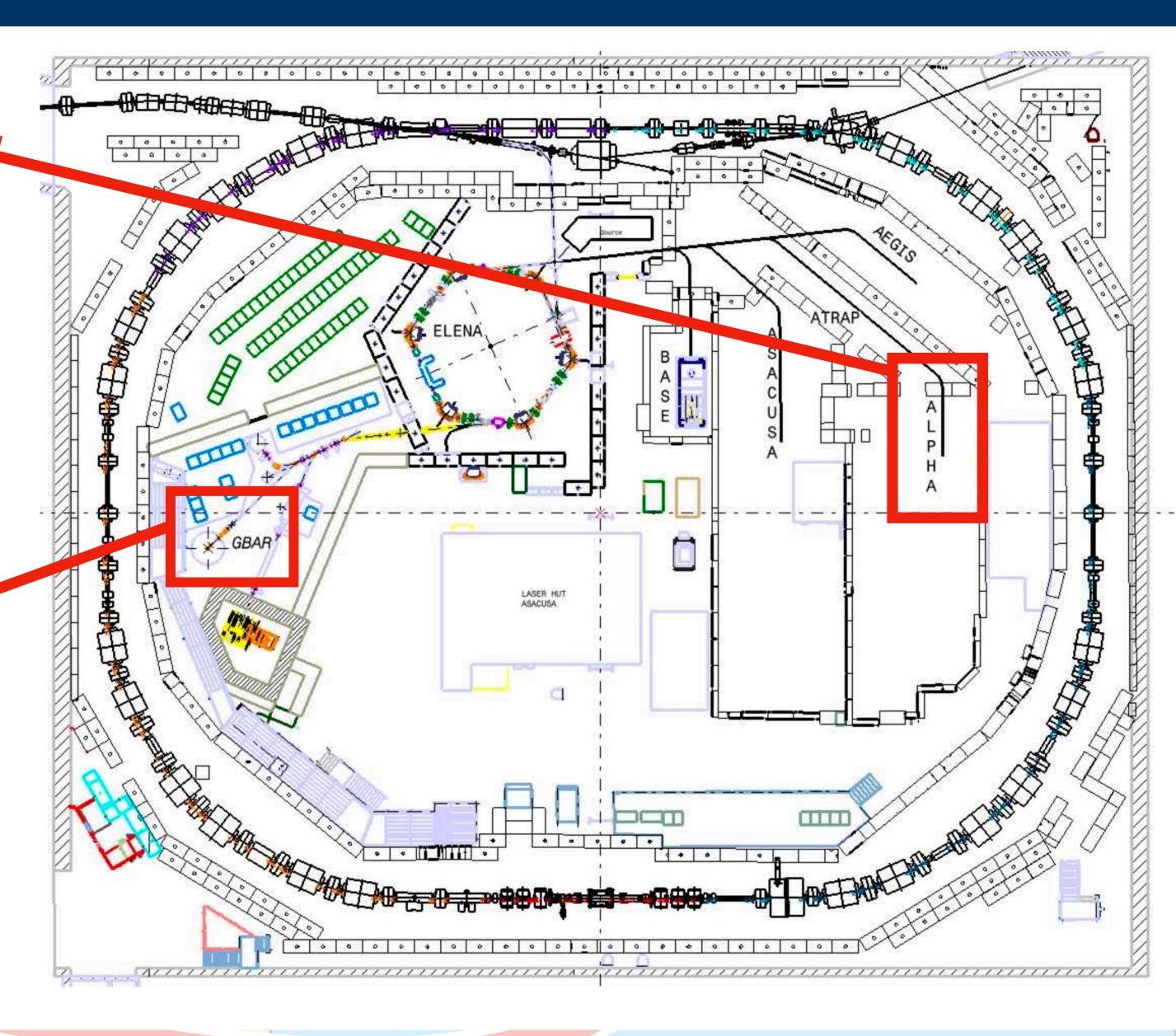
Antihydrogen Laser Physics Apparatus 🔫 (ALPHA)

create, capture and then cool antihydrogen to use for experiment



Gravitational **Behaviour of Antimatter at Rest** (GBAR)

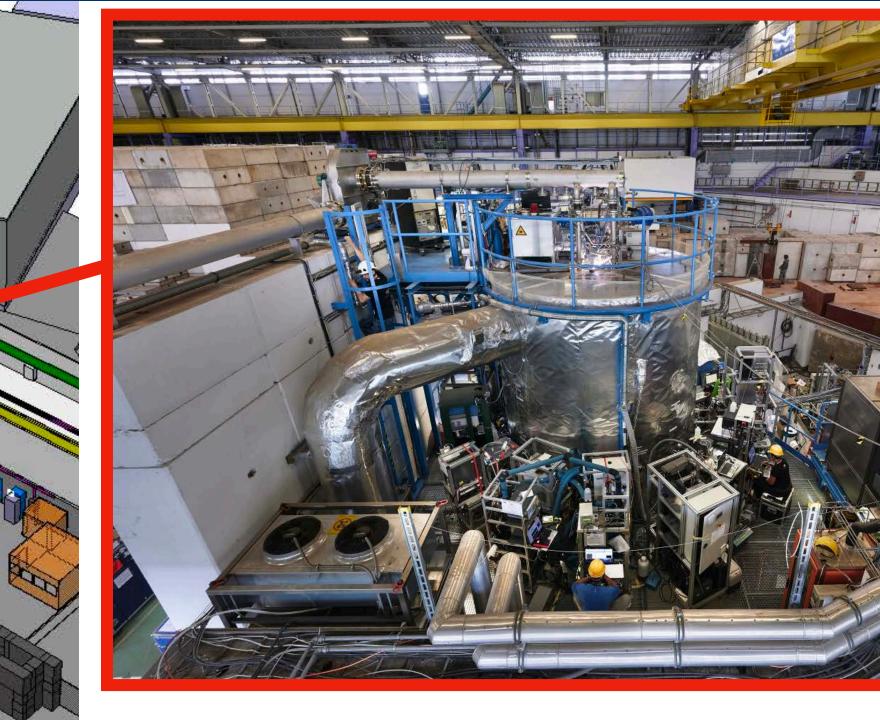
Study different behavior of hydrogen/antihydrogen under gravity (free fall)





To study links between cosmic rays and cloud formation

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CLOUD

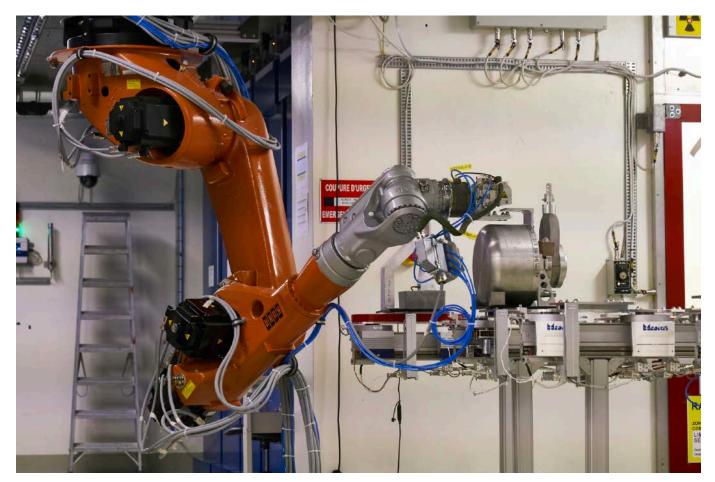
Could there be a link between galactic cosmic rays and cloud formation? An experiment at CERN is using the cleanest box in the world to find out.





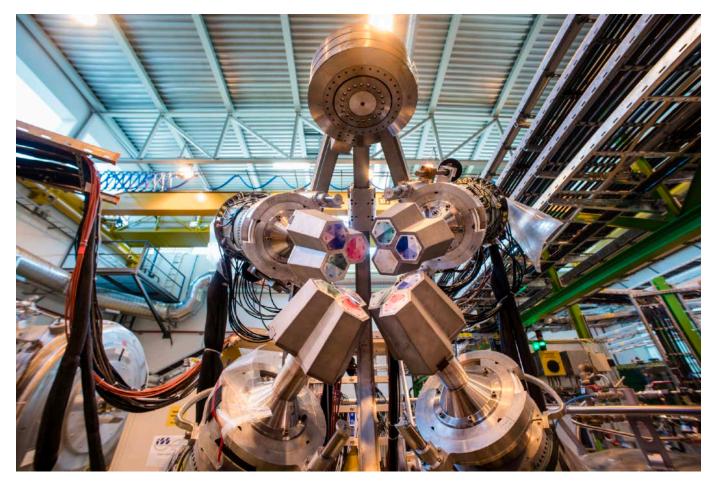


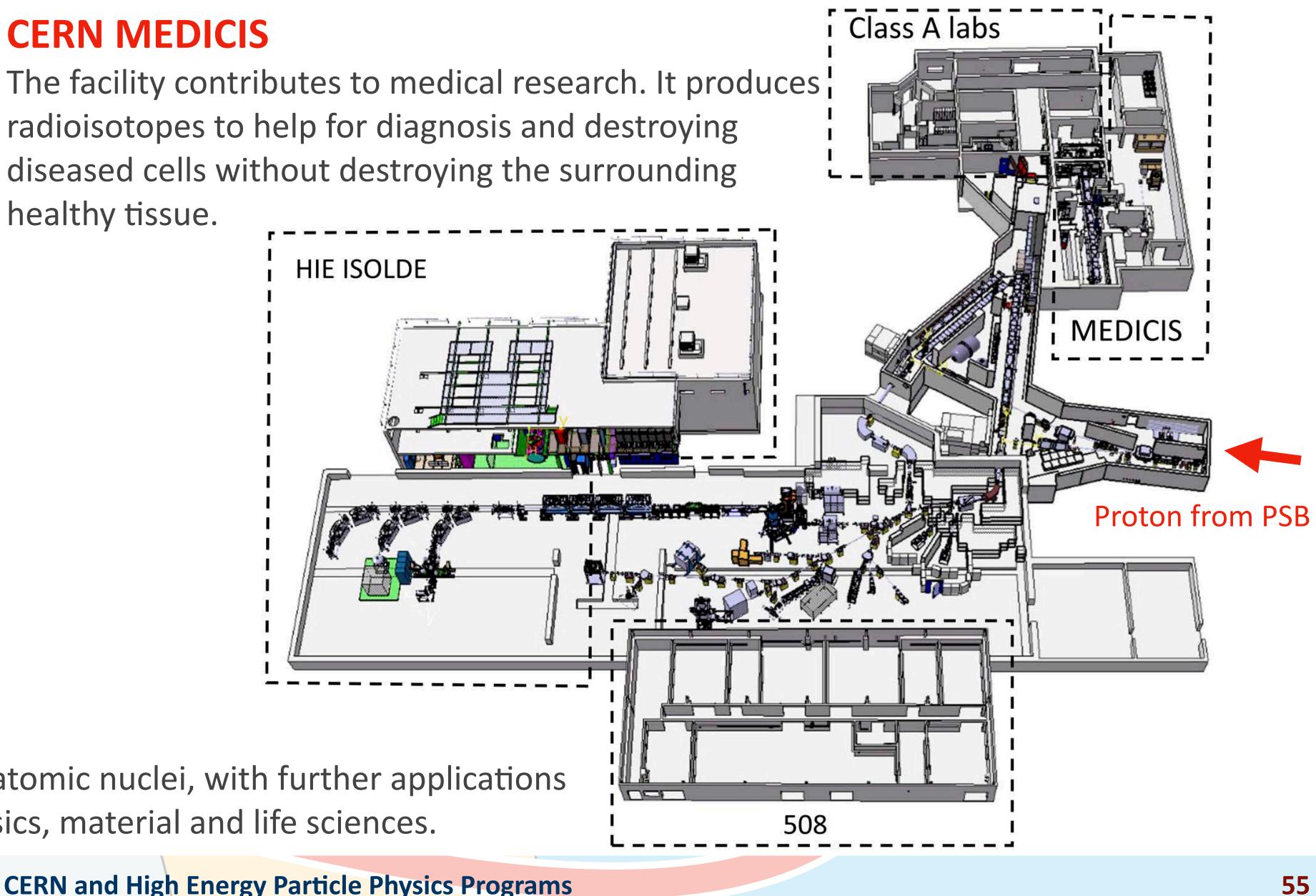
To study on radioisotopes and applications, e.g. medical applications



CERN MEDICIS

healthy tissue.





ISOLDE

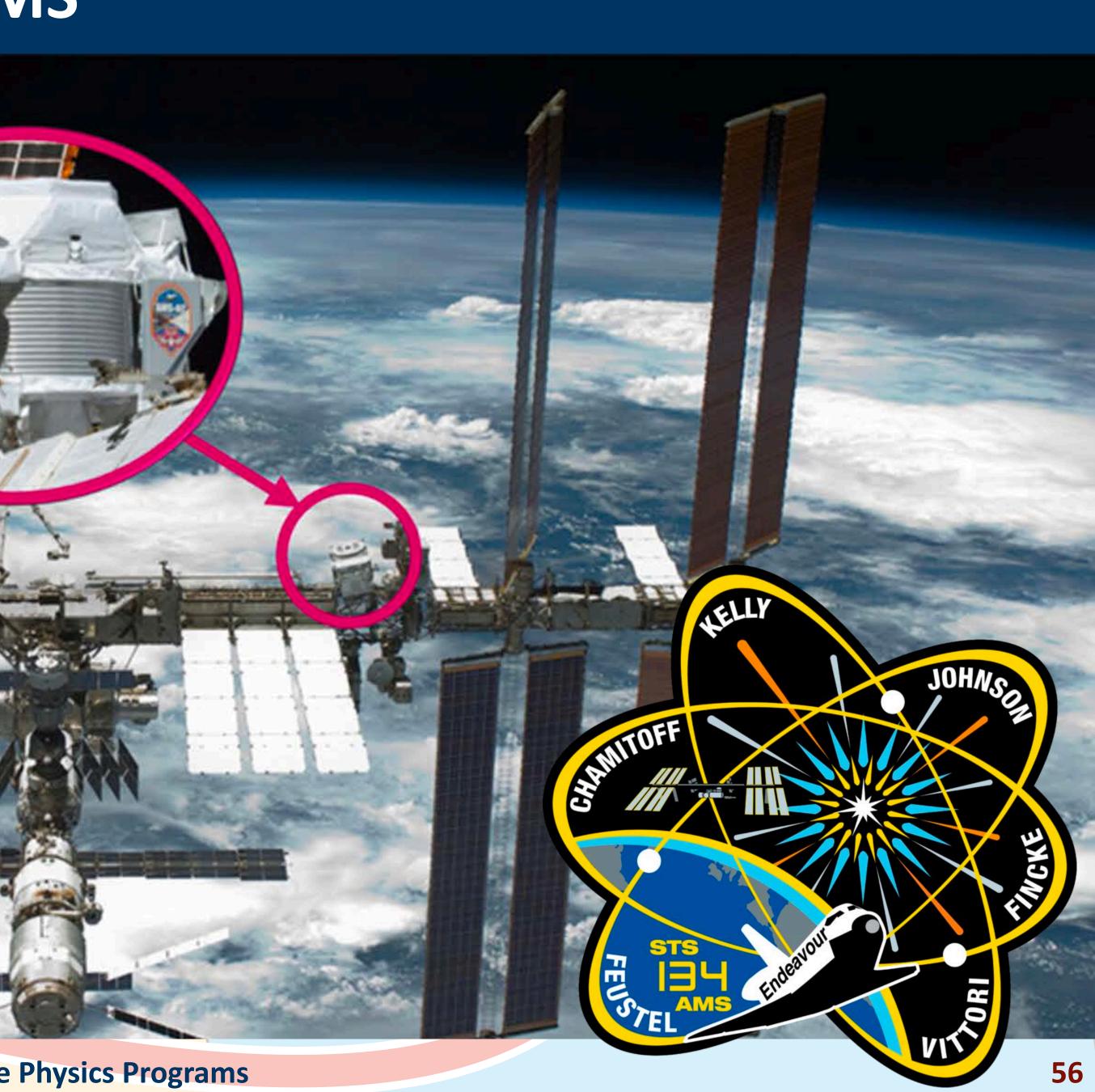
ISOLDE studies the properties of atomic nuclei, with further applications in fundamental studies, astrophysics, material and life sciences.

External experiments at CERN: AMS

AMS

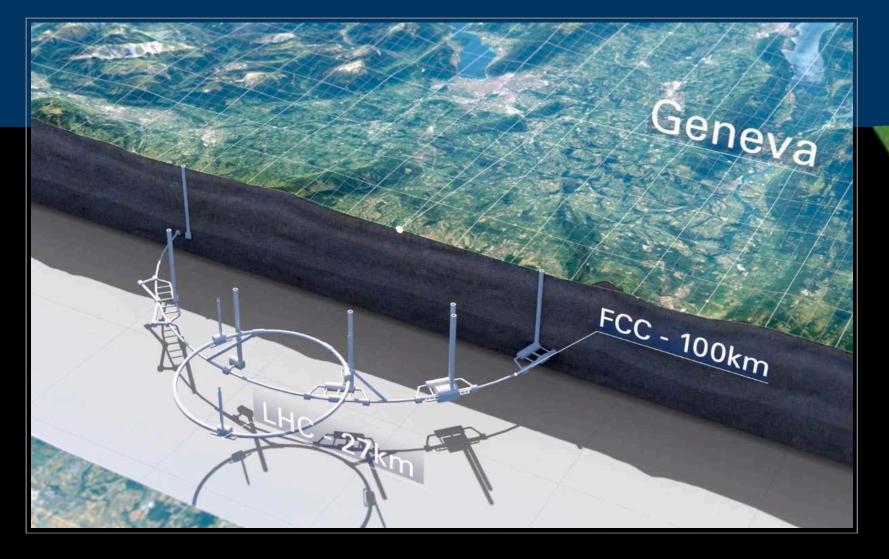
The Alpha Magnetic Spectrometer looks for dark matter, antimatter and missing matter from a module on the International Space Station.

16 May 2011: Space shuttle Endeavour delivered the AMS detector to ISS. 19 May 2011 - Now: Data from AMS is sending back to Earth - to NASA in Houston and then from NASA to CERN for analysis.



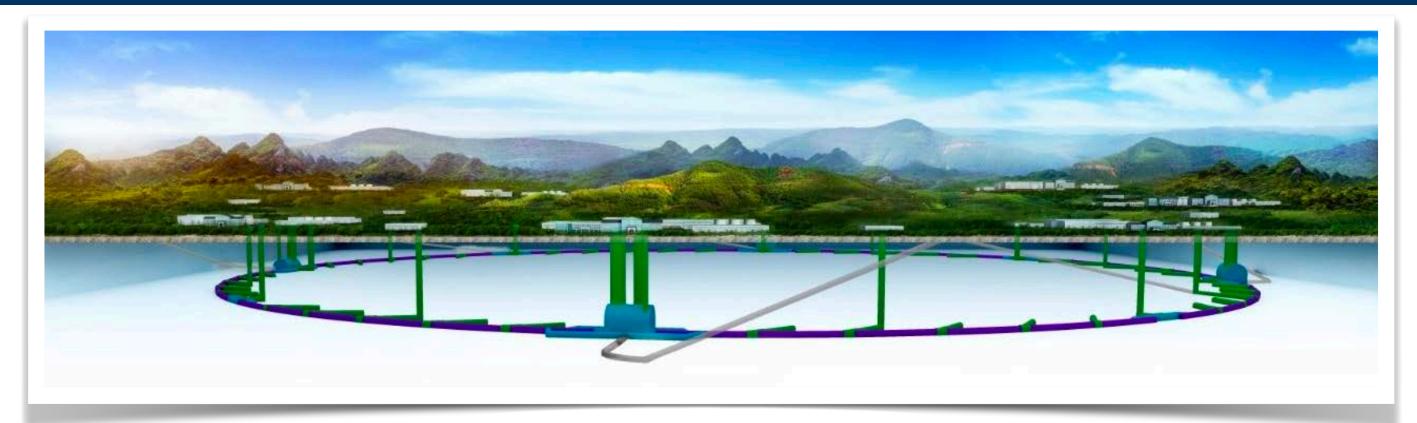
Beyond LHC

- Future Circular Collider (FCC) Circumference: 90 -100 km Energy: 100 TeV (pp) 90-350 GeV (e+e)
- Large Hadron Collider (LHC) Large Electron-Positron Collider (LEP) Circumference: 27 km Energy: 14 TeV (pp) 209 GeV (e+e)
 - Tevatron Circumference: 6.2 km Energy: 2 TeV (pp)



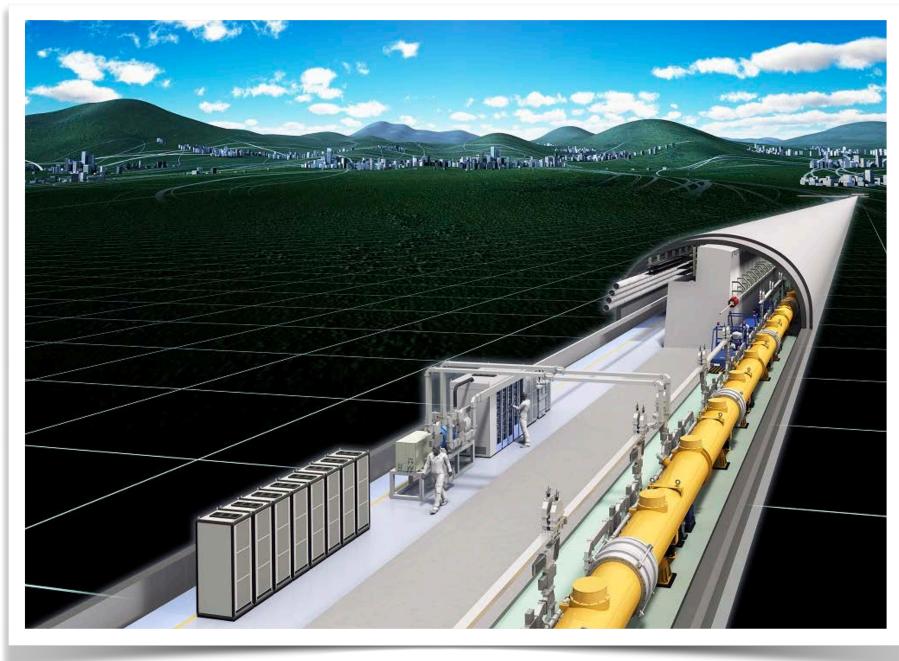


Beyond LHC: Precision Measurement



CEPC (Circular Electron Positron Collider), China

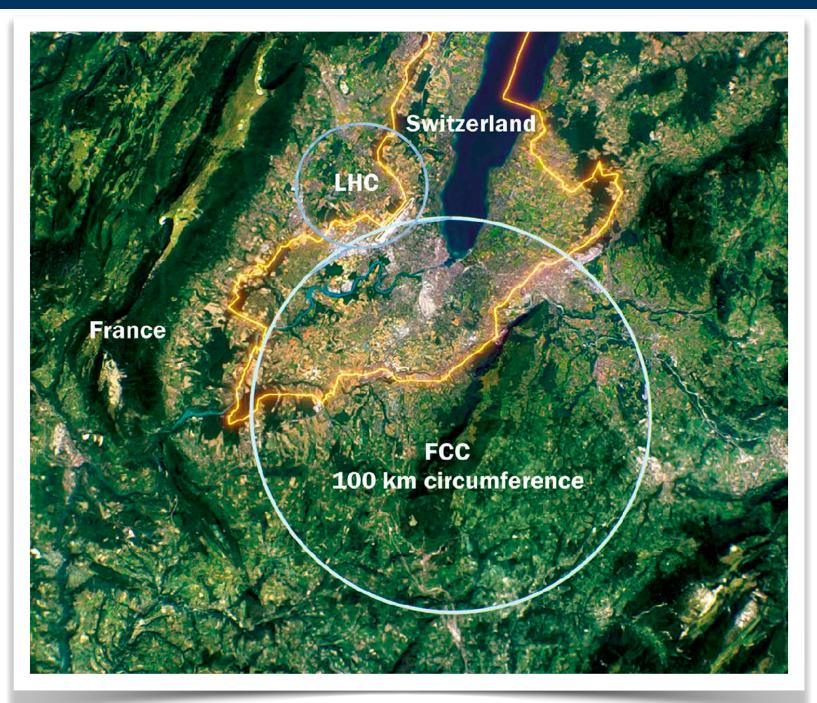
International Linear Collider (ILC)



Compact Linear Collider (CLIC)



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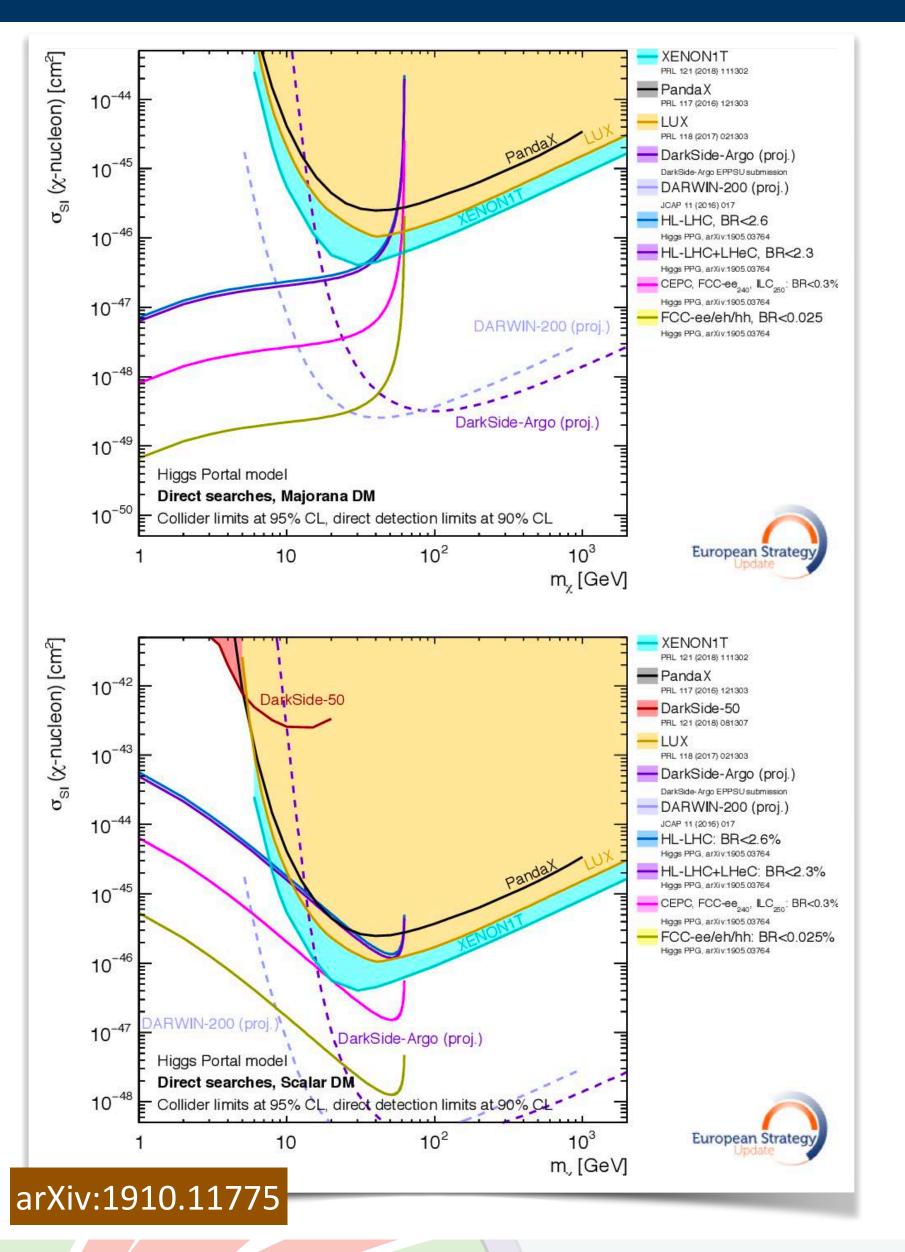


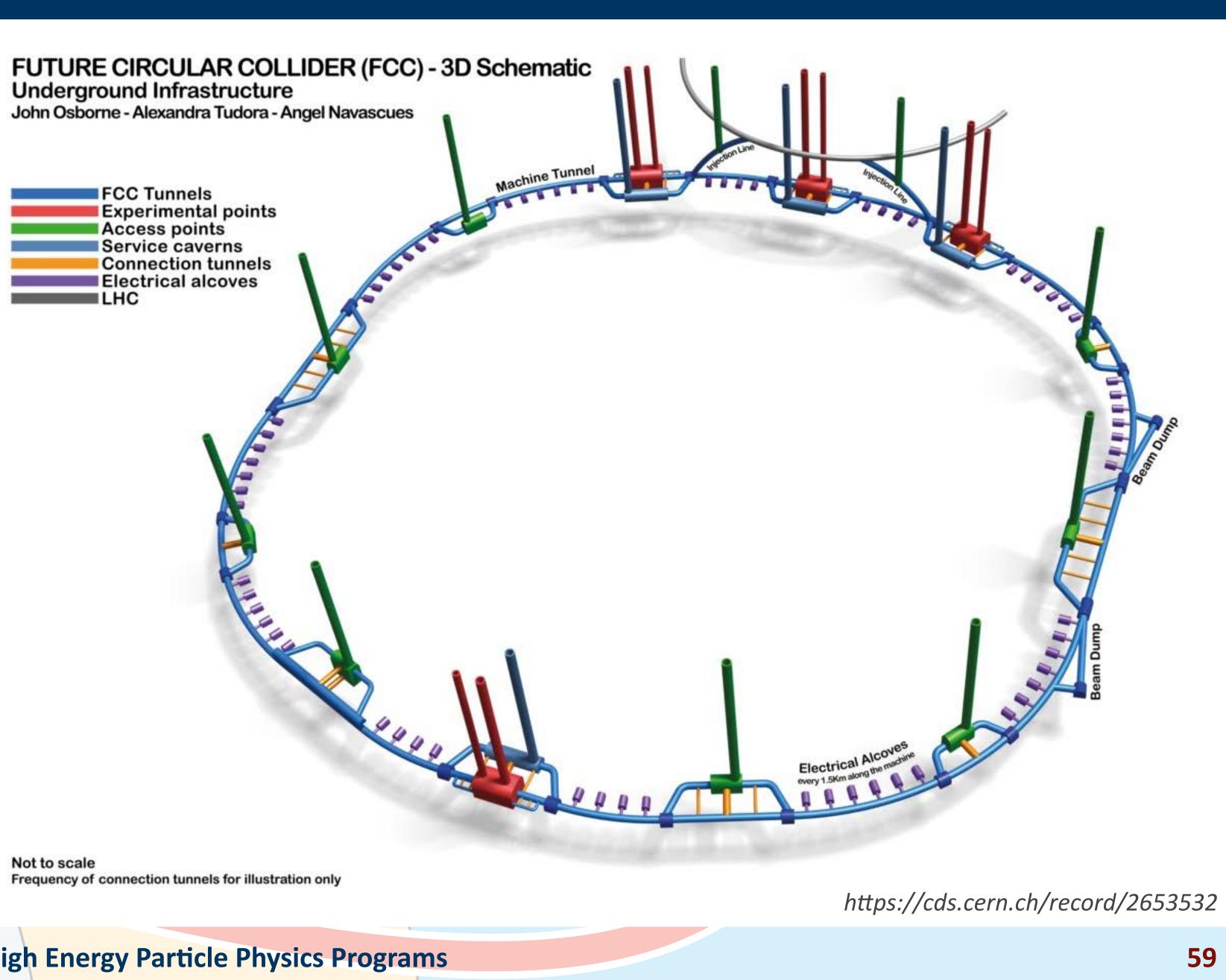
Future Circular Collider (FCC) - ee, **Switzerland-France**

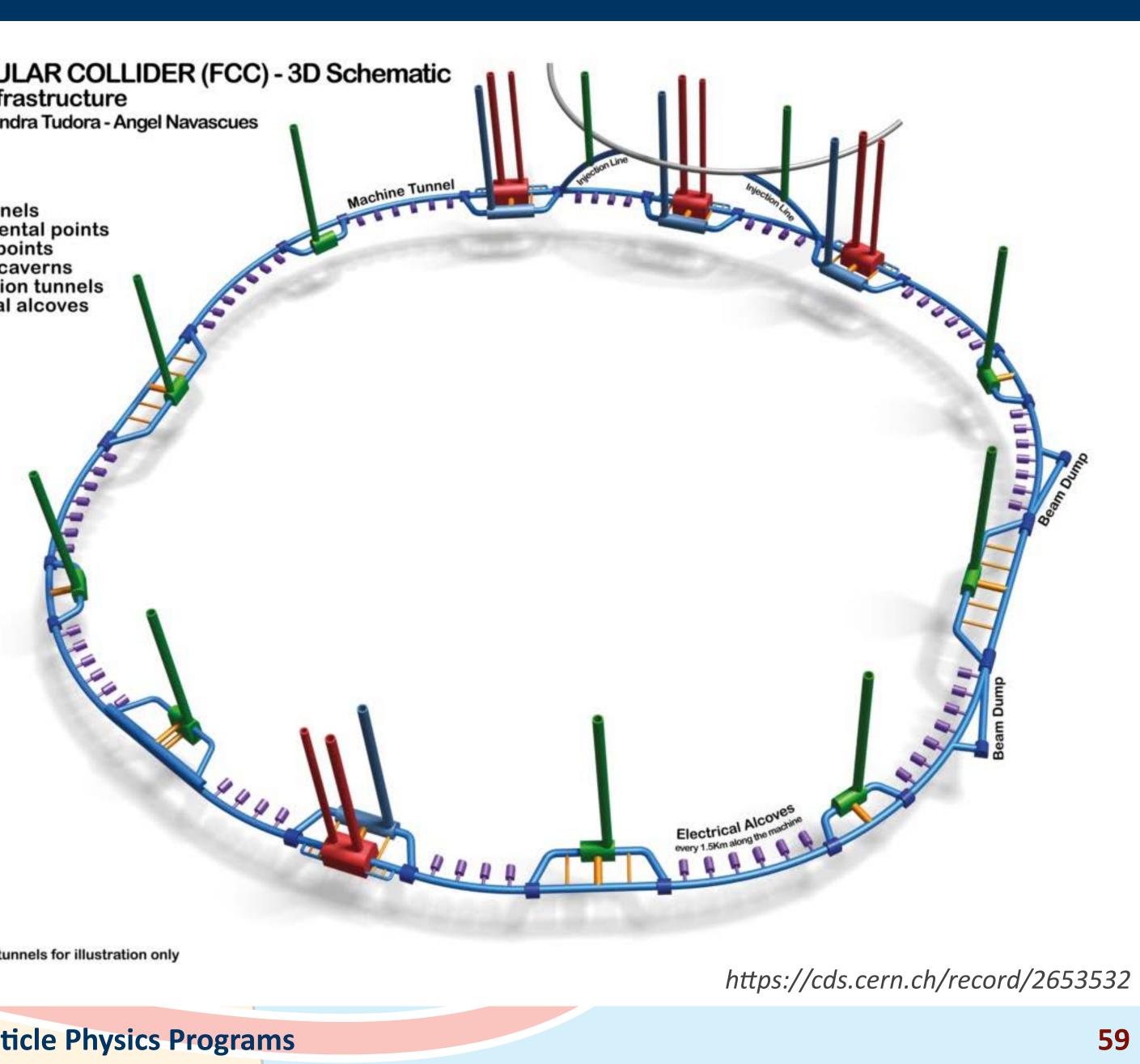




Beyond LHC: 100 TeV pp Collider







Beyond LHC: Muon Collider

Why muon collider?

As electron, muon is elementary particle (AFAWK) with higher mass (x206 electron mass). The mass of muon is the key of high energy collision.

Consider the power emitted from charged particles which are accelerated in a curved path (known as synchrotron radiation), it is proportional to $\frac{1}{m^4}$. The amount of synchrotron radiation from a muon will be reduced by a factor of about 1 billion of an electron.

$$P = \frac{q^2}{6\pi\epsilon_0 m^4 c^5 r^2 \sin^2(\alpha)} (E^2 - m^2 c^4)^2$$

Something to solve?

We need to handle the muon's lifetime. At rest, it will decay in 2 μ s. If we can accelerate a muon close to the speed of light before decaying, its lifetime will stretch longer. This is the result of the special relativity.

Muon Collider **Conceptual Layout**

Project X Accelerate hydrogen ions to 8 GeV using SRF technology.

Target Collisions lead to muons with energy of about 200 MeV.

In a dozen turns, accelerate muons to 20 GeV.

In a number of turns, accelerate muons up to 2 TeV using SRF technology.

Bring positive and negative muons into collision at two locations 100 meters underground.

Compressor Ring Reduce size of beam.

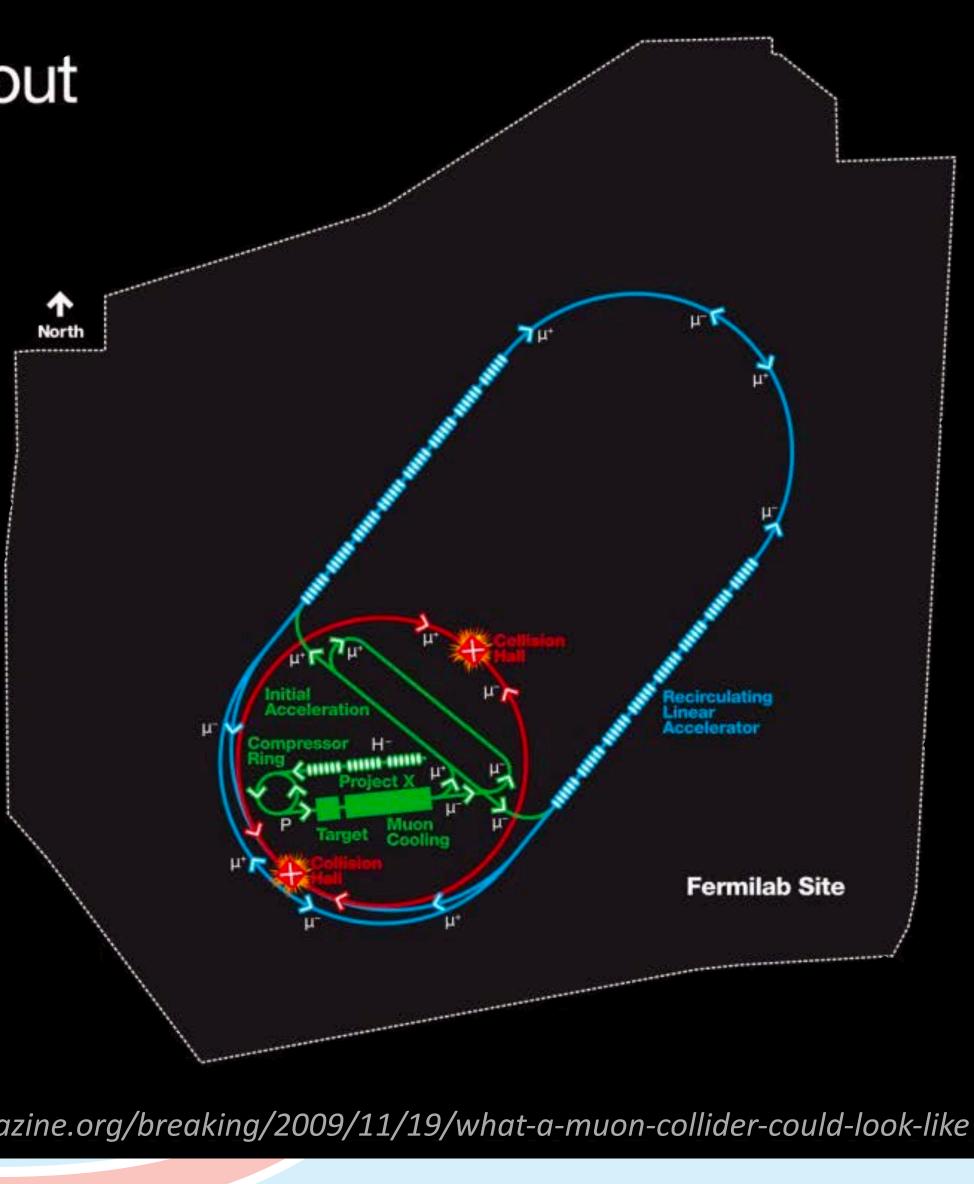
Muon Capture and Cooling

Capture, bunch and cool muons to create a tight beam.

Initial Acceleration

Recirculating Linear Accelerator

Collider Ring



60

https://www.symmetrymagazine.org/breaking/2009/11/19/what-a-muon-collider-could-look-like

Why particle physics matters



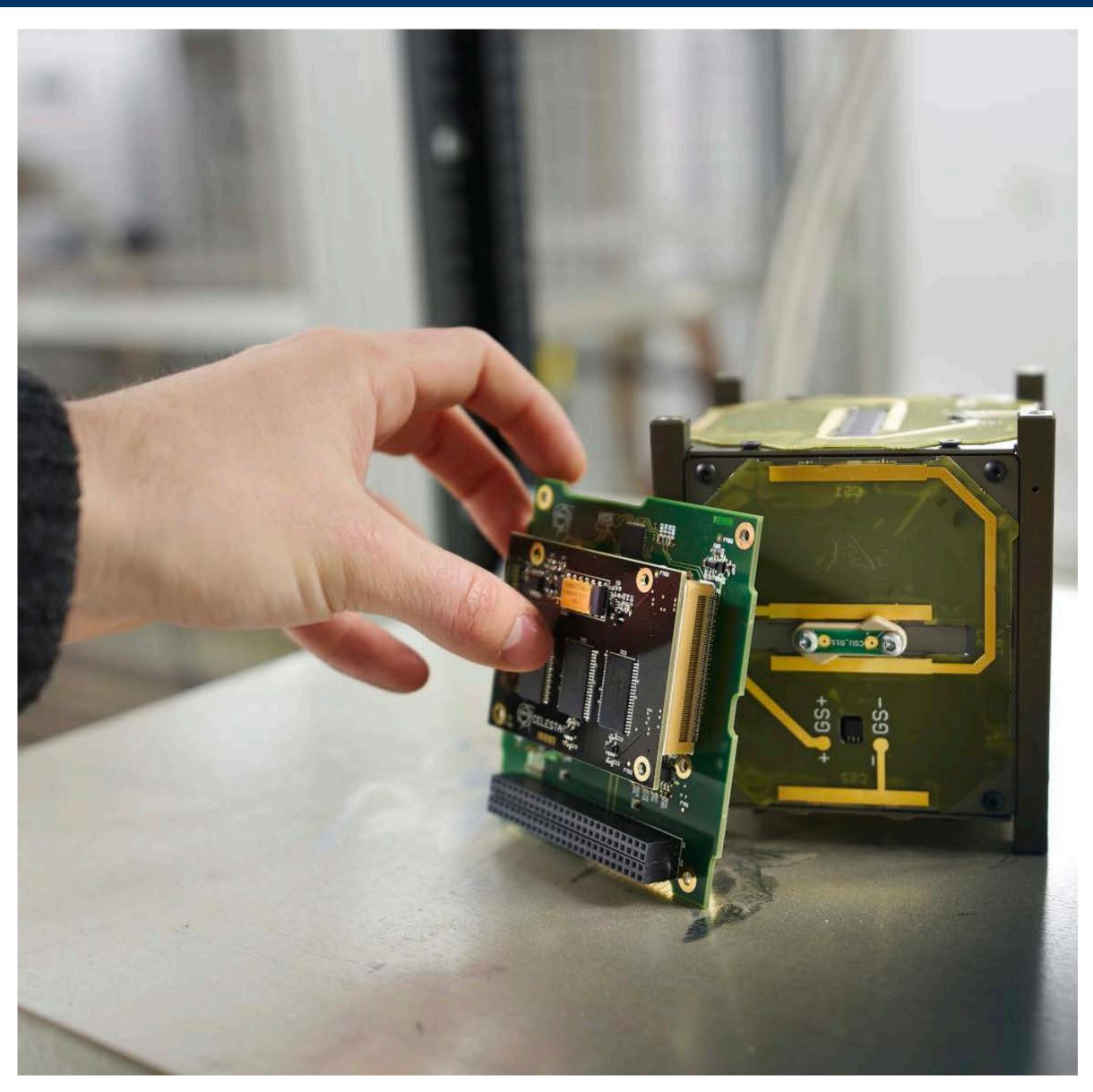
Phat SRIMANOBHAS (CU): Introduction to CERN and High Energy Particle Physics Programs

http://www.symmetrymagazine.org/article/october-2013/why-particle-physics-matters





CERN latchup and radmon experiment student satellite



Phat SRIMANOBHAS (CU): Introduction to CERN and High Energy Particle Physics Programs

Ref: <u>here</u>

CELESTA (CERN latchup and radmon experiment student satellite) is a 1U CubeSat with weight of one kilogram and measuring 10 cm on each of its sides. It is designed to study the effects of cosmic radiation on electronics. The satellite carries a Space RadMon, a miniature version of a well-proven radiation monitoring device deployed in CERN's Large Hadron Collider (LHC).

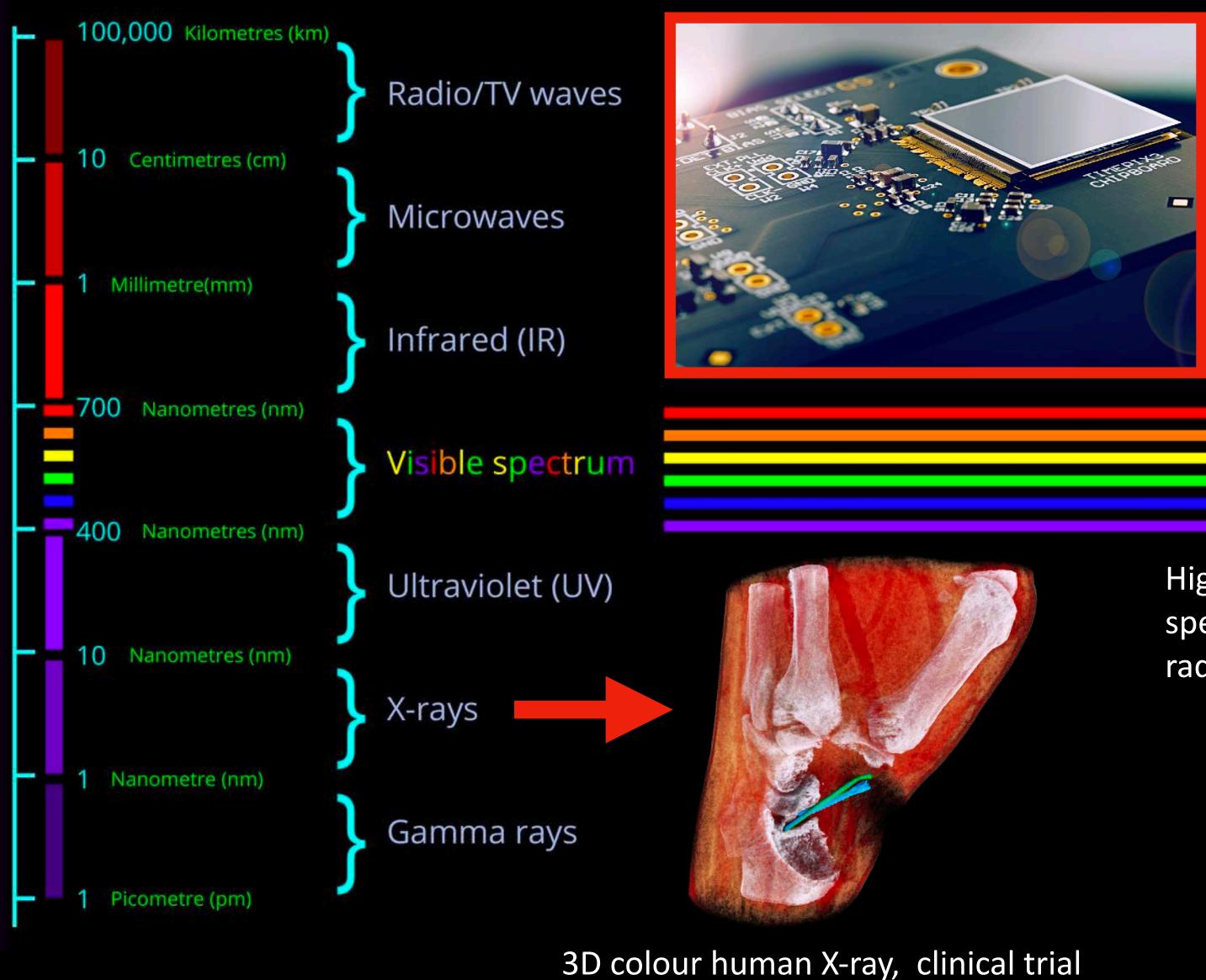


Dipole antenna system. Each antenna unfolds up to 55cm in length Standard format structure Solar panel Payload Electrical power system **UHF/VHF** radio Flight control computer





Blue skies research ... not with imaginations/ideas



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Medipix3; a CMOS pixel detector readout chip designed to be connected to a segmented semiconductor sensor.

High resolution spectroscopic radiography









Blue skies research ... not with imaginations/ideas



https://cms.cern/news/how-can-high-energy-physics-help-water-shortage

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The irrigation system will use fibre optic sensors designed to measure parameters such as temperature, humidity, concentration of pesticides, fertilisers and enzymes in the soil of cultivated fields. This is the same fibre optic sensors developed by CMS experiment in order to monitor the environment in the CMS tracking system.



Blue skies research ... not with imaginations/ideas

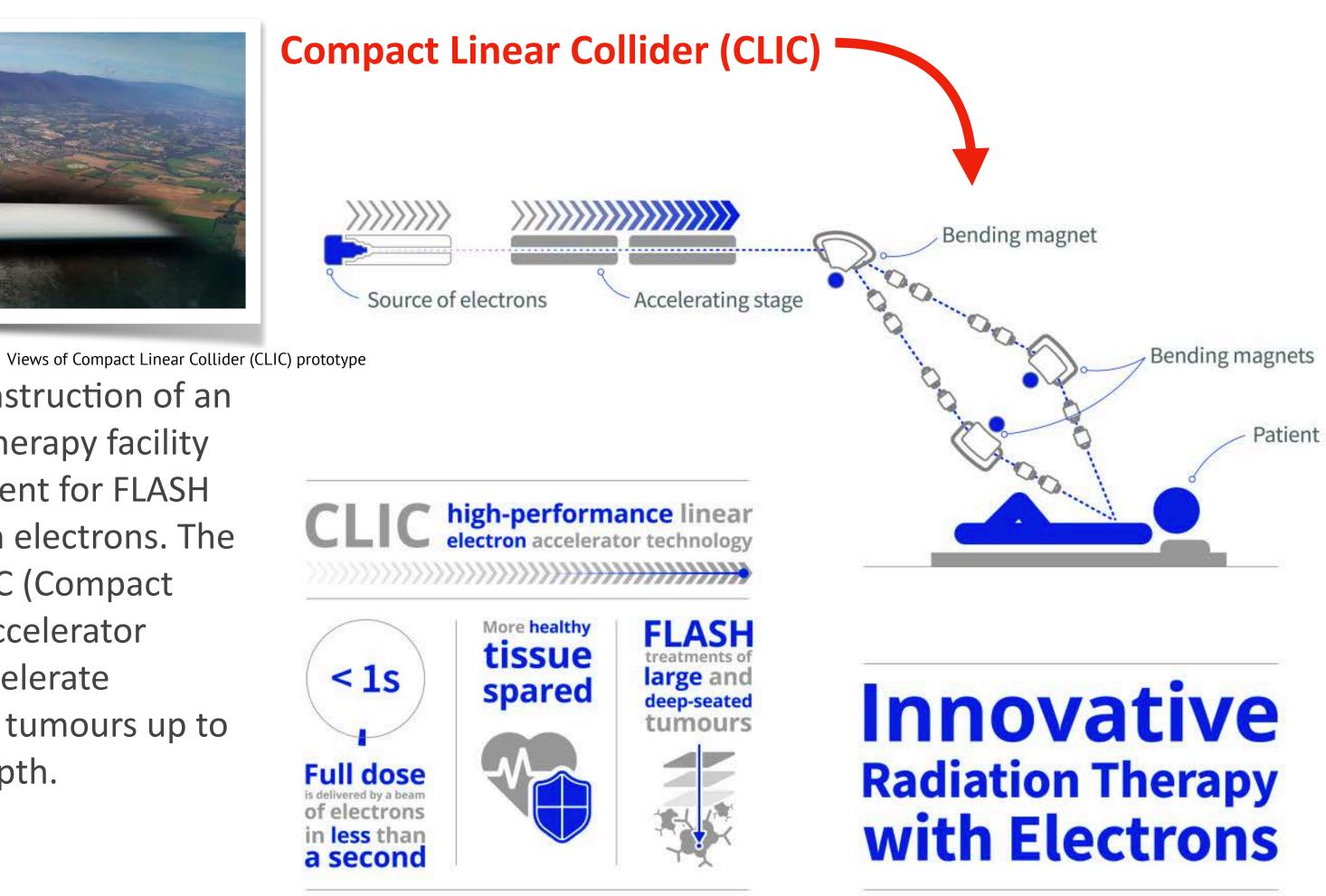
[Wikipedia] Blue skies research (also called blue sky science) is scientific research in domains where "real-world" applications are not immediately apparent. It has been defined as "research without a clear goal"[1] and "curiosity-driven science". It is sometimes used interchangeably with the term "basic research".



To design and construction of an innovative radiotherapy facility for cancer treatment for FLASH radiotherapy with electrons. The machine uses CLIC (Compact Linear Collider) accelerator technology to accelerate electrons to treat tumours up to 15 to 20 cm in depth.

CLIC prototype

https://cds.cern.ch/record/2728727





CERN International Teacher Programmes



โครงการความร่วมมือไทยเซิร์น

Published by Umaratchani Kaewbutta 😰 · 23h · 🕤

ด่วน ด่วน ด่วน หากคุณเป็นครูวิทยาศาสตร์ที่สอนระดับชั้นมัธยมศึกษาตอนปลาย ที่อายุไม่เกิน ่ 45 ปี มีความสามารถในการสื่อสารภาษาอังกฤษได้ดี และมีความสนใจในเรื่องของฟิสิกส์ อนุภาค และฟิสิกส์พลังงานสูง เราขอเชิญชวนท่านสมัครเข้าร่วมโครงการ CERN – International High School Teacher Programme 2023 และโครงการ CERN – International Teacher Weeks Programme 2023 ณ เซิร์น สมาพันธรัฐสวิส ครู ้วิทยาศาสตร์ที่สนใจสามารถสมัครเข้าร่วมโครงการได้ที่

โครงการ CERN HST2023 – International High School Teacher Programme 2023 (ระยะเวลา 2 สัปดาห์ ช่วงระหว่างวันที่ 2-15 กรกฎาคม 2566) https://teacher-programmes.web.cern.ch/international-high...

CERN ITW2023– International Teacher Weeks Programme 2023 (ระยะเวลา 2 สัปดาห์ ช่วงระหว่างวันที่ 6-19 สิงหาคม 2566) https://teacher-programmes.web.cern.ch/international...

สมัครได้ตั้งแต่บัดนี้จนถึงวันที่ 15 มกราคม 2566 ประสบการณ์อันแสนล้ำค่ารอคุณอยู่ที่สวิตเซ อร์แลนด์ค้า

...

- To support teachers' professional development in the field of particle physics. • To promote the teaching of particle physics in high schools.
- To facilitate the exchange of knowledge and experience among teachers of different nationalities.
- To stimulate activities related to the popularisation of physics within and beyond the classroom.
- To help CERN establish closer links with schools all around the world.

<u>https://teacher-programmes.web.cern.ch/international-teacher-programmes</u>

Until 15 January 2023!!!

- CERN offers two 2-week international teacher programmes: The International High School Teacher (HST) Programme and the International Teacher Weeks
- (ITW) Programme. Both programmes are delivered entirely in English and are open to in-service science teachers from around the world. The goals of CERN's international teacher programmes are:

