

CERN Neutrinos to Gran Sasso


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- “Neutrinos” and “Gran Sasso”
- CNGS: main components, layout, expected performance, cost, schedule
- Summary

What are Neutrinos (ν) ?



- ⇨ elementary particles
- ⇨ come in three flavors (pistache, chocolat, vanilla) 
- ⇨ electric charge: zero !
- ⇨ mass: very small zero?
- ⇨ interaction with matter: “very weak”

“ the elusive particle “

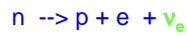
Leptons			
	electric charge		electric charge
e	-1	ν_e	0
μ	-1	ν_μ	0
τ	-1	ν_τ	0

Where are the Neutrinos ?



⇨ “all around us”

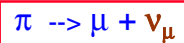
-> radioactive decay of atomic nuclei (e.g. in granite)



-> nuclear reactors

-> from the sun

-> accelerators... (high energy ν)



-> from reactions of cosmic rays in the atmosphere (ν_μ, ν_e)

....

⇨ in the universe: some hundreds per cm^3 (big bang “leftover”)

Neutrino mass ?



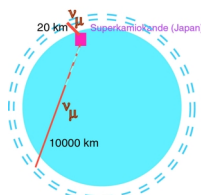
⇨ Standard model of particle physics: ν masses ZERO

⇨ “Direct” mass measurements -> upper limits
(in decay experiments measuring kinetic energy of “the partner”)

$m_{\nu_e} < 5 \text{ eV}$ $m_{\nu_\mu} < 170 \text{ keV}$ $m_{\nu_\tau} < 18 \text{ MeV}$

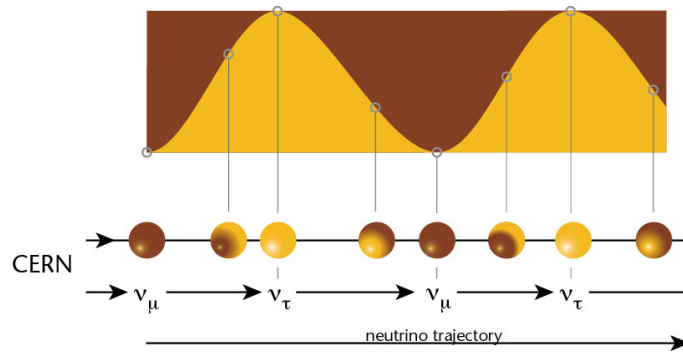
⇨ What’s the problem ? (DARK MATTER ...)

⇨ OBSERVATION 1 : SOLAR NEUTRINO “DEFICIT”
only about 50% of the ν_e expected are actually observed
 ν_e disappear “en route” from the sun to the earth ...?



⇨ OBSERVATION 2 : ATMOSPHERIC NEUTRINO ANOMALY
much less ν_μ “from below” observed w.r.t. expectations
 ν_μ disappear “en route” over 10’000 km ... ?

Neutrino Oscillation



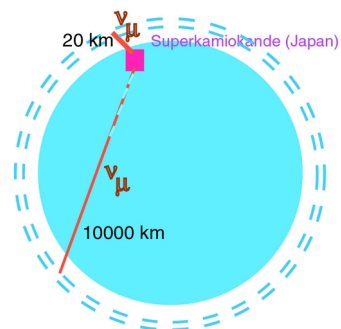
v's change flavor ! Is this possible? -- Yes, "if neutrinos have mass"!

very simplified picture:

$$\text{Probability}(v_\mu \rightarrow v_\tau) \approx \sin^2(1.27 \Delta m^2 L / E)$$

Δm^2 difference of mass between v 's (eV^2)
 L distance source / detector (km)
 E neutrino energy (GeV)

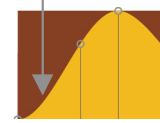
Example: $\Delta m^2 = 3 \times 10^{-3} eV^2$, $E = 10 \text{ GeV}$
 $\rightarrow L \approx 4000 \text{ km}$



4000 km -- Gran Sasso ?



Gran Sasso



In Dec. 1999, CERN council approved the CNGS project

- build an intense ν_μ beam at CERN-SPS
- search for ν_τ at Gran Sasso laboratory (730 km from CERN)

"long base-line" ν_μ -- ν_τ oscillation search

ST-Chamonix, 2 Feb 2001

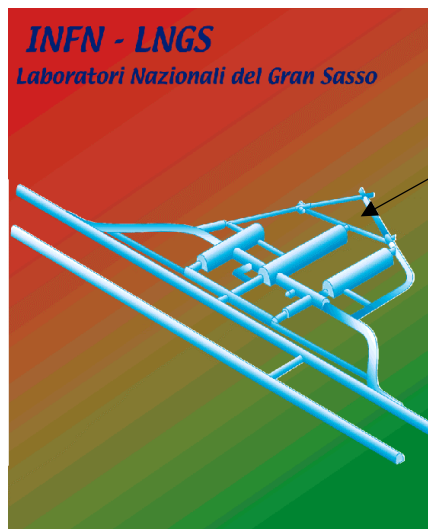
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The Gran Sasso Laboratory

730 km might seem to short -
but look at the details :
(Background low enough,
event rate still acceptable)
--> 730 km almost perfect

AND:

- existing laboratory with its infrastructure (since 1987)
- large halls directed to CERN
- rock overburden:
1500 m of shielding
- tradition in very successful neutrino physics experiments (solar ν 's)

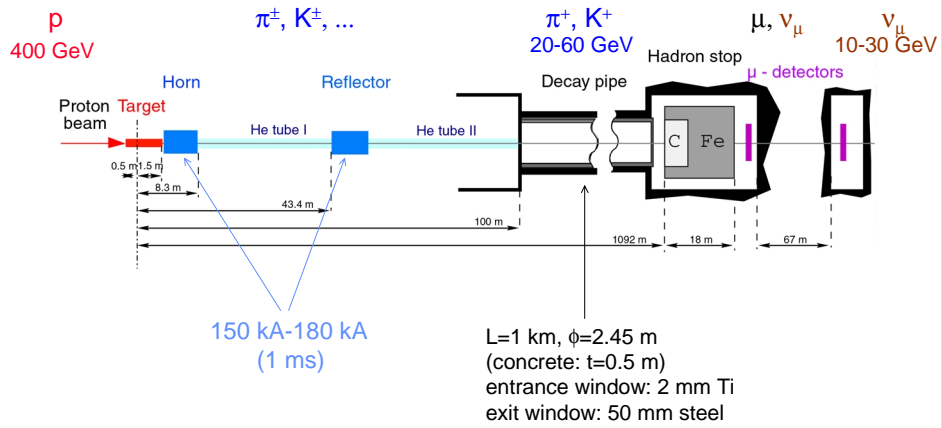


CERN

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CNGS Overview

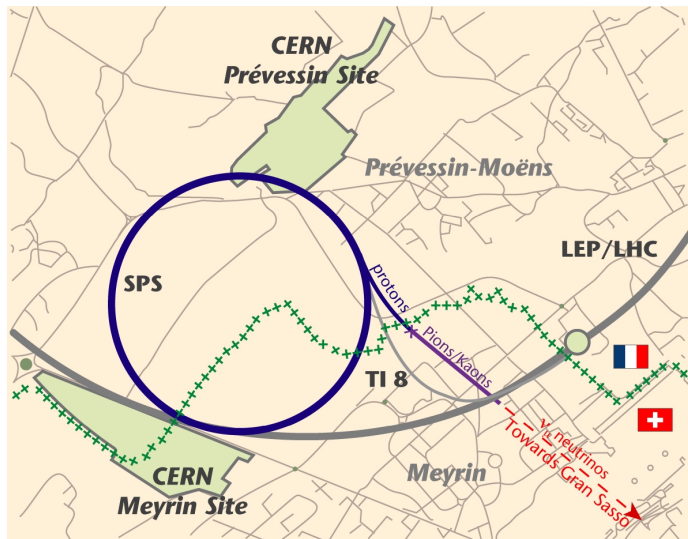


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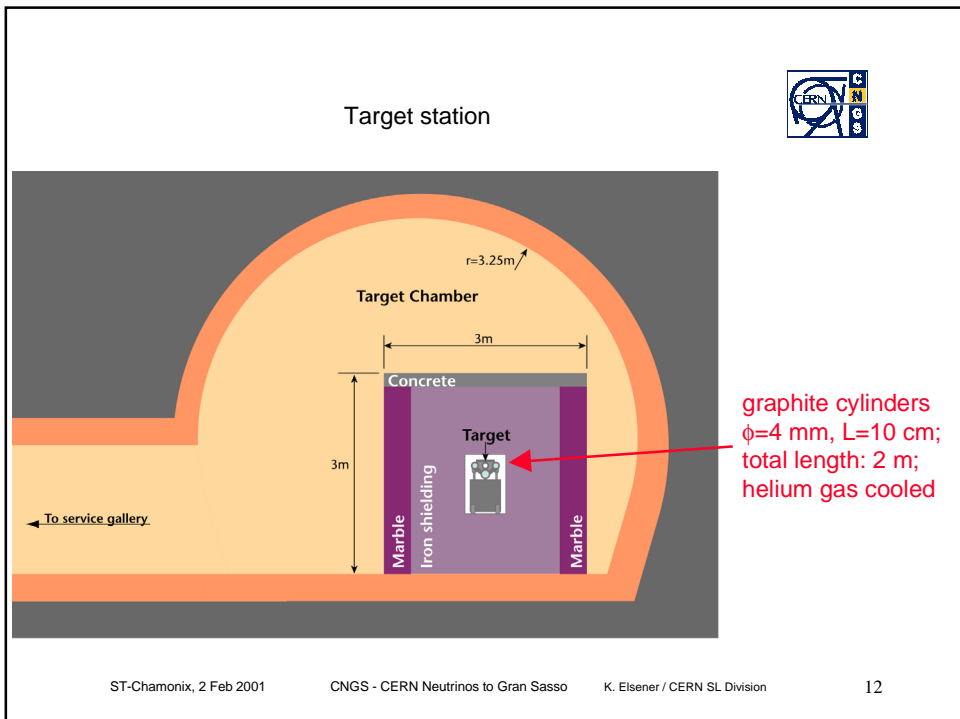
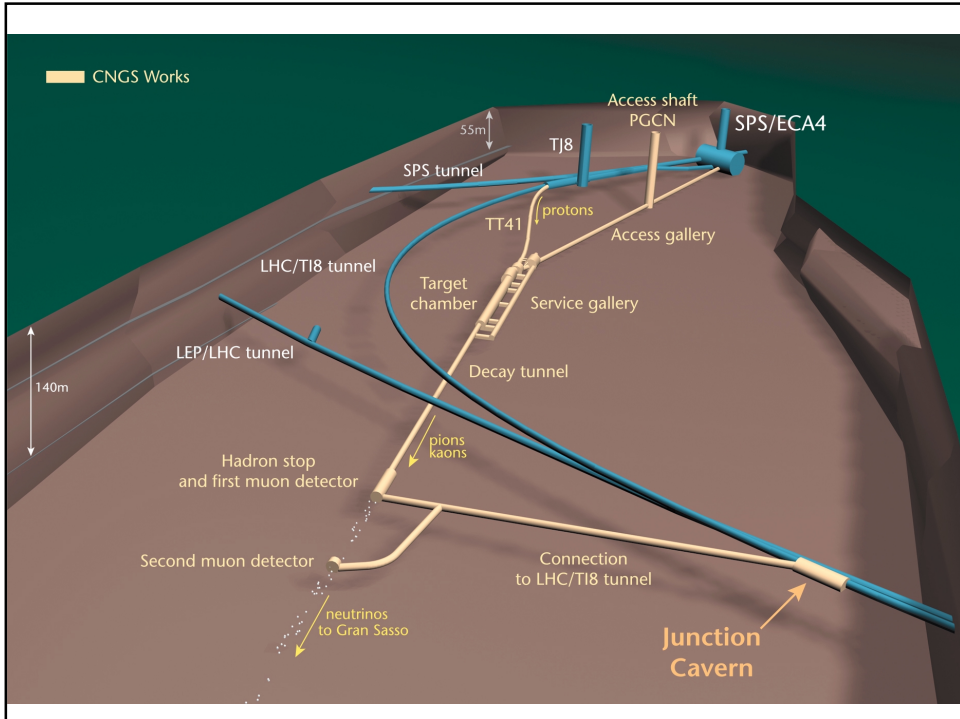


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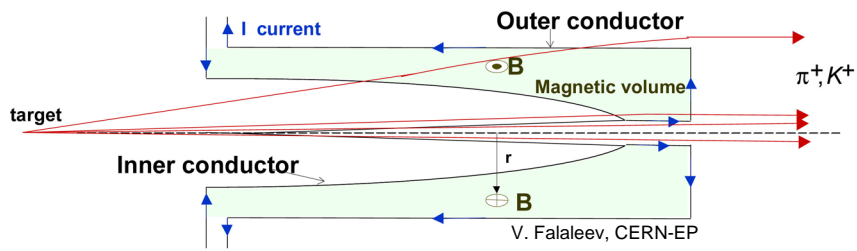
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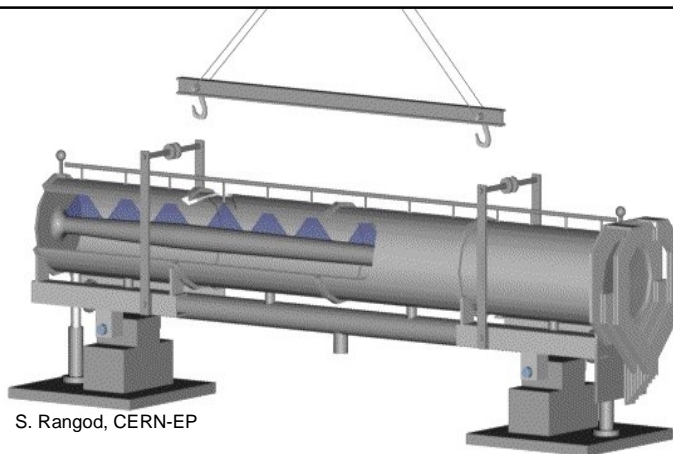


Horn focusing: principle (S. van der Meer)



Inner conductor - paraboloid of revolution
Outer conductor - cylinder

over/under-focused particles: second horn ("reflector")



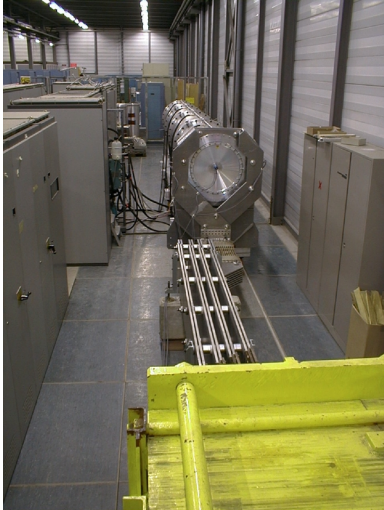
S. Rangod, CERN-EP



energy dissipated in the Horn / Reflector
by electrical current and beam : water cooling.

(overall length: 6.5 m, diameter of outer conductor: 0.7 m)

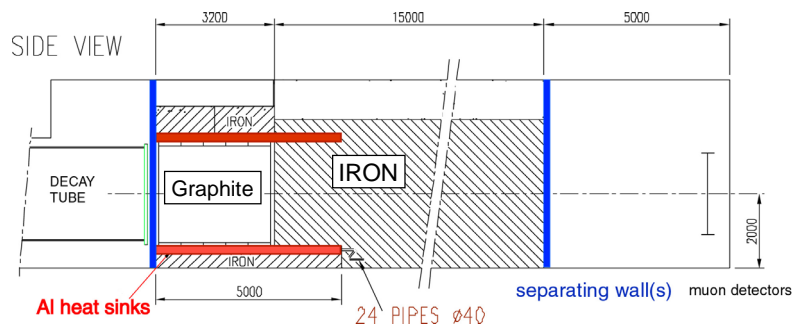
CNGS horn prototype tests in BA7/SPS



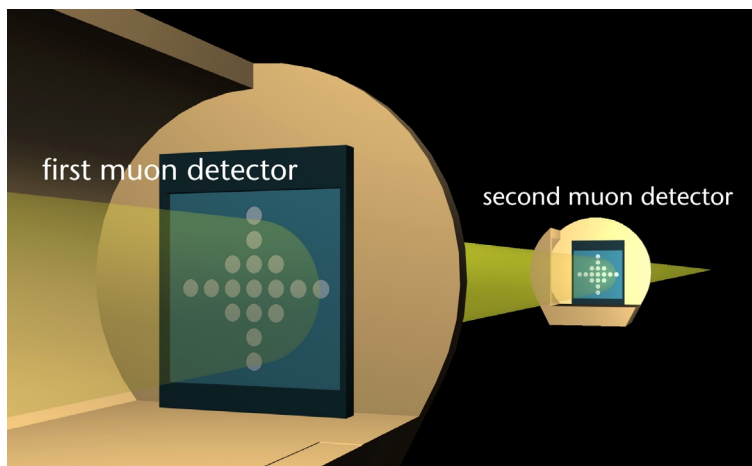
View inside the horn - "spider" support of inner conductor



Hadron stop (beam dump) area



Muon monitoring stations



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Experiments at Gran Sasso



Note:

- (1) after having travelled 730 km underground...
neutrinos are not easily "caught" in a detector
- (2) the beam at Gran Sasso has **diameter of about 2 km**

OPERA approved

Pb sheets and photographic emulsions;
1.8 kton target material (final config.)
bricks -> walls -> modules (incl. scintillator) -> S-modules
235'000 bricks in total - "hit" registered in electronic detectors
-> robot extracts **30 bricks per day for shipment / analysis**
computing power for analysis: "worldwide effort" !

ICANOE under approval

Liquid argon time projection chamber;
4 kton target material (?)
initial module **600 tons**, now testing
(this test is considered an important milestone)

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OPERA

Emulsion Pb

τ

μ

π

1mm

v_τ

Neutrino interaction in emulsion:

- ◆ Sandwich of emulsion film and lead sheets
- ◆ Mass > 1 kton, resolution 1 μm
- ◆ Proven technique: **CHORUS (CERN), DONUT (Fermilab)**

RUN 624 EVENT 576 Z-projection

EMULSION LAYER

ACETATE BASE

EMULSION LAYER

μ

π

τ

1mm

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ICARUS

- ◆ General purpose detector for CNGS + solar and atmospheric neutrinos, proton decay
- ◆ ultrapure liquid Argon equipped with wire planes: true 3D image of tracks produced e.g by a v_τ
- ◆ Technique proven with West Area neutrinos (v_μ event -->)

CNGS: Expected Performance



neutrino beam performance "detectable events"

No oscillation:

2450 ν_μ cc. events per kt per year

Assume $\sin^2(2\theta)=1$ and $\nu_\mu - \nu_\tau$ is the correct hypothesis:

$\Delta m^2 = 5 \times 10^{-3} \text{eV}^2$:	56 ν_τ	cc. events per kt per year
$\Delta m^2 = 3 \times 10^{-3} \text{eV}^2$:	21 ν_τ	
$\Delta m^2 = 1 \times 10^{-3} \text{eV}^2$:	2.3 ν_τ	

including exp. inefficiencies "expected events"

OPERA 1.8 kton x 5 years running:

$\Delta m^2 = 3.2 \times 10^{-3} \text{eV}^2$: 18.3 ν_τ 0.57 b.g.

ICANOE 20 kton x year running:

$\Delta m^2 = 3.5 \times 10^{-3} \text{eV}^2$: 37 ν_τ 4.4 b.g.

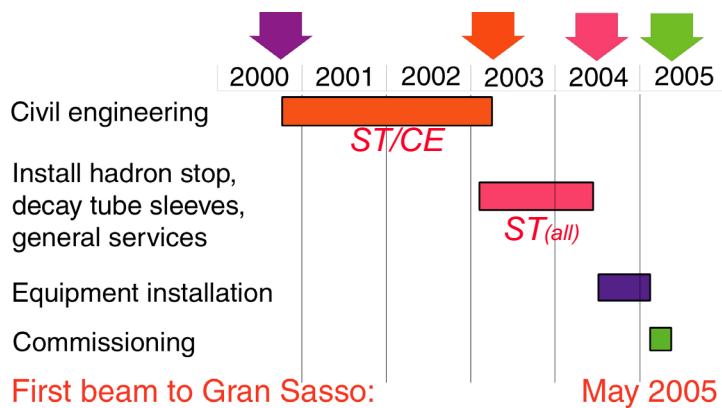
Cost of the CNGS facility*

*without detectors at Gran Sasso



Civil Engineering (including decay tube)	41.6 MCHF	Voluntary contributions announced: Belgium 1 MCHF France 1 MCHF Germany 5 MCHF Italy 47 MCHF Comp. S. Paolo (Italy) 1.6 MCHF Spain 4 MCHF Switzerland (Dec. 2000) 1.5 MCHF Total 61.1 MCHF
Equipment	19.6 MCHF	
Infrastructure	7.3 MCHF	
Contingency	2.5 MCHF	
Total	71 MCHF	
+ the value of recuperated equipment and investment for LHC: estimated 22 MCHF		

CNGS schedule / status

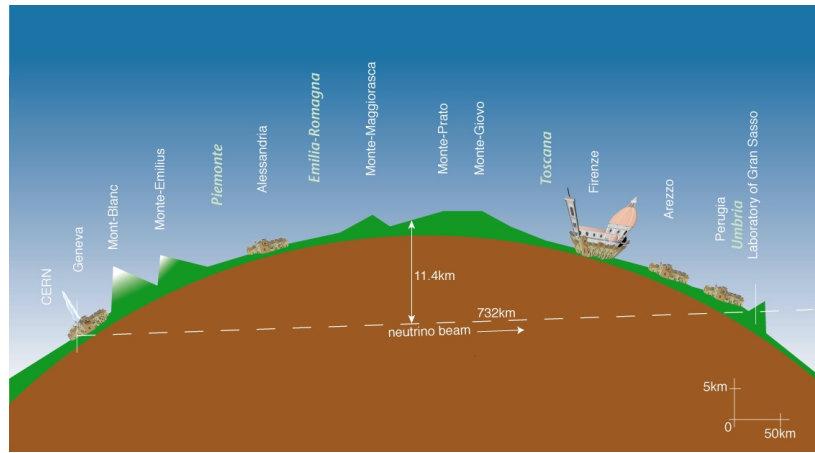


Summary

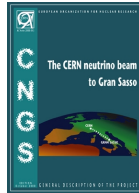


- Following the approval of CNGS in Dec. 1999, construction started in Sept. 2000.
- First CNGS beam for the ν_τ appearance search is expected in May 2005.

Thanks for your contribution to CNGS!



for more information:



CNGS general description

+ **NEW** CNGS www pages:
<http://proj-cngs.web.cern.ch/proj-cngs>